Angular Observables Sensitive to NP in $\bar{B} \longrightarrow D^* l^- \bar{\nu}_l$

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Intro I

- Evidence for NP in $\bar{B} \longrightarrow D^* \tau^- \bar{\nu}_{\tau}$ is mounting
 - \implies Next question: What kind of NP?
 - \implies New observables! (also good as cross checks)
- Here we consider the angular distribution of $\overline{B} \longrightarrow D^*(\rightarrow D\pi) l^- \overline{\nu}_l$. Example: Forward-backward asymmetry:

$$A_{ heta} := \left[\int_{-1}^{0} - \int_{0}^{1}
ight] \mathrm{d} \mathrm{cos}\, heta_{\ell} rac{\mathrm{d} \Gamma}{\mathrm{d} \mathrm{cos}\, heta_{\ell}}$$

Generalization: Observables \mathcal{O}_a built from *binned* measurements of $\frac{d^3\Gamma}{d\chi \, d\theta_\ell \, d\theta_{D^*}}$ Similar to 12 observables constructed in 1602.03030¹ (using polarizations in addition to angles)

¹Becirevic, Fajfer, Nisandzic, Tayduganov

Intro II

However (so far): Only proofs of concepts, little experimental considerations!

Aim:

- + Characterize general form of 12 observables \mathcal{O}_a
- + Introduce experimental error estimate
- + Minimize expected errors using d.o.f. in construction of \mathcal{O}_a
- + Minimal number of bins required for construction of \mathcal{O}_a
- + Optimal bin spacing for each \mathcal{O}_a
- + Consider discriminatory power of observables on basis of operators corresponding to relevant NP mediators

The Differential Cross Section

Slightly adapted² from 1405.3719