The 4th KMI International Symposium (KMI 2019) Quest for the Origin of Particles and the Universe

Looking at Particle - AntiParticle asymmetry with Heavy Flavors

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February 18, 2019

Kobayashi-Maskawa Institute for the Origin of Particles and the Universe

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Talk Outline

- Topical session focused on "Mysteries of the matterantimatter asymmetry in the Universe".
- Introductory overview for the 2nd KMI school.
- Introduction
- CP violation in the SM
 - Flavor changing quark interaction
- Approaching New Physics with heavy flavors
 - LHC and B experiments
- Summary

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 Ist generation
 Ind generation

 Image: Charmer of the perturbation of the pert

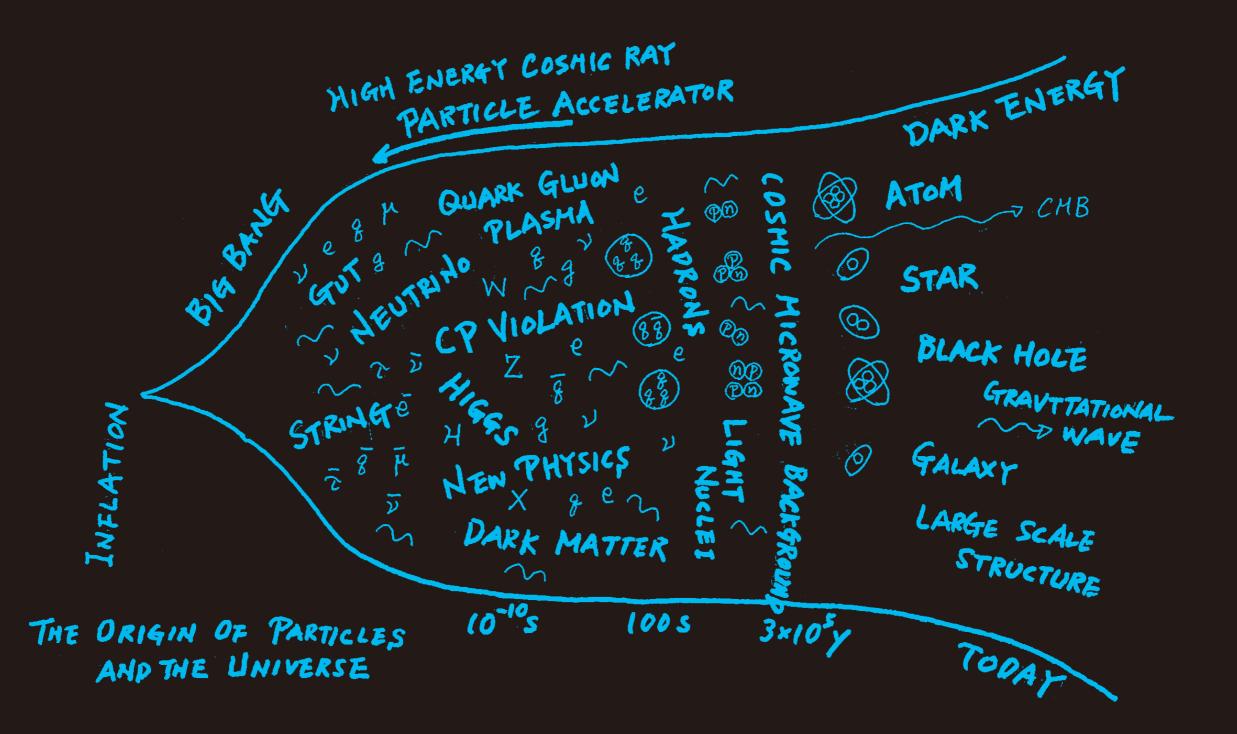
Strong CP (M.Ramsey-Musolf)

Neutrino (F. di. Lodovico)

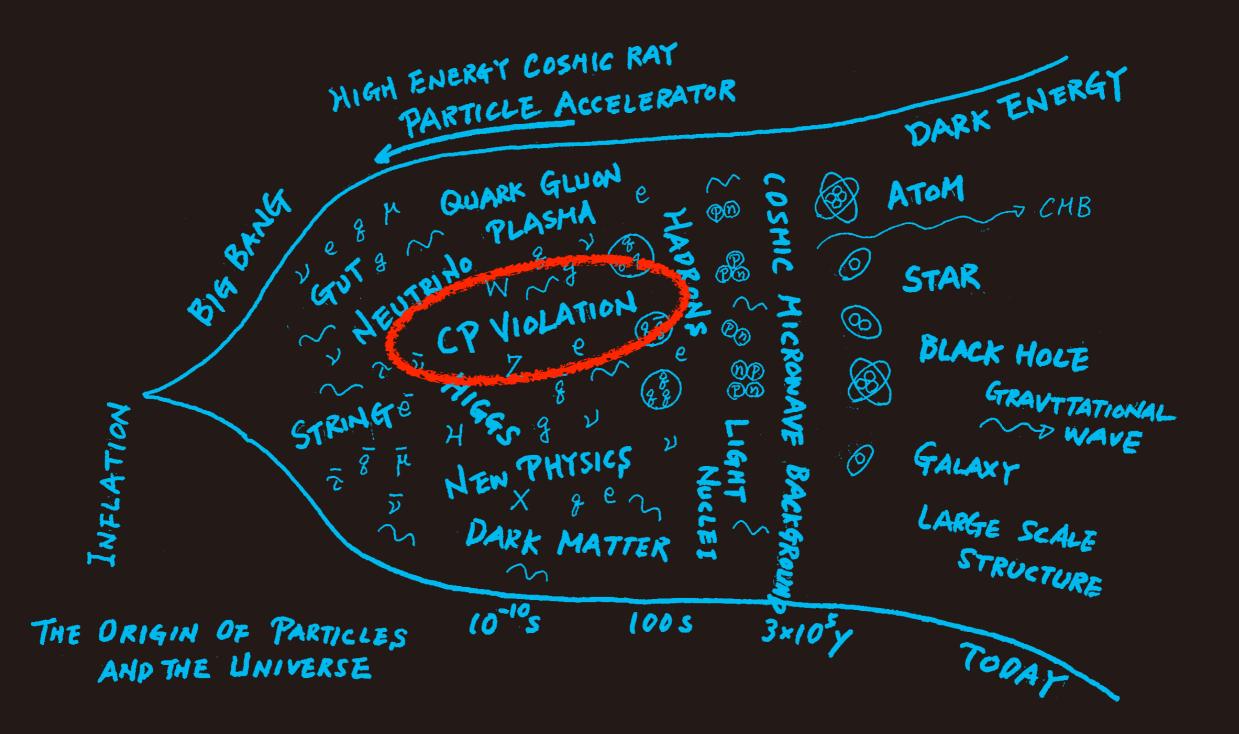
Charged leptons (K. Hayasaka)

Cosmic ray research (S. Haino)

History of the Universe

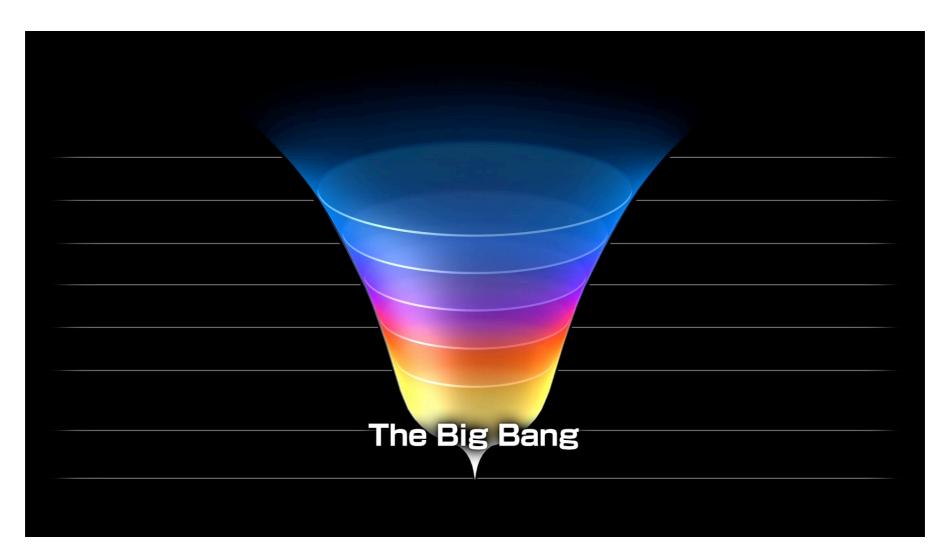


History of the Universe



Mystery of Anti-matter Disappearance

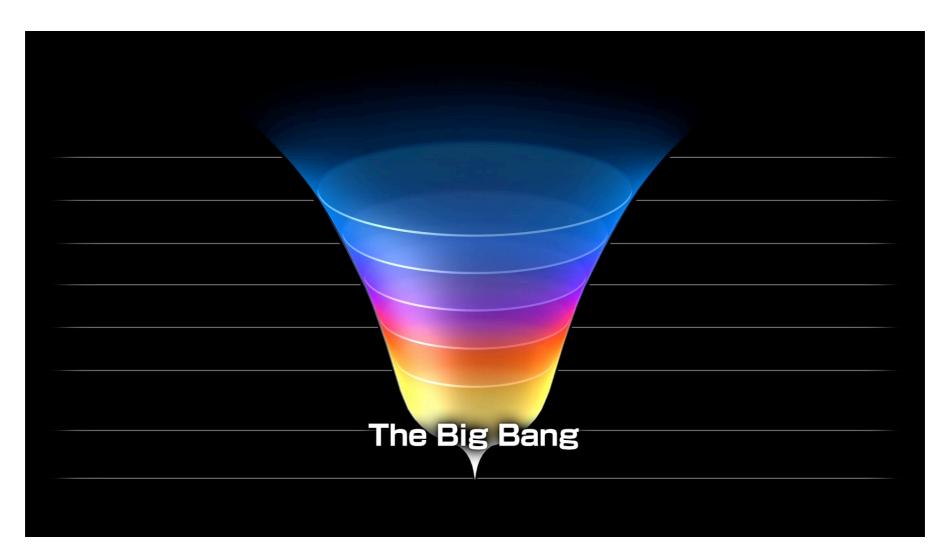
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Physicists believe that equal number of particles and anti-particles are produced from Big Bang. However, anti-particles disappeared somehow, and the present Universe is dominated by matters. Why ?

Mystery of Anti-matter Disappearance

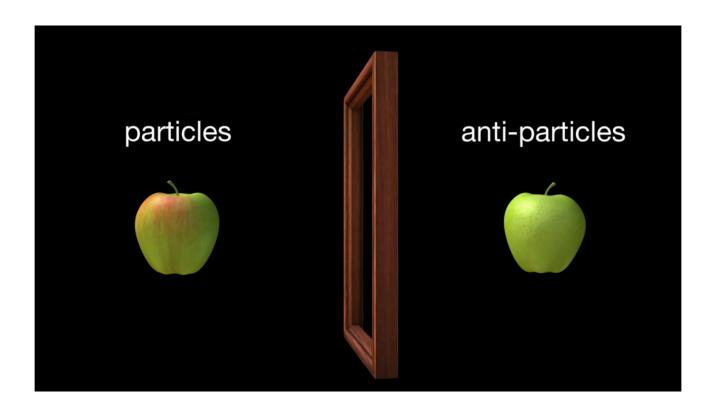
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Physicists believe that equal number of particles and anti-particles are produced from Big Bang. However, anti-particles disappeared somehow, and the present Universe is dominated by matters. Why ?

CPViolation

To make the matter dominated Universe, we need I. Baryon number violation II. **CP violation** III. Loss of thermal equilibrium [Sakharov's 3 conditions]



Discovery of CPV (1964)

• V. Fitch, J. Cronin et al.

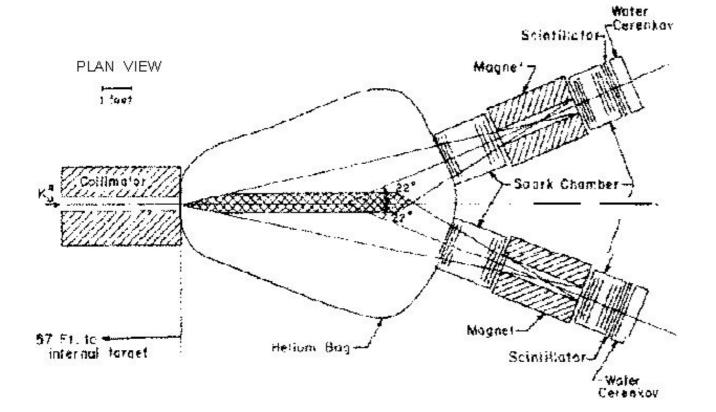


Fig. I. Plan view of the apparatus as located at the A. G. S.

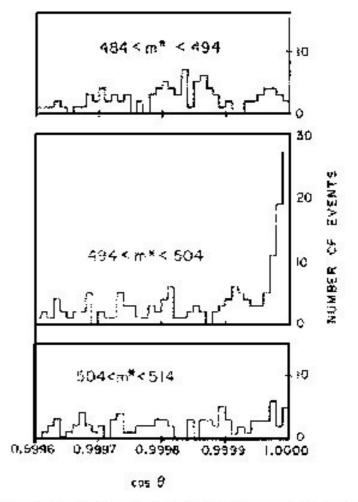
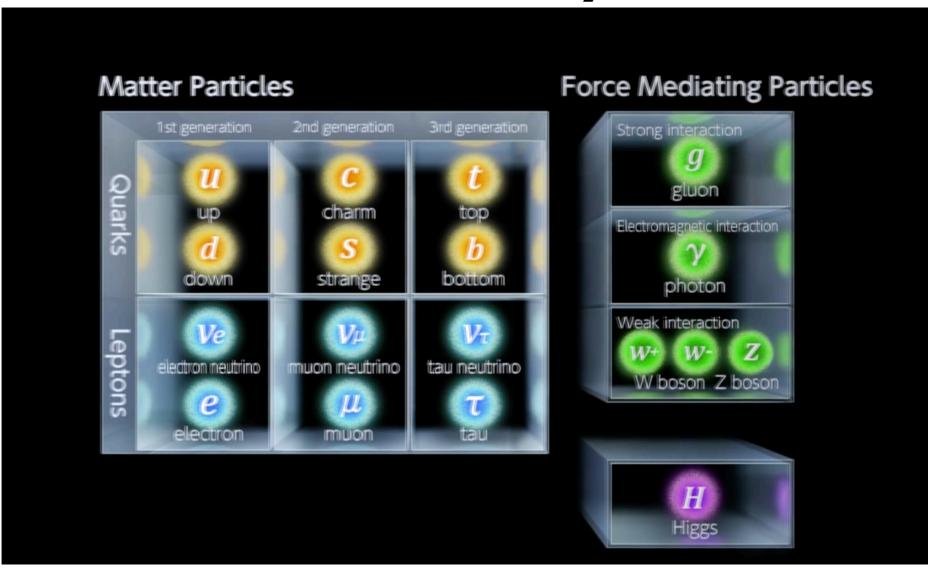


Fig. 3. Angular distribution of the events after measurement by a precise machine in three relevant mass regions.

$$\begin{split} |K_{S}\rangle = &(1+\left|\epsilon\right|^{2})^{1/2} \left[\left|K_{1}\rangle+\epsilon\right|K_{2}\rangle\right] \\ |K_{L}\rangle = &(1+\left|\epsilon\right|^{2})^{1/2} \left[\epsilon\left|K_{1}\rangle+\left|K_{2}\rangle\right] \end{split}$$

The Standard Model of Particle Physics



Elementary particles make up the Universe Matter particles (quarks and leptons) Force mediating particles (bosons) and Higgs !

2008 Nobel Prize in Physics



Photo: University of Chicago

Yoichiro Nambu Prize share: 1/2



© The Nobel Foundation Photo: U. Montan

Makoto Kobayashi Prize share: 1/4



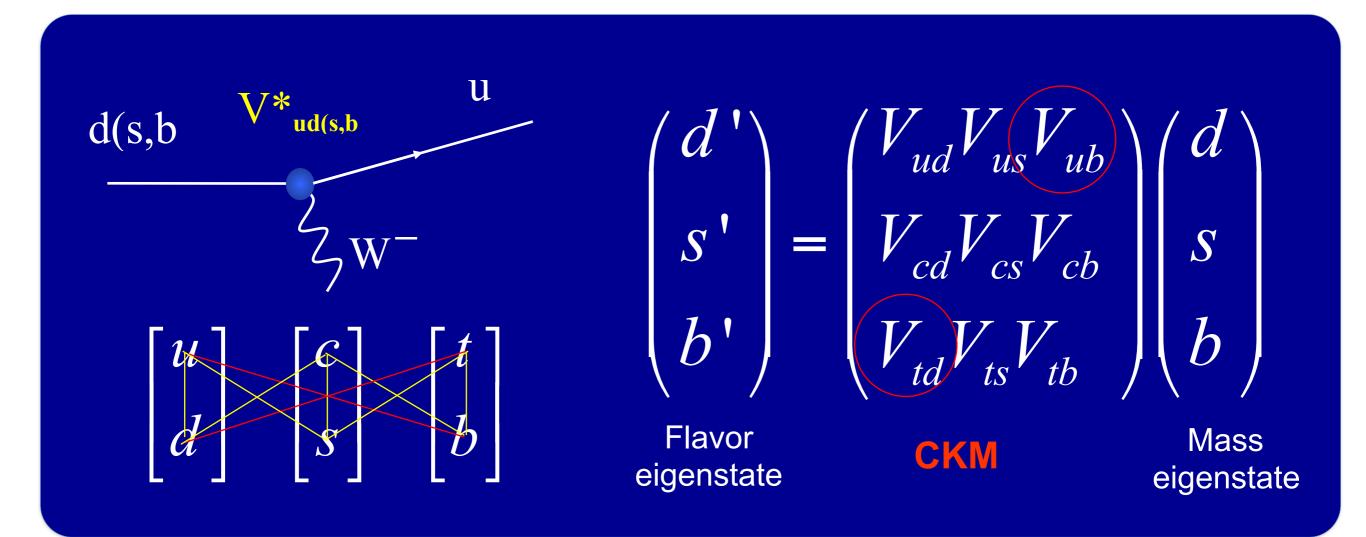
© The Nobel Foundation Photo: U. Montan

Toshihide Maskawa Prize share: 1/4

The Nobel Prize in Physics 2008 was divided, one half awarded to Yoichiro Nambu "for the discovery of the mechanism of spontaneous broken symmetry in subatomic physics", the other half jointly to Makoto Kobayashi and Toshihide Maskawa "for the discovery of the origin of the broken symmetry which predicts the existence of at least three families of quarks in nature."

Kobayashi-Maskawa Theory

- A quark change its flavor by emitting virtual W.
- 3X3 unitarity triangle has 3 rotational angles and one complex phase.
 CPViolation



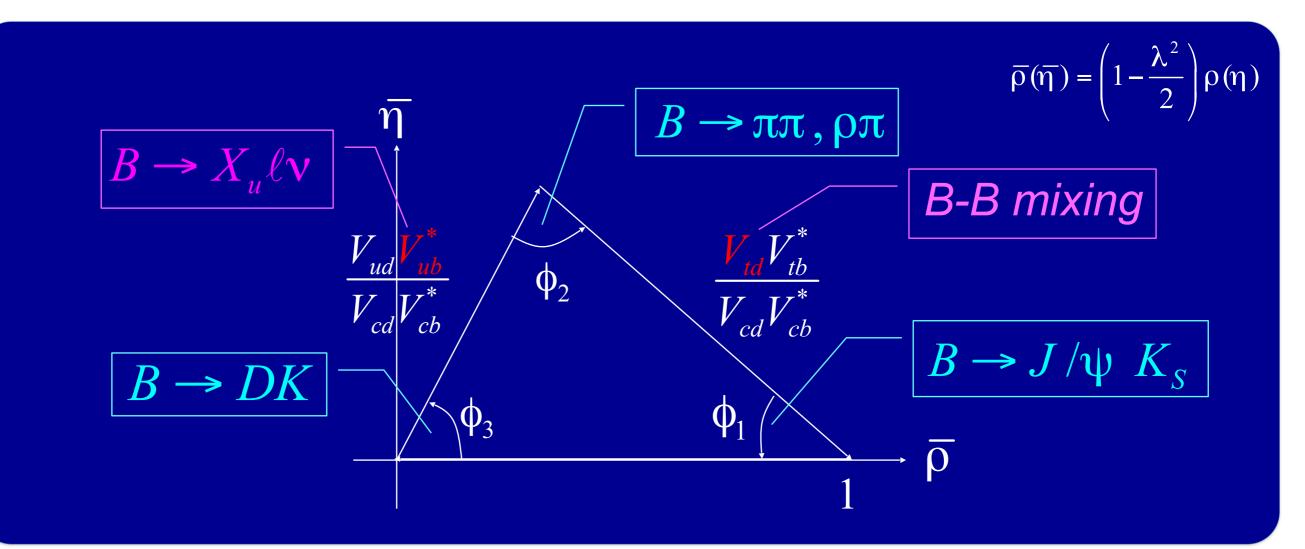
Unitarity Triangle

unitarity condition

$$V_{ud}V_{us}^{*} + V_{cd}V_{cs}^{*} + V_{td}V_{ts}^{*} = 0$$

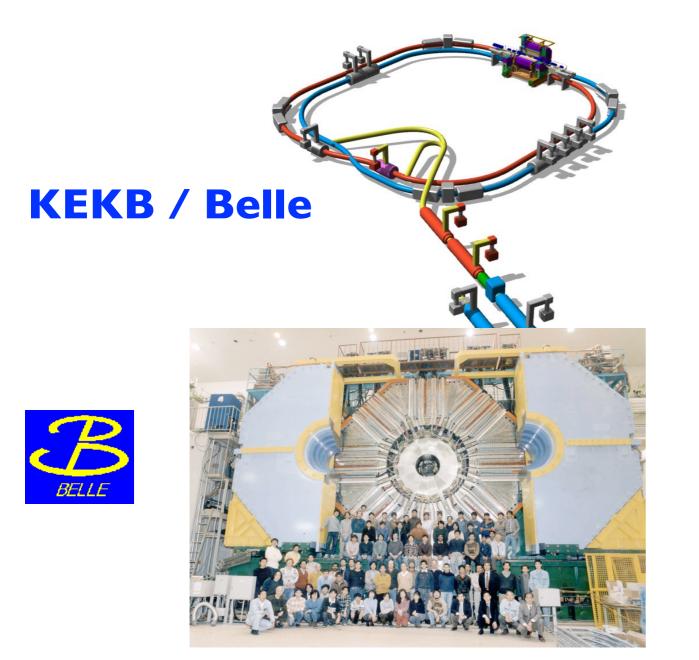
$$V_{ud}V_{ub}^{*} + V_{cd}V_{cb}^{*} + V_{td}V_{tb}^{*} = 0$$

$$V_{us}V_{ub}^{*} + V_{cs}V_{cb}^{*} + V_{ts}V_{tb}^{*} = 0$$



B-Factory Experiments

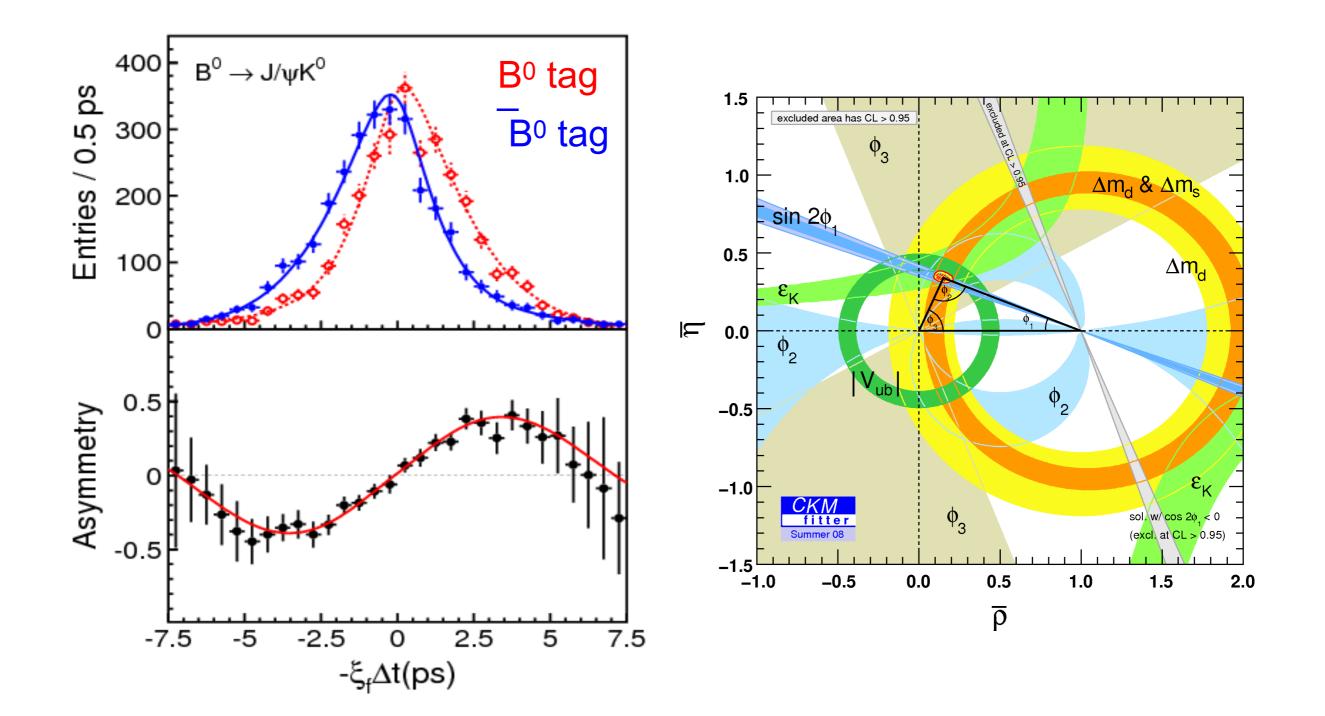
In 2001, the Belle experiment at KEK and the BaBar experiment at SLAC successfully measured CP violation in B meson decays as predicted by the Kobayashi-Maskawa theory.



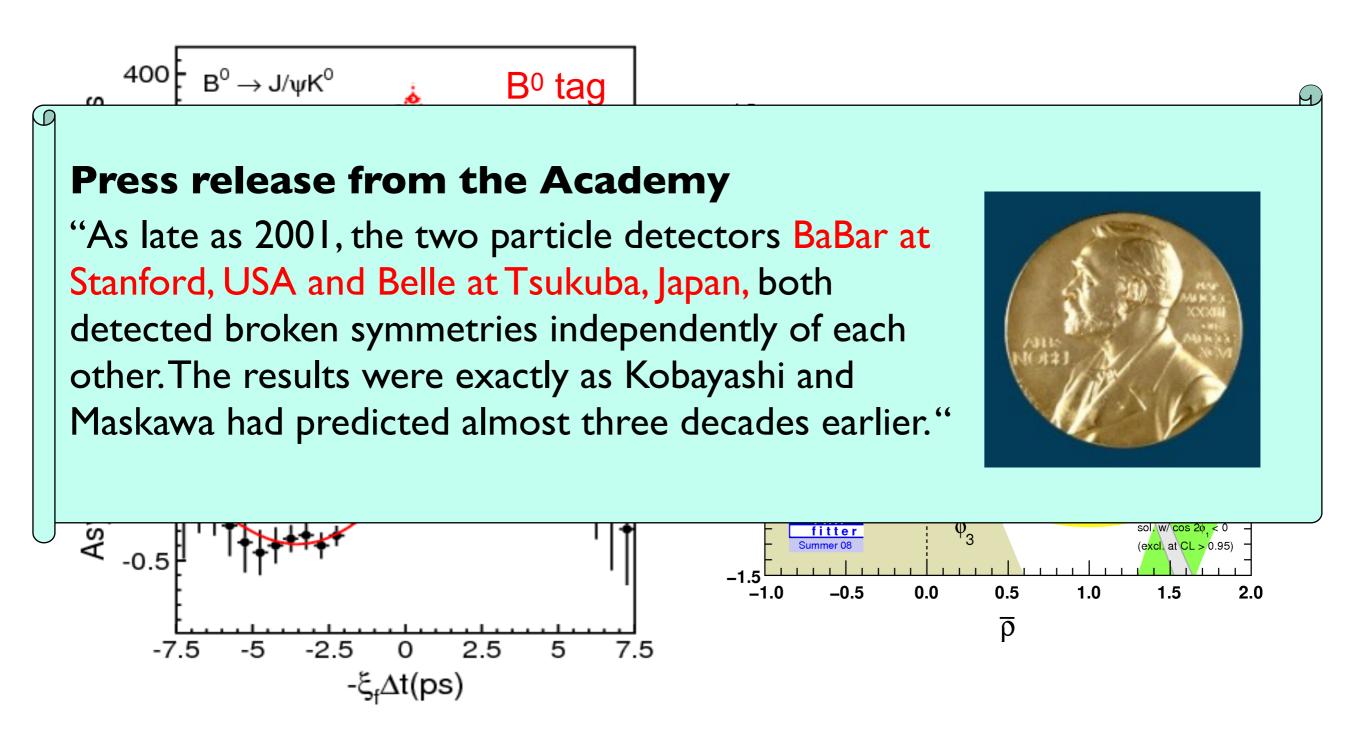




Confirmation of Kobayashi-Maskawa

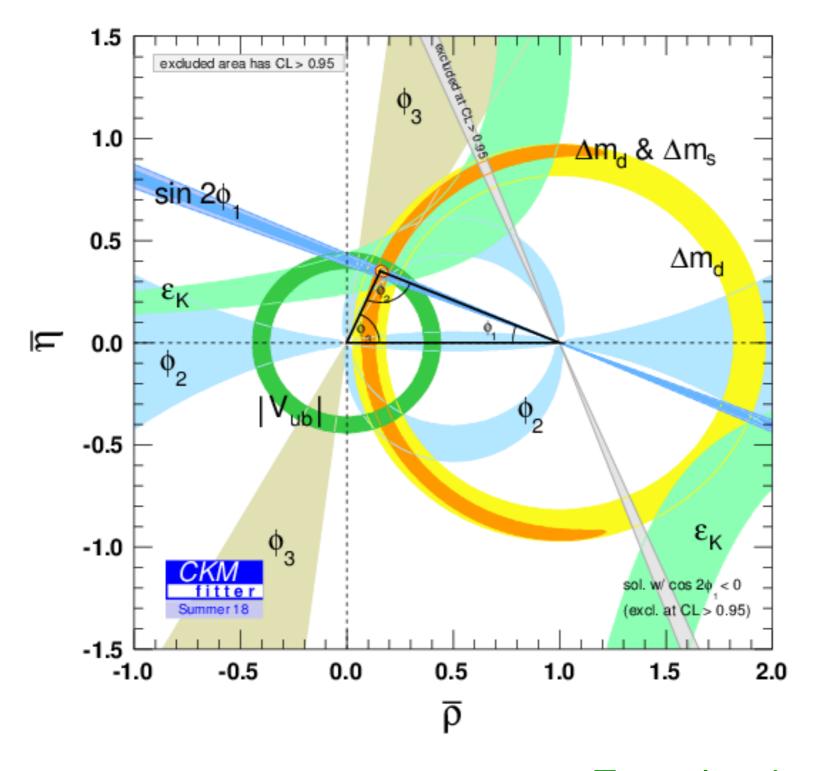


Confirmation of Kobayashi-Maskawa



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Latest Results (Summer 2018)



Tutorial at the 2nd KMI School

Large Hadron Collider

フランス

周長27 Km

7TeVの陽子と7TeVの陽子を衝突 世界最高エネルギー14TeVの世界

Large Hadron Collider



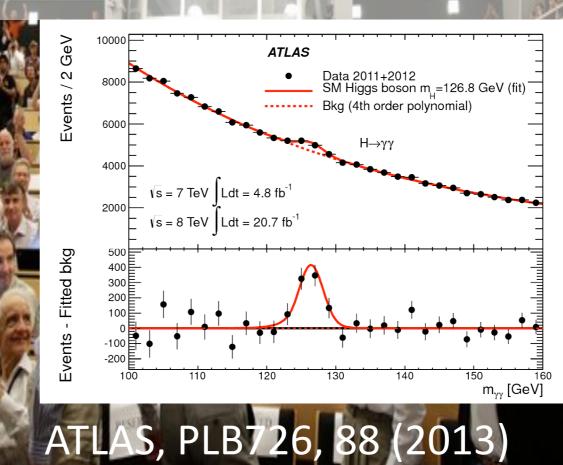


フランス

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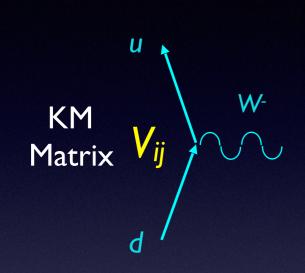


周長27 Km

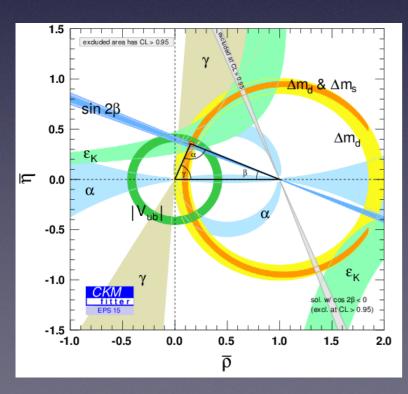
ヒッグス粒子発見!

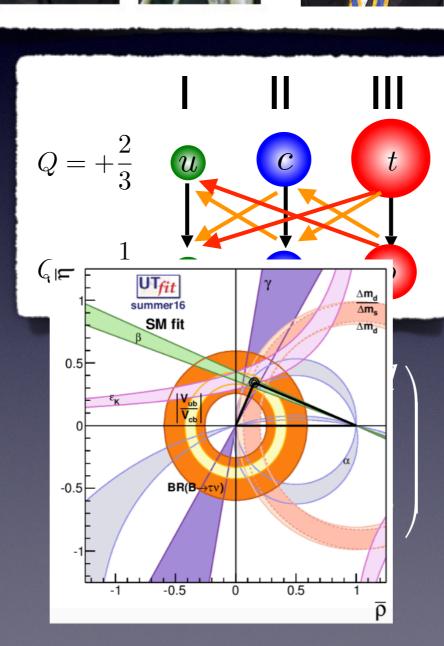
TeVの陽子と7TeVの陽子を衝突 世界最高エネルギー14TeVの世界

CPViolation and Mass

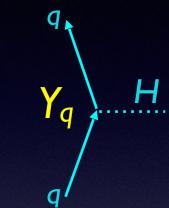


Test of CPV origin

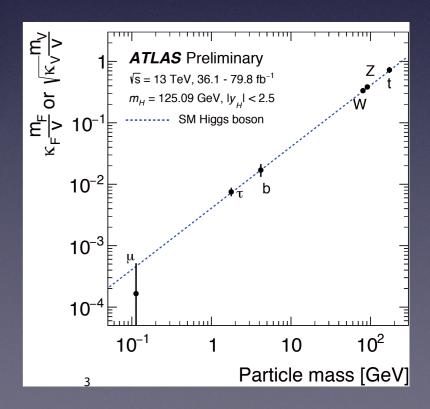




Yukawa coupling



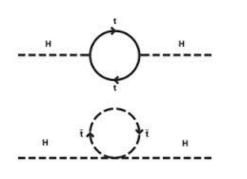
Test of mass origin

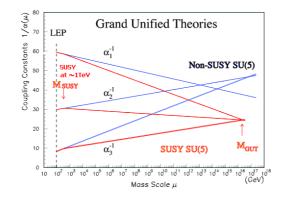


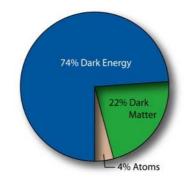
Why do we need go further ?

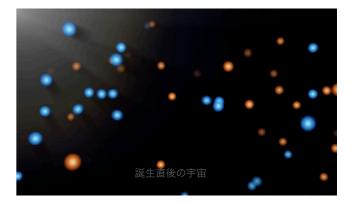
SM explains almost every phenomena so far, but cannot explain

- The Higgs mass (fine tuning problem)
- Grand unification
- Dark matter
- Baryon asymmetry in the Universe
- Origin of the 3 generations







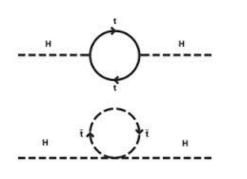


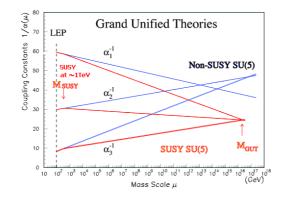


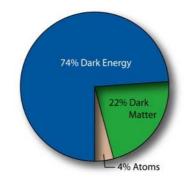
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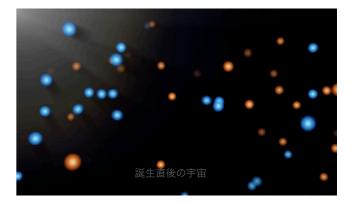
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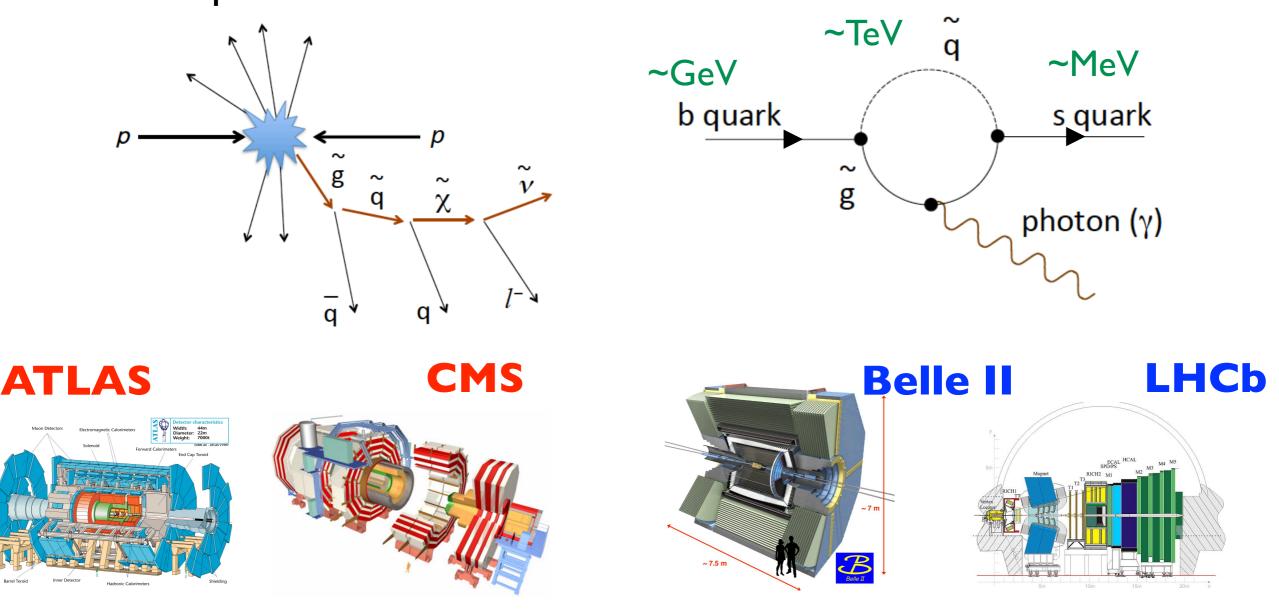




Two Ways to Find New Physics

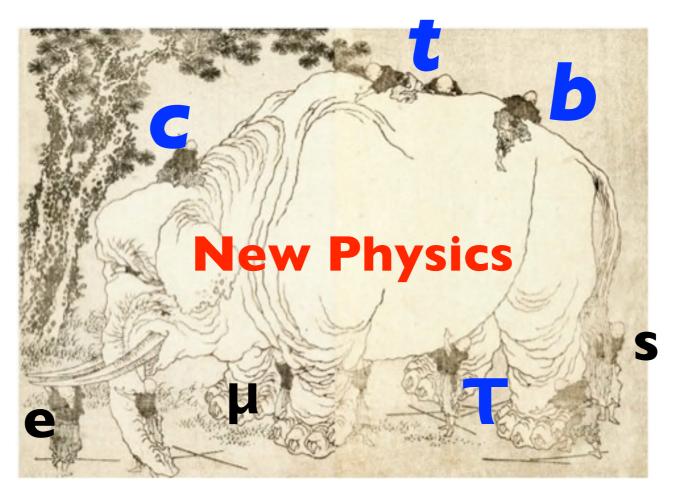
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- Energy Frontier : produces and detects a new particle directly in collisions of extremely high energy beams.
- Luminosity Frontier : measures reactions of known particles very precisely, and finds deviations from the Standard Model predictions.



Importance of Heavy Flavors

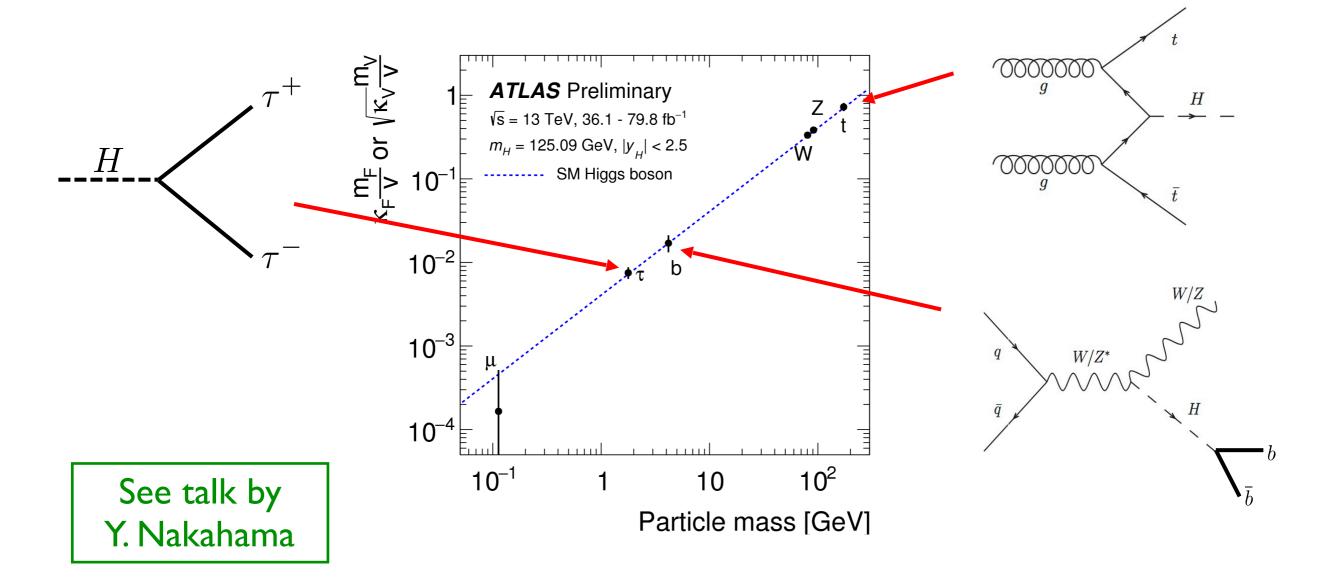
- New Physics is unknown.
- We need a variety of approaches to find and know it.
- Heavy flavor particles (t, b, c, T) are good probes
 - Sensitive to New Physics



"群盲象を撫でる"

Yukawa couplings to t, b, T

- Higgs has been discovered, and its couplings to fermions are being measured.
- Couplings to t, b, τ are just on the stage, and we need more precise measurements to test the SM and also to find NP.



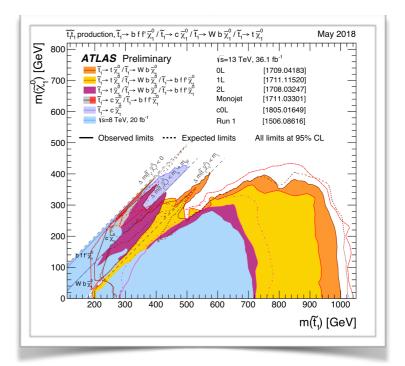
Direct NP Searches

- LHC Run2 (6.5TeV+6.5TeV) completed.
- No NP so far from partial data analyses.
- Extensive analyses for accumulated data (~150fb-1) are underway.

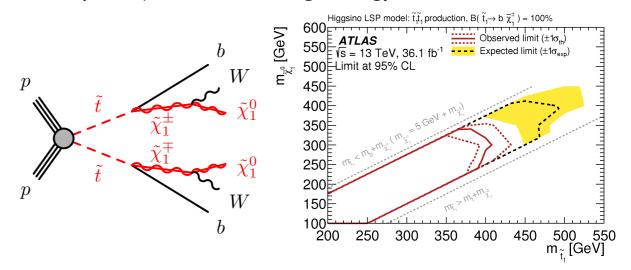
Summer 2018 (mostly 36.1fb⁻¹)

	ATLAS SUSY Searches* - 95% CL Lower Limits July 2018 Model e, μ, τ, γ Jets E_{τ}^{miss} $\int \mathcal{L} dt (fh^{-1})$ Mass limit $\sqrt{s} = 7, 8 \text{ TeV}$ $\sqrt{s} = 13 \text{ TeV}$										ATLAS Preliminar $\sqrt{s} = 7, 8, 13$ TeV
	Model	<i>e</i> , μ, τ, γ	Jets	$E_{\rm T}^{\rm mass}$	∫£ dt[ft	-1]	Mass limit		$\sqrt{s} = 7, 8$	TeV $\sqrt{s} = 13$ TeV	Reference
S	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_{1}^{0}$	0 mono-jet	2-6 jets 1-3 jets	Yes Yes	36.1 36.1	<i>q</i> [2x, 8x Degen.] <i>q</i> [1x, 8x Degen.]	0.43	0.9	1.55	m($\tilde{\chi}_1^0$)<100 GeV m(\tilde{q})-m($\tilde{\chi}_1^0$)=5 GeV	1712.02332 1711.03301
Inclusive Searches	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q \tilde{q} \tilde{\chi}_{1}^{0}$	0	2-6 jets	Yes	36.1	ğ ğ		Forbidden	2.0 0.95-1.6	m($\tilde{\chi}_{1}^{0}$)<200 GeV m($\tilde{\chi}_{1}^{0}$)=900 GeV	1712.02332 1712.02332
ve Se	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell \ell)\tilde{\chi}_{1}^{0}$	3 e, μ ee, μμ	4 jets 2 jets	- Yes	36.1 36.1	ê ê ê			1.85	m(ℓ ₁ ⁰)<800 GeV m(ℓ ₁ ⁰)=50 GeV	1706.03731 1805.11381
Iclusi	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_{1}^{0}$	о З е, µ	7-11 jets 4 jets	Yes	36.1 36.1	êg iêg		0.98	1.8	$m(\tilde{\chi}_1^0) < 400 \text{ GeV}$ $m(\tilde{g})-m(\tilde{\chi}_1^0)=200 \text{ GeV}$	1708.02794 1706.03731
-	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t t \tilde{\chi}_1^0$	0-1 e,μ 3 e,μ	3 b 4 jets	Yes -	36.1 36.1	ğ ğ			2.0	m(\tilde{t}_1^0)<200 GeV m(\tilde{g})-m(\tilde{t}_1^0)=300 GeV	1711.01901 1706.03731
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1{\rightarrow}b\tilde{\chi}_1^0/t\tilde{\chi}_1^\pm$		Multiple Multiple Multiple		36.1 36.1 36.1	$egin{array}{ccc} egin{array}{ccc} eta_1 & & Forbid \ eta_1 & & eta_1 \ eta_1 & & eta_1 \ eta_1 & & eta_1 \end{array}$	den Forbidden Forbidden	0.9 0.58-0.82 0.7	m(ž	$m(\tilde{\chi}_{1}^{0})=300 \text{ GeV}, BR(h\tilde{\chi}_{1}^{0})=1$ $m(\tilde{\chi}_{1}^{0})=300 \text{ GeV}, BR(h\tilde{\chi}_{1}^{0})=BR(r\tilde{\chi}_{1}^{0})=0.5$ $)=200 \text{ GeV}, m(\tilde{\chi}_{1}^{0})=300 \text{ GeV}, BR(r\tilde{\chi}_{1}^{0})=1$	1708.09266, 1711.03301 1708.09266 1706.03731
ks on	$\tilde{b}_1\tilde{b}_1,\tilde{\imath}_1\tilde{\imath}_1,M_2=2\times M_1$		Multiple Multiple		36.1 36.1	Ĩ₁ Ĩ₁ Forbidden		0.7		m(\tilde{k}_{1}^{0})=60 GeV m(\tilde{k}_{1}^{0})=200 GeV	1709.04183, 1711.11520, 1708.03247 1709.04183, 1711.11520, 1708.03247
3 ^{///} gen. squarks direct production	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow W b \tilde{\chi}_1^0 \text{ or } t \tilde{\chi}_1^0$ $\tilde{t}_1 \tilde{t}_1, \tilde{H} LSP$	0-2 e, µ (0-2 jets/1-2 Multiple Multiple	b Yes	36.1 36.1 36.1	<i>ī</i> 1 <i>ī</i> 1	dan	1.0 0.4-0.9 0.6-0.8	m(ž	$m(\tilde{t}_{1}^{0})=1 \text{ GeV}$ $m(\tilde{t}_{1}^{0})=1 \text{ GeV}, m(\tilde{t}_{1}^{+})\cdot m(\tilde{t}_{1}^{0})=5 \text{ GeV}, \tilde{t}_{1} \approx \tilde{t}_{L}$ $m(\tilde{t}_{1}^{0})=300 \text{ GeV}, m(\tilde{t}_{1}^{+})\cdot m(\tilde{t}_{1}^{0})=5 \text{ GeV}, \tilde{t}_{1} \approx \tilde{t}_{L}$	1506.08616, 1709.04183, 1711.11520 1709.04183, 1711.11520 1709.04183, 1711.11520
3" ge direct	$\tilde{t}_1 \tilde{t}_1$, Well-Tempered LSP $\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow c \tilde{\chi}_1^0 / \tilde{c} \tilde{c}, \tilde{c} \rightarrow c \tilde{\chi}_1^0$	0	Multiple 2c	Yes	36.1 36.1	\tilde{t}_1 Forbid \tilde{t}_1 \tilde{t}_1		0.48-0.84 0.85		$\tilde{\ell}_{1}^{0}$)=150 GeV, m($\tilde{\chi}_{1}^{\pm}$)-m($\tilde{\chi}_{1}^{0}$)=5 GeV, $\tilde{r}_{1} \approx \tilde{r}_{L}$ m($\tilde{\chi}_{1}^{0}$)=0 GeV	1709.04183, 1711.11520 1805.01649
		0	mono-jet	Yes	36.1		0.46 0.43			$m(\tilde{t}_1, \tilde{c})-m(\tilde{t}_1^0)=50 \text{ GeV}$ $m(\tilde{t}_1, \tilde{c})-m(\tilde{t}_1^0)=5 \text{ GeV}$	1805.01649 1711.03301
	$\tilde{t}_2 \tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + h$	1-2 e, µ	4 <i>b</i>	Yes	36.1	ĩ ₂		0.32-0.88		$m(\tilde{k}_{1}^{0})=0$ GeV, $m(\tilde{r}_{1})-m(\tilde{k}_{1}^{0})=180$ GeV	1706.03986
	$\tilde{\chi}_1^{\pm} \tilde{\chi}_2^0$ via WZ	2-3 e,μ ee,μμ	≥ 1	Yes Yes	36.1 36.1	$ \tilde{\chi}_{1}^{\pm}/\tilde{\chi}_{2}^{0} $ $ \tilde{\chi}_{1}^{\pm}/\tilde{\chi}_{2}^{0} $ 0.17		0.6		$m(\tilde{\chi}_{1}^{0})=0$ $m(\tilde{\chi}_{1}^{0})-m(\tilde{\chi}_{1}^{0})=10 \text{ GeV}$	1403.5294, 1806.02293 1712.08119
EW direct	$ \begin{array}{l} \tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \; \text{via} \; Wh \\ \tilde{\chi}_{1}^{\pm}\tilde{\chi}_{1}^{\mp} / \tilde{\chi}_{2}^{0}, \tilde{\chi}_{1}^{+} {\rightarrow} \tilde{\tau} \nu(\tau \tilde{\nu}), \tilde{\chi}_{2}^{0} {\rightarrow} \tilde{\tau} \tau(\nu \tilde{\nu}) \end{array} $	<i>ℓℓ/ℓγγ/ℓbb</i> 2 τ	-	Yes Yes	20.3 36.1	$\begin{array}{ccc} \tilde{\chi}_{1}^{\pm}/\tilde{\chi}_{2}^{0} & 0.26 \\ \\ \tilde{\chi}_{1}^{\pm}/\tilde{\chi}_{2}^{0} & \\ \tilde{\chi}_{1}^{\pm}/\tilde{\chi}_{2}^{0} & 0.22 \end{array}$		0.76	m(ž [±] ₁)-m	$m(\tilde{\chi}_{1}^{0})=0$ $m(\tilde{\chi}_{1}^{0})=0, m(\tilde{\tau}, \tilde{\tau})=0.5(m(\tilde{\chi}_{1}^{+})+m(\tilde{\chi}_{1}^{0}))$ $\tilde{\chi}_{1}^{0})=100 \text{ GeV}, m(\tilde{\tau}, \tilde{\tau})=0.5(m(\tilde{\chi}_{1}^{+})+m(\tilde{\chi}_{1}^{0}))$	1501.07110 1708.07875 1708.07875
山道	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell}{\rightarrow}\ell\tilde{\chi}_1^0$	2 e, μ 2 e, μ	0 ≥ 1	Yes Yes	36.1 36.1	₹ ₹ 0.18	0.5			$m(\tilde{\ell}_{1}^{0})=0$ $m(\tilde{\ell})-m(\tilde{\ell}_{1}^{0})=5 \text{ GeV}$	1803.02762 1712.08119
	$\tilde{H}\tilde{H}, \tilde{H} \rightarrow h\tilde{G}/Z\tilde{G}$	0 4 e, µ	$\geq 3b$ 0	Yes Yes	36.1 36.1	<u>Й</u> 0.13-0.23 <u>Й</u>	0.3	0.29-0.88		$BR(\tilde{\chi}_1^0 \rightarrow h\tilde{G})=1$ $BR(\tilde{\chi}_1^0 \rightarrow Z\tilde{G})=1$	1806.04030 1804.03602
0 0	$\operatorname{Direct} \tilde{\chi}_1^* \tilde{\chi}_1^- \operatorname{prod.}, \operatorname{long-lived} \tilde{\chi}_1^\pm$	Disapp. trk	1 jet	Yes	36.1	$ \tilde{\chi}_{1}^{\pm} \\ \tilde{\chi}_{1}^{\pm} $ 0.15	0.46			Pure Wino Pure Higgsino	1712.02118 ATL-PHYS-PUB-2017-019
Long-Ilved particles	Stable g R-hadron	SMP	- Multiple	-	3.2 32.8	ğ ğ [τ(ğ) =100 ns, 0.2 ns]			1.6	2.4 m(λ ⁰ ₁)=100 GeV	1606.05129 1710.04901, 1604.04520
pai	Metastable \tilde{g} R-hadron, $\tilde{g} \rightarrow qq \tilde{\chi}_1^0$ GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$, long-lived $\tilde{\chi}_1^0$ $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow eev/e\mu v/\mu\mu v$	2γ displ. ee/eµ/µ	-	Yes	20.3 20.3	\tilde{x}_{1}^{0} \tilde{s}	0.44		1.3	2.4 $m(x_1)=100 \text{ GeV}$ $1 < r(\tilde{\chi}_1^0) < 3 \text{ ns, SPS8 model}$ $6 < cr(\tilde{\chi}_1^0) < 1000 \text{ mm, m}(\tilde{\chi}_1^0)=1 \text{ TeV}$	1409.5542
	LFV $pp \rightarrow \tilde{\nu}_{\tau} + X, \tilde{\nu}_{\tau} \rightarrow e\mu/e\tau/\mu\tau$ $\tilde{\chi}_{\pm}^{\pm}\tilde{\chi}_{\pm}^{T}/\tilde{\chi}_{2}^{0} \rightarrow WW/Z\ell\ell\ell\ell\nu\gamma$	eμ,eτ,μτ 4 e,μ	-	- Yes	3.2 36.1	$\tilde{\nu}_{\tau}$ $\tilde{\chi}_{1}^{\pm}/\tilde{\chi}_{2}^{0}$ $[\lambda_{233} \neq 0, \lambda_{12k} \neq 0]$		0.82	1.9	λ ₃₁₁ =0.11, λ _{132/133/233} =0.07 m($\tilde{\chi}_{1}^{0}$)=100 GeV	1607.08079 1804.03602
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow qqq$		-5 large- <i>R</i> j Multiple		36.1 36.1	$\tilde{g} = [m(\tilde{k}^0)=200 \text{ GeV}, 1100 \text{ GeV} \\ \tilde{g} = [\lambda'_{112}=2e\cdot4, 2e\cdot5]$]	1.0	1.3 1.9	Large $\lambda_{112}^{\prime\prime}$ m $(\tilde{\chi}_1^0)$ =200 GeV, bino-like	1804.03568 ATLAS-CONF-2018-003
RPV	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow tbs / \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow tbs$	Multiple 3			36.1	ğ [𝑋'' ₃₂₃ =1, 1e−2]			1.8 2.1		ATLAS-CONF-2018-003
	$\tilde{t}\tilde{t}, \tilde{t} \rightarrow t\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow tbs$ $\tilde{t}_{1}\tilde{t}_{1}, \tilde{t}_{1} \rightarrow bs$	0	Multiple 2 jets + 2 i		36.1 36.7	$\tilde{g} = [\lambda''_{323}=2e-4, 1e-2]$ $\tilde{t}_1 = [qq, bs]$	0.5	5 1.0 0.61	5	m($\tilde{\chi}_1^0$)=200 GeV, bino-like	ATLAS-CONF-2018-003 1710.07171
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow b \delta$ $\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow b \delta$	2 e, µ	2 jets + 2 i 2 b	-	36.7 36.1	$\overline{t}_1 = [qq, bs]$ \overline{t}_1	0.42	0.01	0.4-1.45	$BR(\tilde{t}_1 \rightarrow be/b\mu) > 20\%$	1710.05544
phén	a selection of the available ma omena is shown. Many of the	1									
simpl	ified models. c.f. refs. for the	assumptions	made.								

Top quark final states are important part of the LHC NP search program.



Search for the stop quark of compressed SUSY scenarios in I-lepton, jets, and missing energy final state

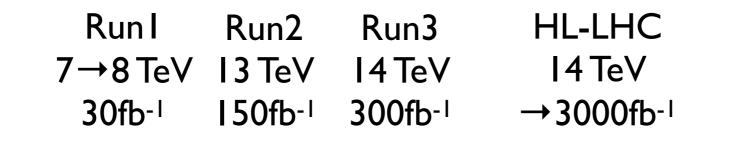


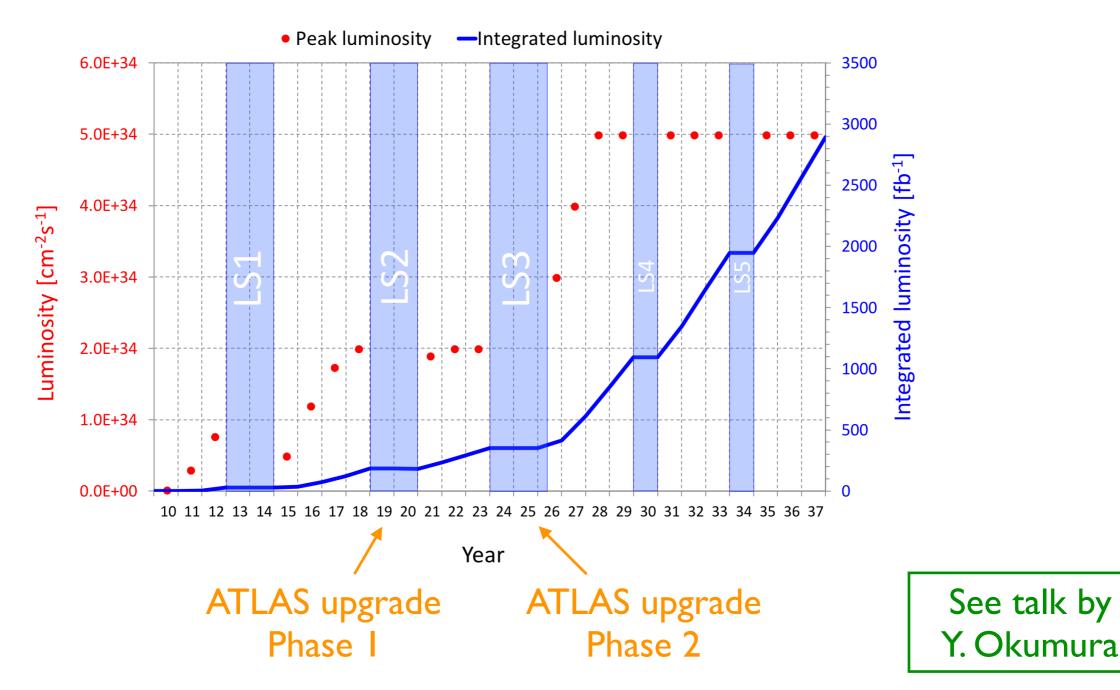
See talk by

Y. Nakahama

LHC Long-term Plan

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SuperKEKB/Belle II

New intensity frontier facility at KEK

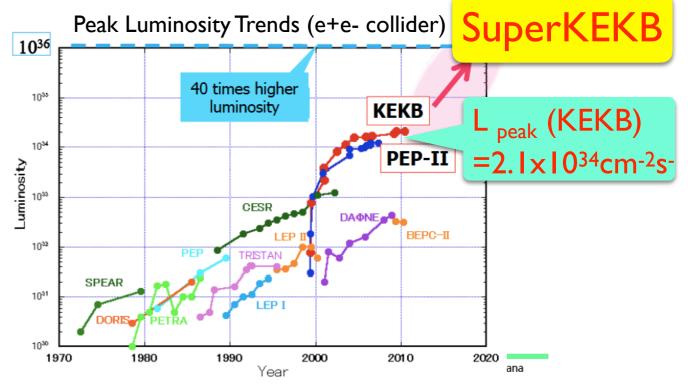
Target luminosity; L_{peak} = 8 x 10³⁵cm⁻²s⁻¹

 \Rightarrow ~1010 BB, T⁺T⁻ and charms per year !

 $L_{int} > 50 \text{ ab-I}$

- Rich physics program
 - Search for New Physics through processes sensitive to virtual heavy particles.
 - New QCD phenomena (XYZ, new states including heavy flavors) + more





The first particle collider after the LHC !

See talk by

K.Trabelsi

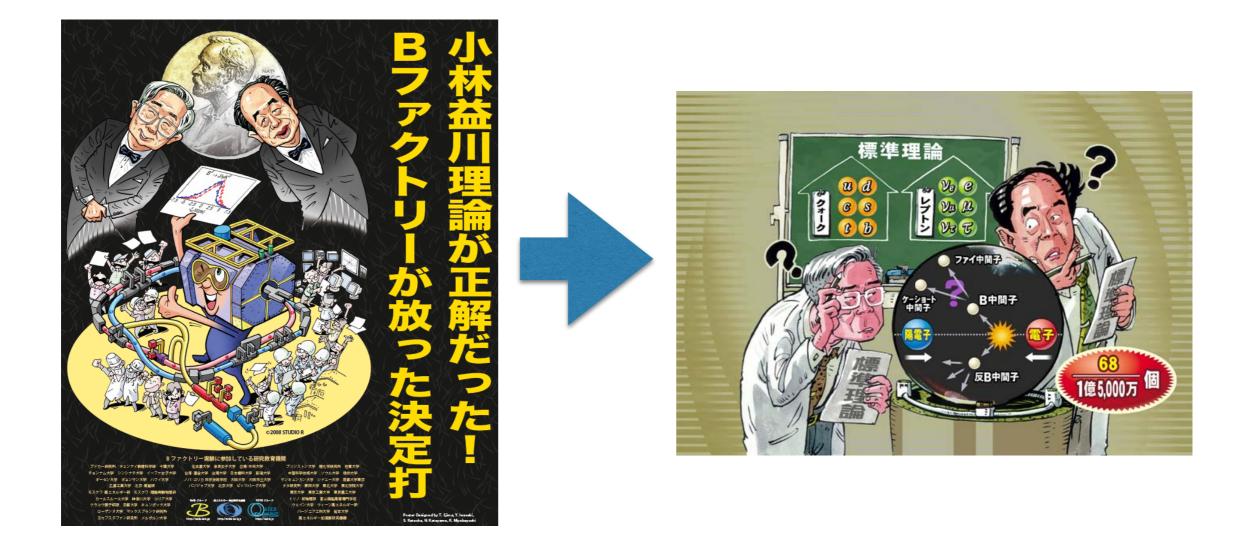
10 years later from KM's Nobel Prize, we are starting a new experiment SuperKEKB/Belle II

to search for new phenomena beyond KM.

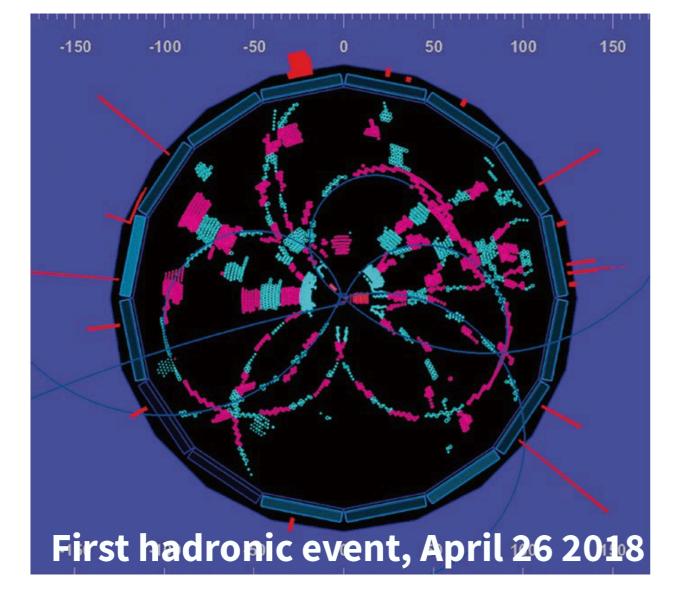


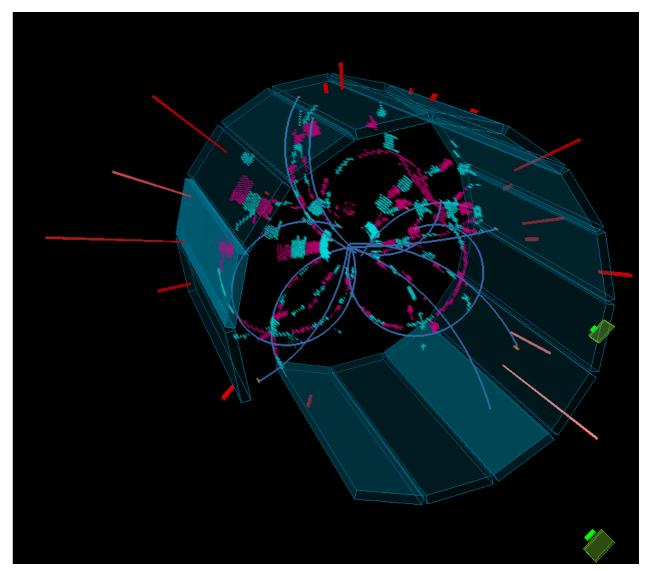
10 years later from KM's Nobel Prize, we are starting a new experiment SuperKEKB/Belle II

to search for new phenomena beyond KM.



First Collision ! 0:38, April 26, 2018





First Collision ! 0:38, April 26, 2018

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First Collision ! 0:38, April 26, 2018

Kobayashi @ First collisions ceremony

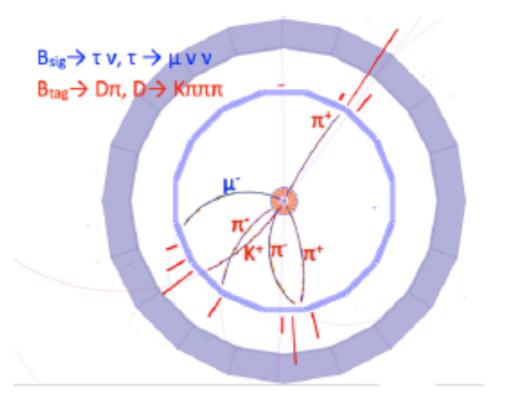
Super

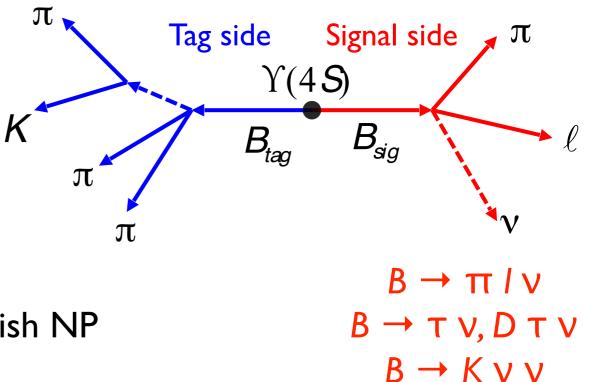
EKB

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Advantage of e⁺e⁻ Flavor Factory²⁵

- Clean environment
 - Efficient detection of neutrals (γ , π^0 , η , ...)
- Quantum correlated B⁰B⁰ pairs
 - High effective flavor tagging efficiency : ~34%(Belle II) ~3% (LHCb)
- Large sample of τ leptons
 - Search for LFV τ decays at O(10-9)
- Full reconstruction tagging possible
 - A powerful tool to measure;
 - b→u semileptonic decays (CKM)
 - decays with large missing energy
- Systematics different from LHCb
 - Two experiments are required to establish NP



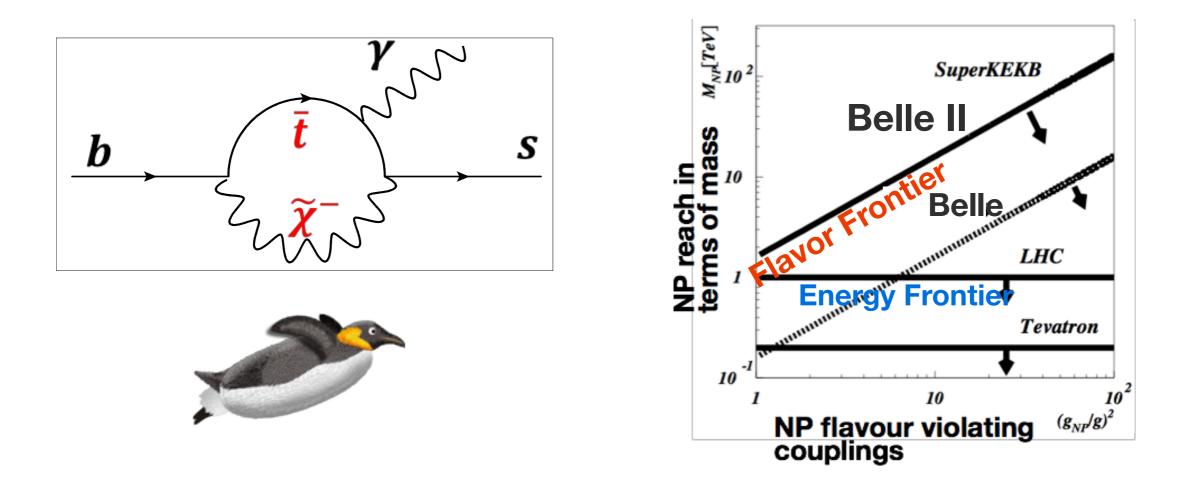


Advantage of e⁺e⁻ Flavor Factory²⁵



Role of the Belle II Experiment

- Complementary to direct search in LHC high P_T programs.
- Reach in mass scale is not limited by the collision energy.
 - Depend on NP flavor violating couplings.



Key Measurements at Belle II

arXiv:1002.5012

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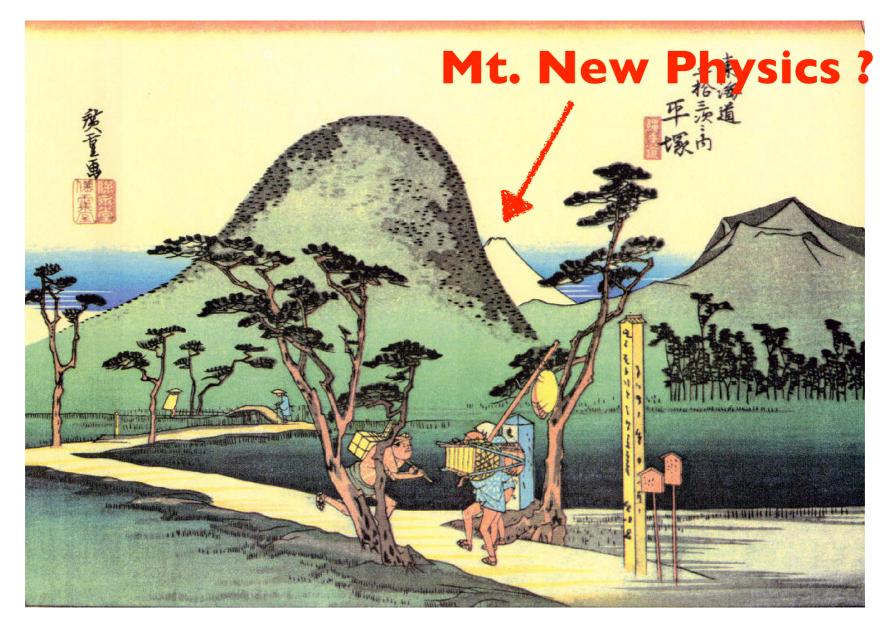
- CPV in $b \rightarrow s$ penguin decays
- FCNC
- Tauonic decays
- LFV T decays



Observable	Belle 2006	SuperK	EKB
	$(\sim 0.5 \text{ ab}^{-1})$	(5 ab^{-1})	(50 ab^{-1})
Hadronic $b \rightarrow s$ transitions			
$\Delta S_{\phi K^0}$	0.22	0.073	0.029
$\Delta S_{\eta' K^0}$	0.11	0.038	0.020
$\Delta S_{K_g^0 K_g^0 K_g^0}$	0.33	0.105	0.037
$\Delta A_{\pi^0 K_S^0}$	0.15	0.072	0.042
$\mathcal{A}_{\phi\phi K^+}$	0.17	0.05	0.014
$\phi_1^{eff}(\phi K_S)$ Dalitz		3.3°	1.5°
Radiative/electroweak $b \rightarrow s$ transitions			
$S_{K_{\pi}^{0}\pi^{0}\gamma}$	0.32	0.10	0.03
$\mathcal{B}(\ddot{B} \rightarrow X_s \gamma)$	13%	7%	6%
$A_{CP}(B \rightarrow X_s \gamma)$	0.058	0.01	0.005
$C_9 \text{ from } \overline{A}_{FB}(B \to K^* \ell^+ \ell^-)$	-	11%	4%
$C_{10} \text{ from } \overline{A}_{FB}(B \to K^* \ell^+ \ell^-)$	-	13%	4%
C_7/C_9 from $\overline{A}_{FB}(B \to K^* \ell^+ \ell^-)$	-		5%
R_K		0.07	0.02
$\mathcal{B}(B^+ \to K^+ \nu \nu)$	†† < 3 $\mathcal{B}_{\rm SM}$		30%
$\mathcal{B}(B^0 \to K^{*0} \nu \bar{\nu})$	$^{\dagger\dagger} < 40 \mathcal{B}_{SM}$		35%
Radiative/electroweak $b \rightarrow d$ transitions			
$S_{\rho\gamma}$	-	0.3	0.15
$\mathcal{B}(B \rightarrow X_d \gamma)$	-	24% (syst.)	
Leptonic/semileptonic B decays			
$\mathcal{B}(B^+ \to \tau^+ \nu)$	3.5σ	10%	3%
$\mathcal{B}(B^+ \to \mu^+ \nu)$	$^{\dagger\dagger} < 2.4 B_{SM}$	4.3 ab^{-1} for 5σ discovery	
$\mathcal{B}(B^+ \to D \tau \nu)$	-	8%	3%
$\mathcal{B}(B^0 \to D \tau \nu)$	-	30%	10%
LFV in τ decays (U.L. at 90% C.L.)			
$\mathcal{B}(\tau \to \mu \gamma) \ [10^{-9}]$	45	10	5
$\mathcal{B}(au o \mu \eta) \ [10^{-9}]$	65	5	2
$\mathcal{B}(au o \mu \mu \mu) \ [10^{-9}]$	21	3	1

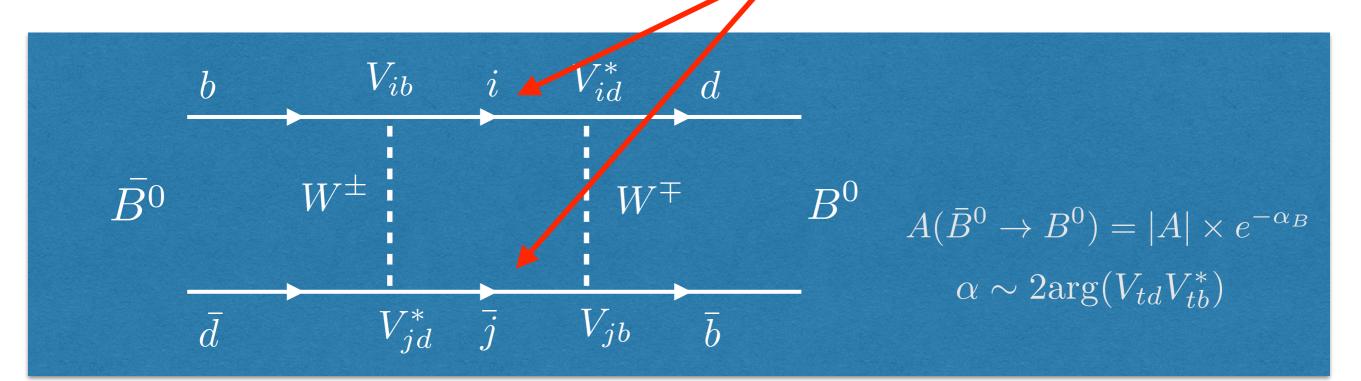
Ultimate measurements down to theory error !

Physics w/ b-quark We know this old road...



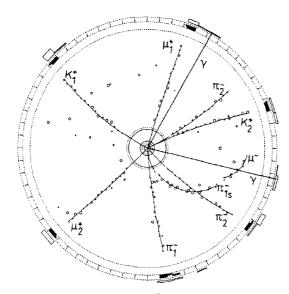
by Hiroshige Utagawa (1797-1858)

B-B Mixing and Top Quark



OBSERVATION OF B⁰-B⁰ MIXING

ARGUS Collaboration



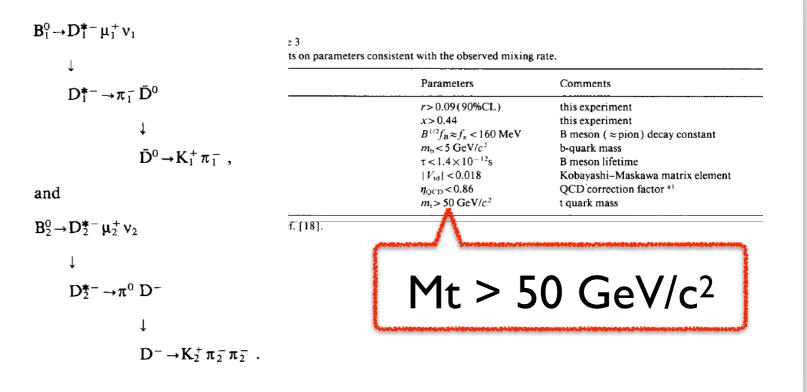
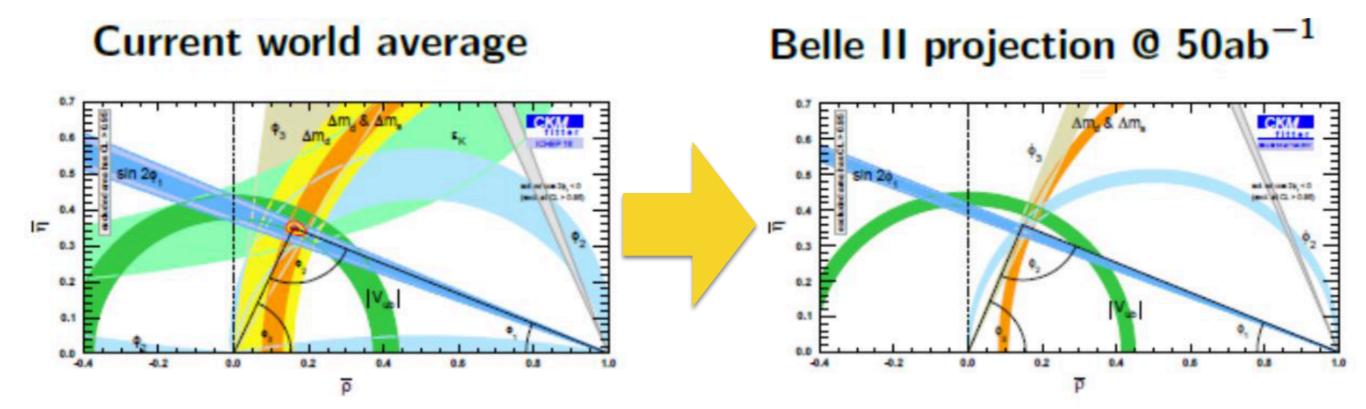


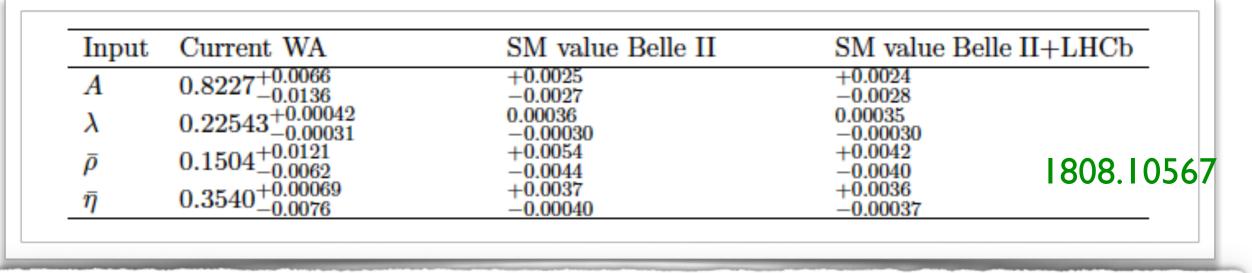
Fig. 2. Completely reconstructed event consisting of the decay Υ (4S) \rightarrow B⁰B⁰.

CKM fit w/ Belle II + LHCb

A	$0.8227^{+0.0066}_{-0.0126}$	+0.0025	+0.0024
	$0.22543^{+0.00042}_{-0.00031}$	-0.0027 0.00036 -0.00030	-0.0028 0.00035 -0.00030
	$0.1504_{-0.0062}^{+0.0121}$	$+0.0054 \\ -0.0044$	$+0.0042 \\ -0.0040$ [808]0 ⁴
$ar\eta$	$0.3540^{+0.00069}_{-0.0076}$	+0.0037 -0.00040	+0.0036 -0.00037

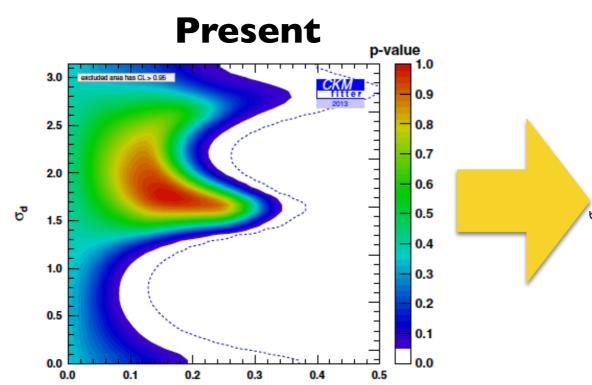


CKM fit w/ Belle II + LHCb



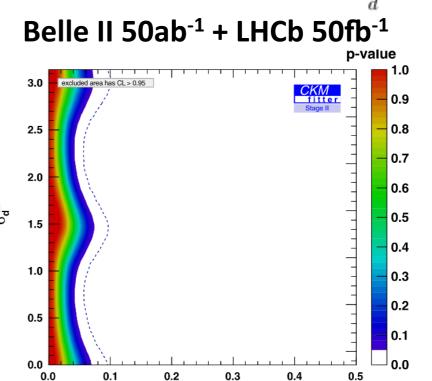
$$M_{12}^{d,s} = (M_{12}^{d,s})_{\rm SM} \times (1 + h_{d,s}e^{2i\sigma_{d,s}})$$

Relative amplitude (h) phase (σ)



h,

15



h_d

b

u, c, t

W

W

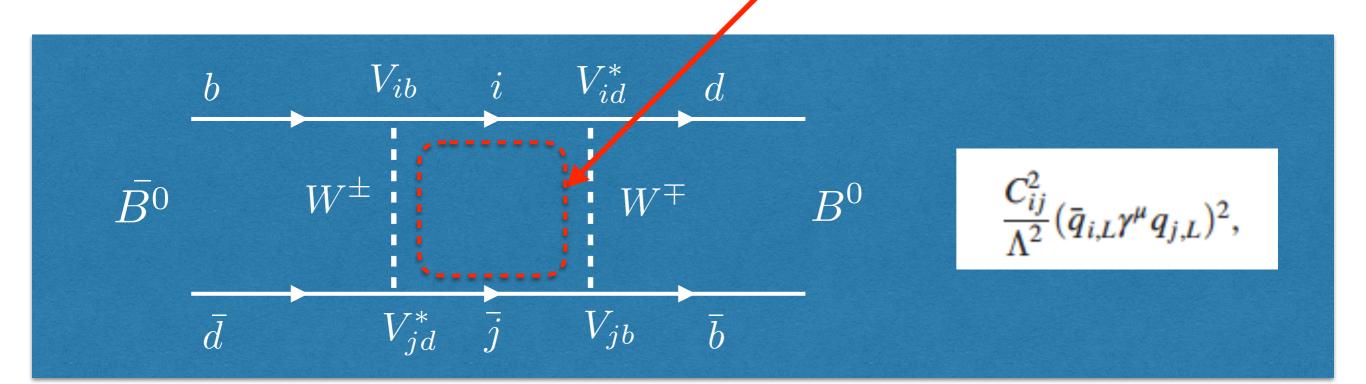


d

u, c, t

 \overline{b}

B-B Mixing and New Physics



$$\begin{split} h &\simeq 1.5 \frac{|C_{ij}|^2}{|\lambda_{ij}^t|^2} \frac{(4\pi)^2}{G_F \Lambda^2} \simeq \frac{|C_{ij}|^2}{|\lambda_{ij}^t|^2} \left(\frac{4.5 \text{ TeV}}{\Lambda}\right)^2, \\ \sigma &= \arg(C_{ij} \lambda_{ij}^{t*}), \\ \lambda_{ij}^t &= V_{ti}^* V_{tj} \end{split}$$

J. Charles et al., PRD89,033016(2014)

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Couplings	NP loop	Scales (in TeV) probed by	
couplings	order	B_d mixing	B_s mixing
$ C_{ij} = V_{ti}V_{tj}^* $	tree level	17	19
(CKM-like)	one loop	1.4	1.5
$ C_{ij} = 1$	tree level	2×10^3	5×10^2
(no hierarchy)	one loop	2×10^2	40

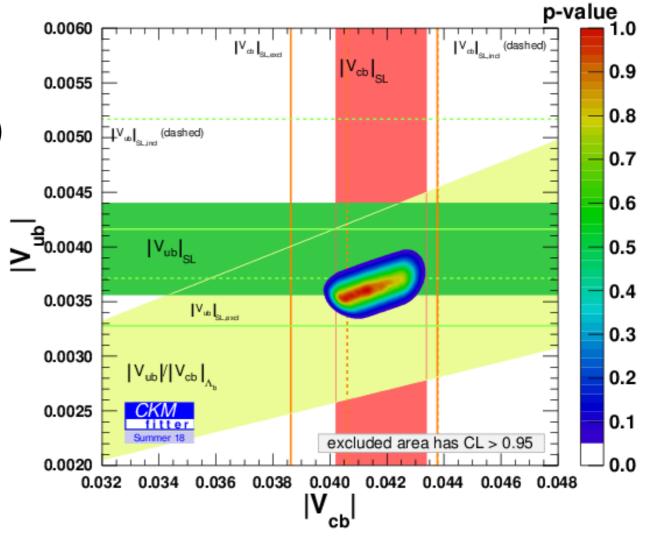
Mass reach (CKM-like): O(I)TeV $\rightarrow O(I0)$ TeV !

Some theoretical issues

 Hadronic uncertainties to extract fundamental quantities from experimentally measured rates.

• ex.
$$|V_{c(u)b}|$$
 from $\Gamma(B \rightarrow X_{c(u)} | v)$.

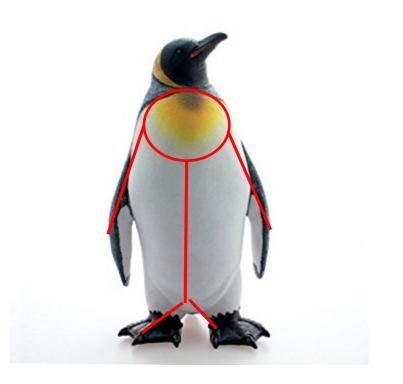
- Form factors $f(B \rightarrow D)$, $f(B \rightarrow \pi)$
- Inclusive vs Exclusive
- Need more theoretical investigations.

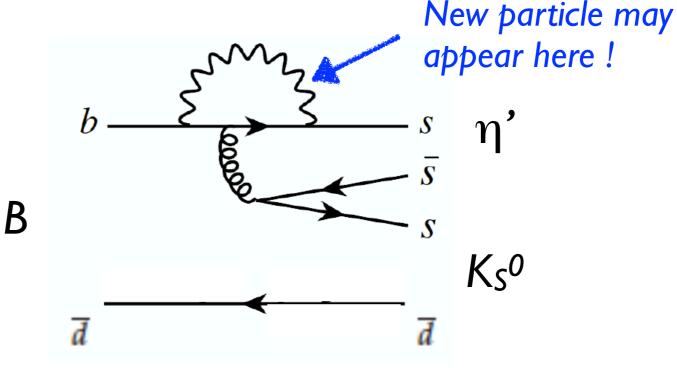


CPViolation by New Physics

Belle II can measure types of B meson decays which rarely happen (~ one per million B decays), known as "Penguin" decays.

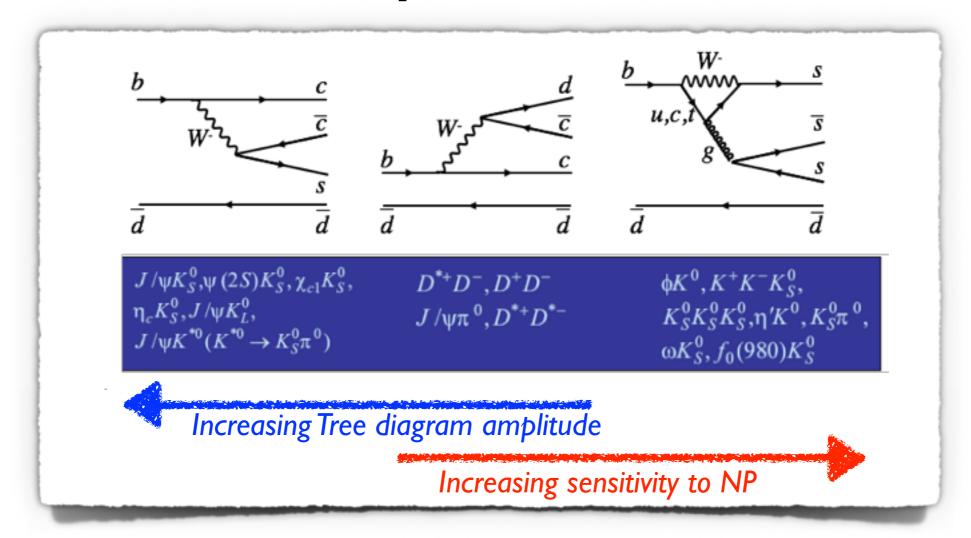
Q: Does CP violation in "Penguin" decays deviates from the SM ?



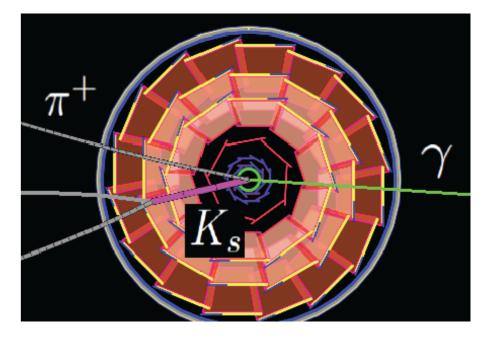


A. Gaz is leading the physics analysis in Belle II

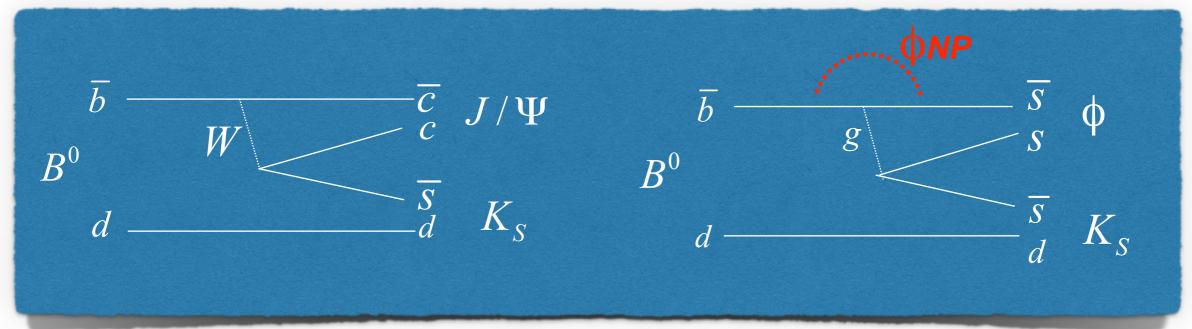
Time-dependent CPV



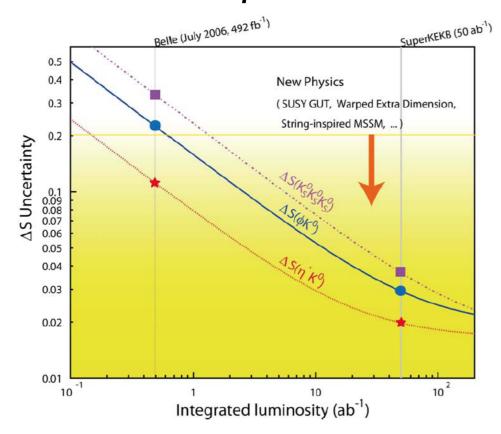
- Larger acceptance for Ks decay vertex +30%
- Improved vertex resolution $\sigma(Z) \sim 18 \mu m @$ Belle II $\leftrightarrow ~61 \mu m @$ Belle
 - → less systematic error

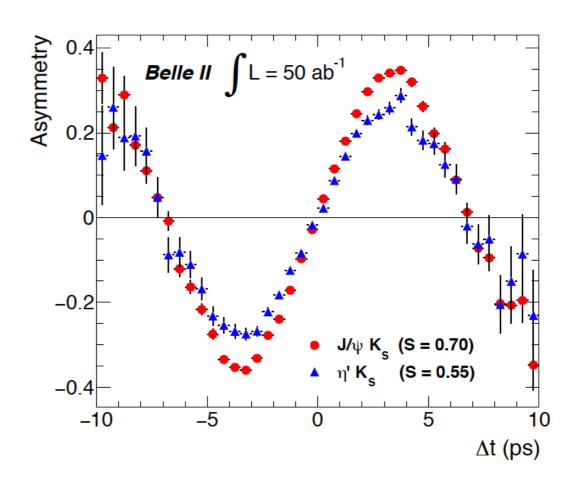


$\Phi_{I} (\equiv \beta)$ Projection



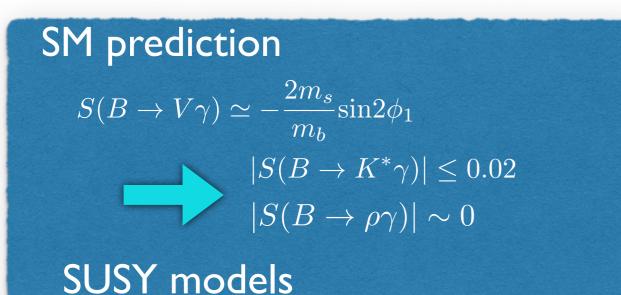
Belle II projection for $sin 2\varphi_1^{eff}$ from $b \rightarrow s \bar{s} s processes$

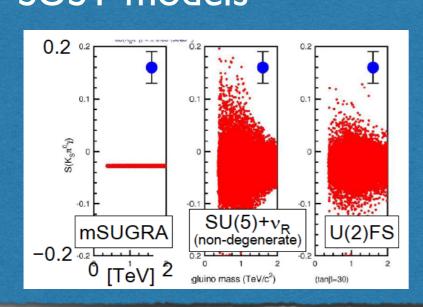


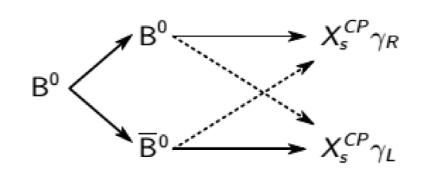


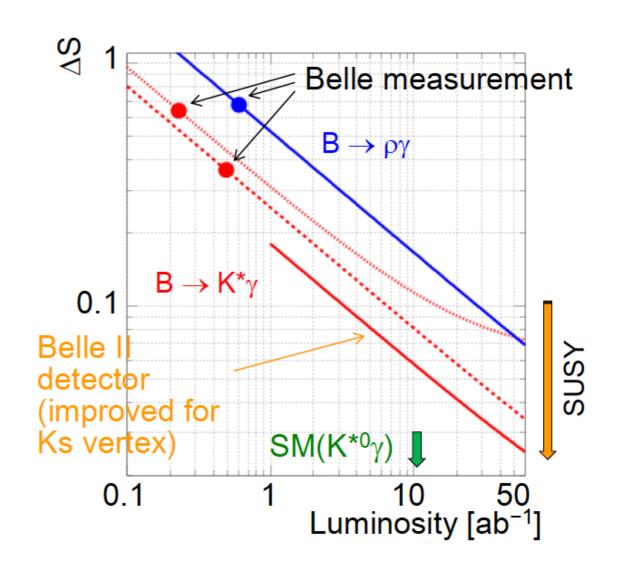
Time-dep. CPV in $b \rightarrow s, d + \gamma$

- In SM, photon from b→s,d + γ is almost left-handed.
- Right-handed photon causes interference, and large CPV.





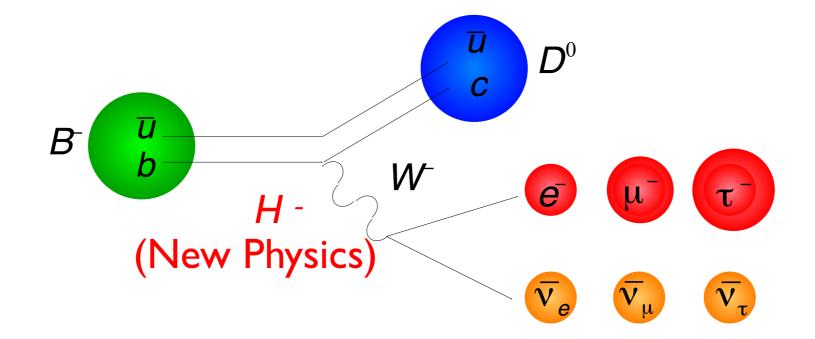




Lepton non-Universality (tree)

There are 3 modes in the B meson weak decay into the lepton final states; $B \rightarrow D \in v$, $D \mu v$, $D \tau v$.

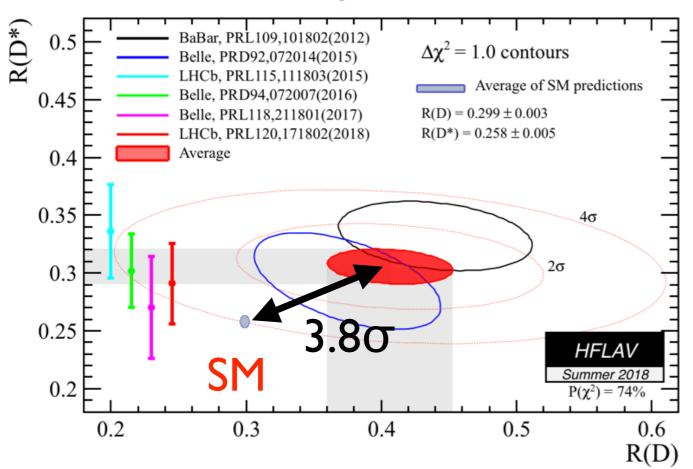
Q: Are they just the same (as in the SM) ?



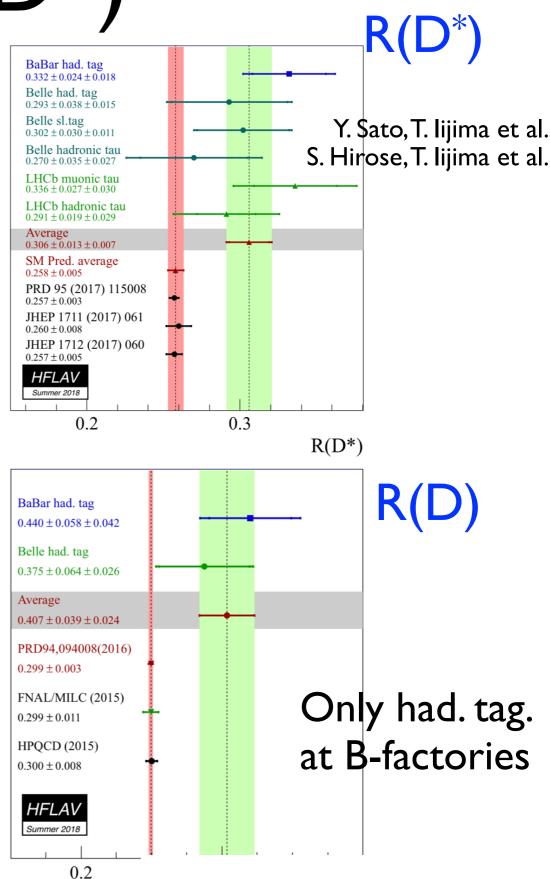
Hints of deviations in recent BaBar, Belle, LHCb data ...

R(D), R(D*)

Summer 2018 update



Deviation from SM slightly decreased from 4.1 \rightarrow 3.8 σ , mainly due to change in theoretical SM prediction.



Belle II Projections

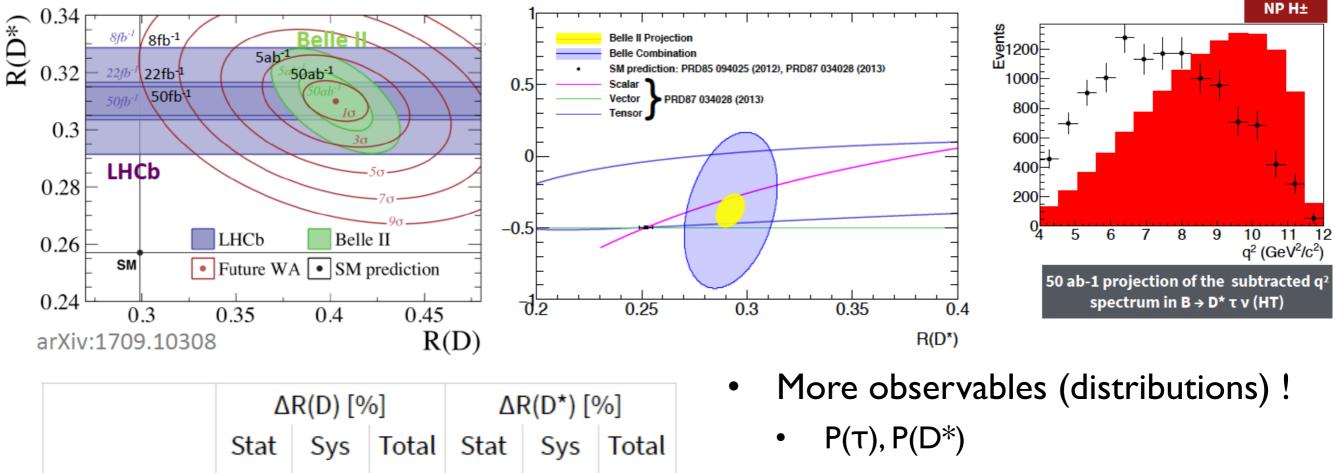
- Lepton universality violation may be established even with 5ab⁻¹ (2020).
- High statistics data will provide more detailed information, such as τ polarization, q² distribution, to discriminate type of NP.

Belle 0.7 ab⁻¹

Belle II 5 ab⁻¹

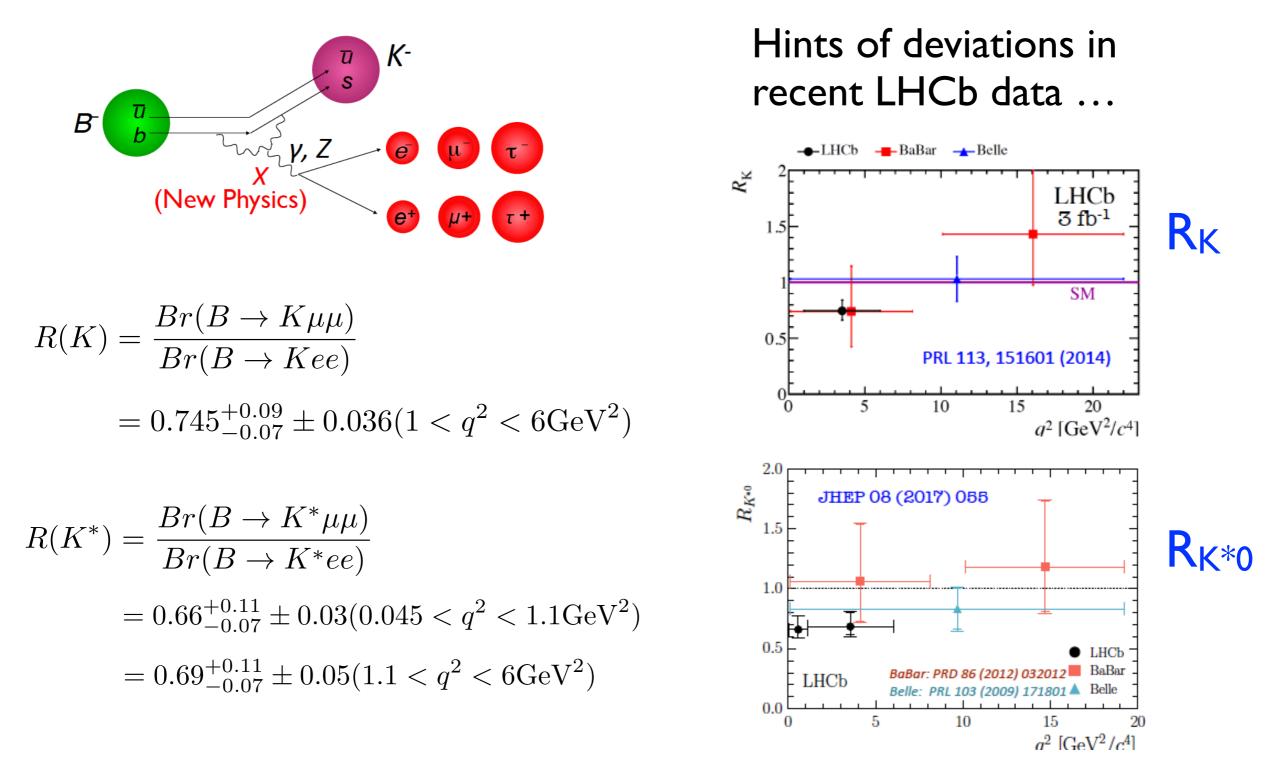
Belle II 50 ab⁻¹

Will soon hit the systematic limit !



- $d\Gamma/dq^2$, $d\Gamma/dp_{D(*)}$, $d\Gamma/dp_e$, ...
- More modes !
 - $B \rightarrow \pi \tau \nu$,
 - $B_S \rightarrow D_S \tau \nu$ (at 5S runs), ...

Lepton non-Universality (loop)



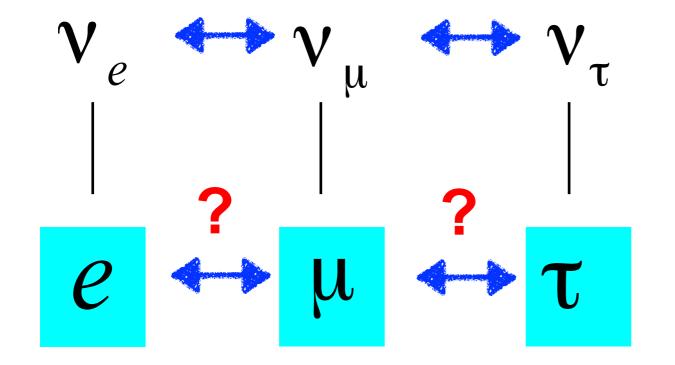
Precision at Belle II : \sim 3% at 50ab⁻¹ and also inclusive measurement (less theory ambiguity) possible.

Lepton Flavor Violation



SuperKEKB produces also a lot of tau-leptons, which can be studies in detail by Belle II.

Q: Does the tau-lepton changes to the muon or electron, similarly to the case of the neutrino (neutrino oscillation).



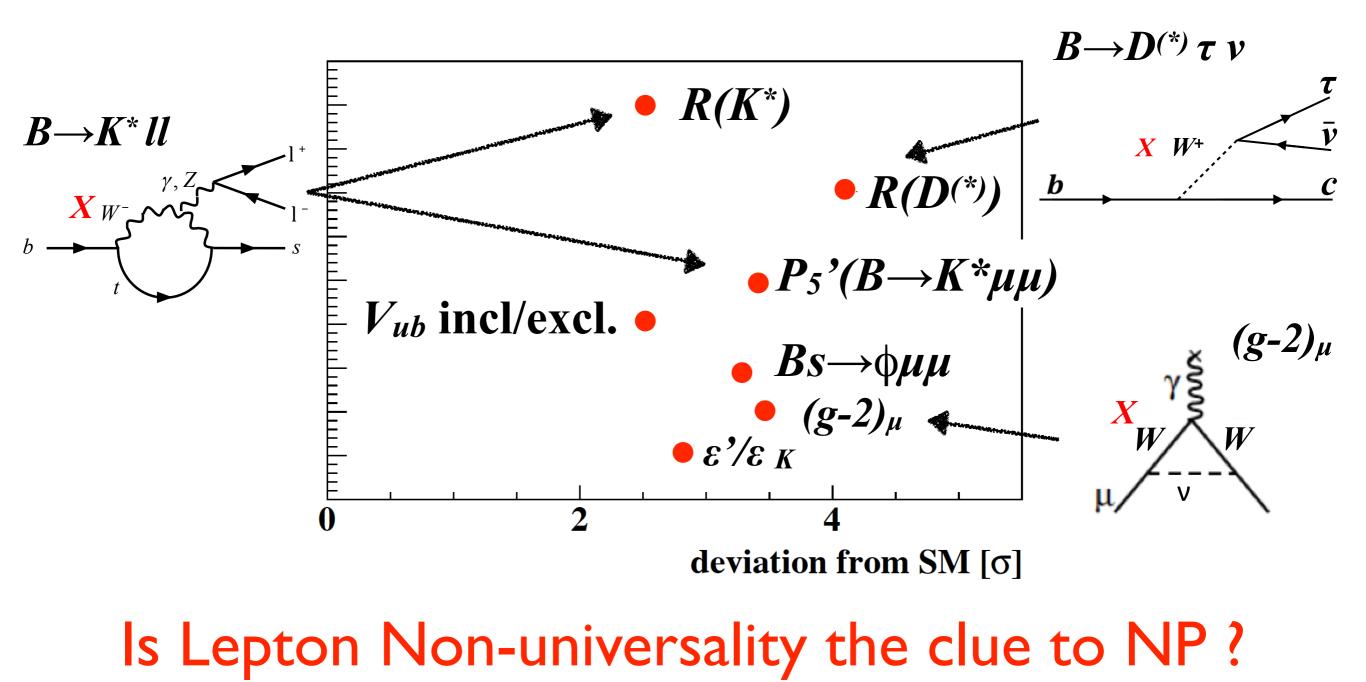
neutrino oscillation

are there decays like $\tau \rightarrow \mu \gamma$? $\tau \rightarrow e \gamma$?

Lepton Non Universality

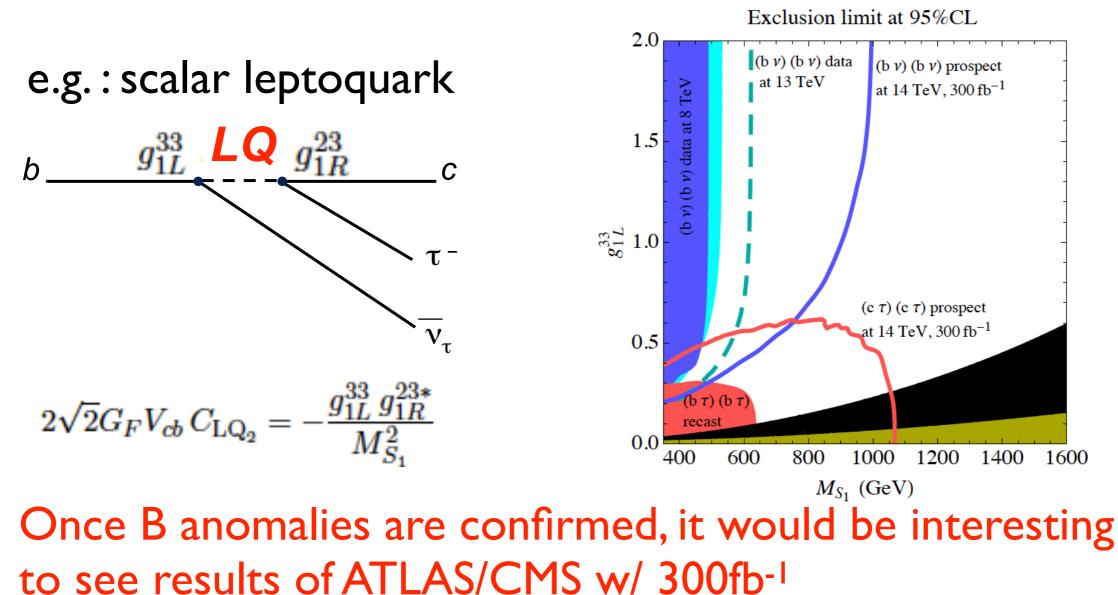
42

Observed deviation from SM



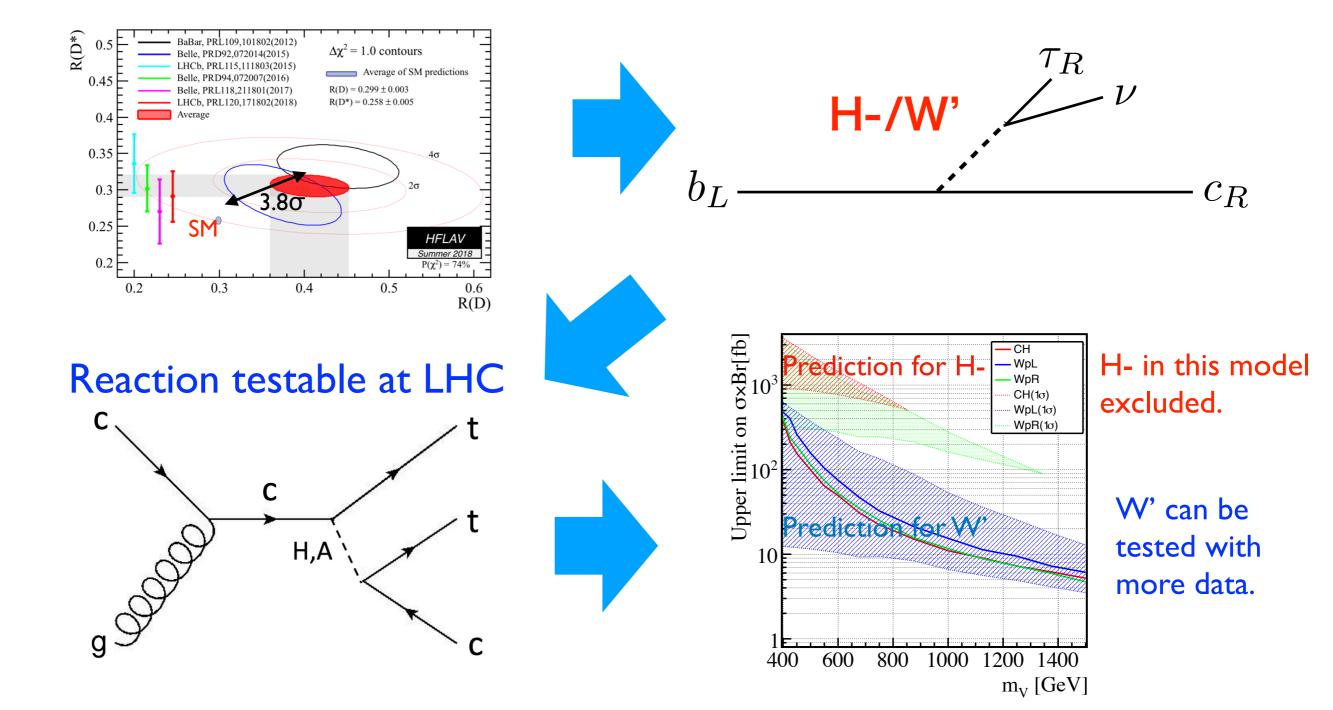
Testing B anomalies at ATLAS/ ⁴³ CMS (e.g. LQ model) ^{1808.10567}

- The Leptoquark (LQ) model is a favored model, which can explain observed anomalies consistently: P5', R_K(*), R(D(*))
 - Coupling to 3rd gen. > to 2nd gen. >> to 1st gen.



Work by Nagoya Theory Group

 Building a model, which can explain the "B anomalies", and predict and suggest tests at LHC.



World Research Unit for Heavy Flavor Particle Physics ("WPI-next")



Toru lijima •B,Tau Physics •Exotic hadrons



Alessandro Gaz



Kodai Matsuoka



Yuji Omura

Theory Junji Hisano Flavor Physics Dark Matter





Makoto Tomoto •Top physics •Higgs



Peter Krizan (Ljubljana)



Tim Gershon (Warwick)

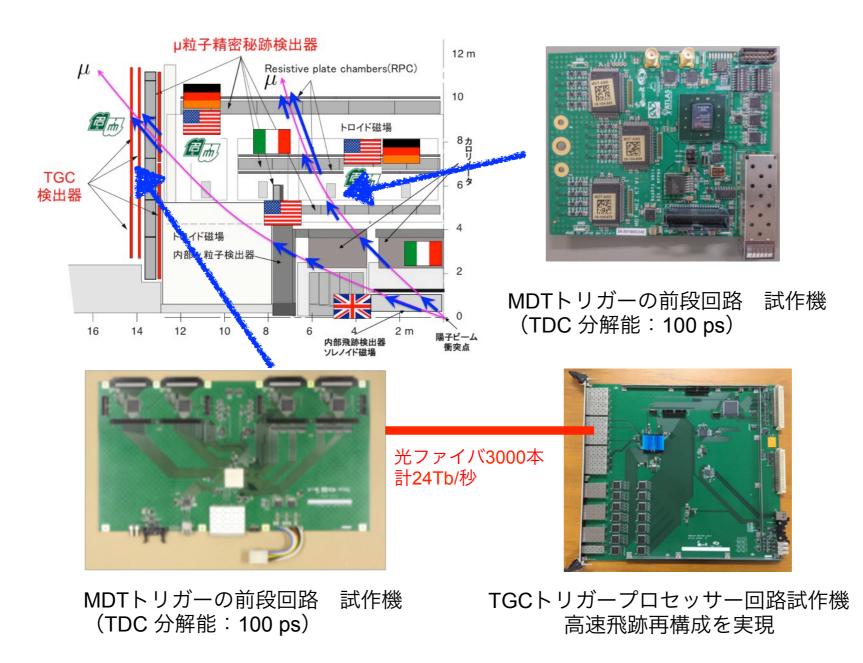
Gino Isidori (Zurich)



Yu Nakahama

R&D for ATLAS Muon Trigger

HL-LHC advanced muon trigger Realized high speed muon track trigger by processing IM channel signals with large-scale FPGA.



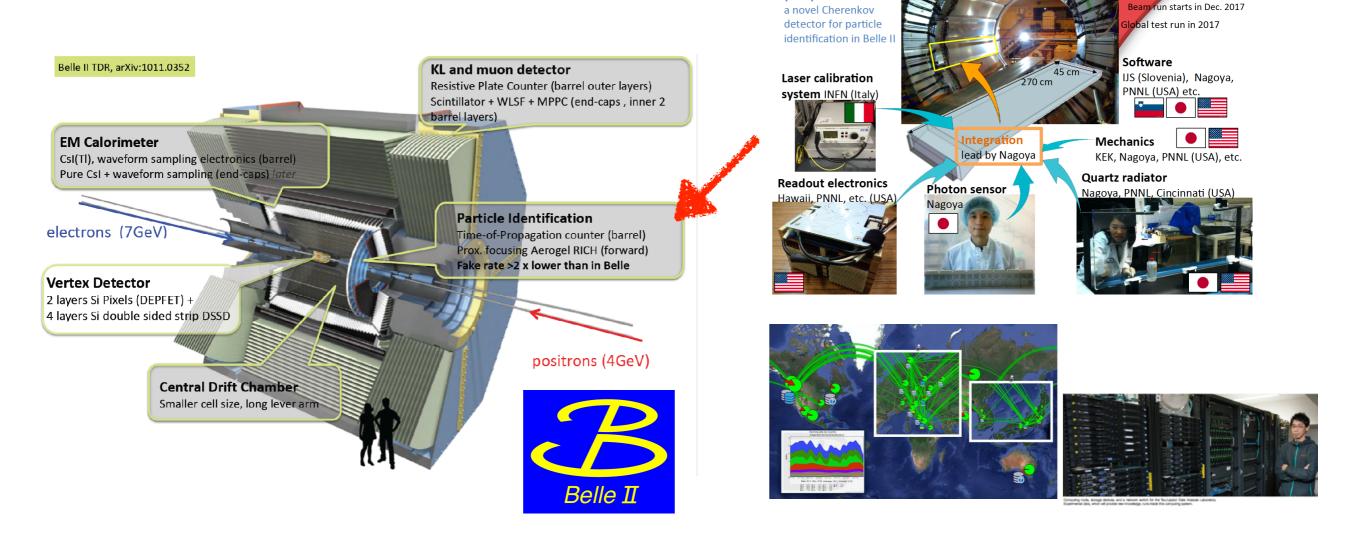
Belle II Experiment

Belle II TOP counters were successfully built in May 2016

Time-Of-Propagation

(TOP) counter is

- Deal with higher background (10-20×), radiation damage, higher occupancy, higher event rates (L1 trigg. $0.5 \rightarrow 30$ kHz)
- Improved performance and hermeticity



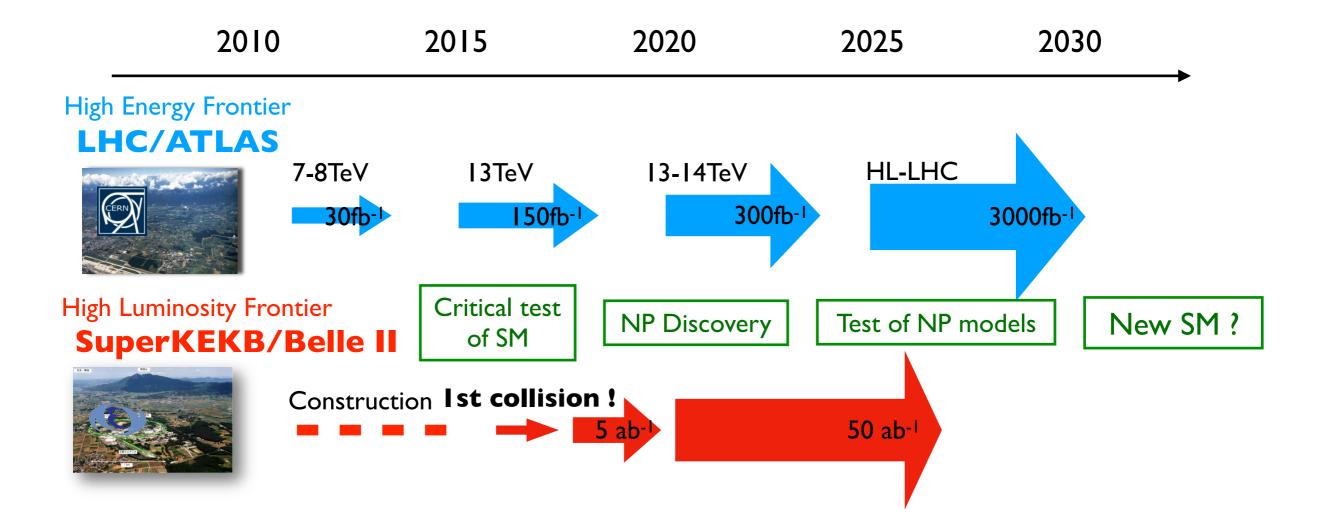
Nagoya group takes leading roles: Particle ID, Computing, Physics analyses T.I will be the spokesperson from June, 2019.

New Physics

n starts in Dec. 2018

Summary

- The CPV phenomena observed in flavor changing quark interactions (in K, B decays so far) can be explained by the Kobayashi-Maskawa.
- But, we are still far from explaining the Matter-Antimatter asymmetry in the Universe.
- Search for New Physics is the clue to investigate more the issue.
- Researches on heavy flavors are important in coming years.



Backup Slides

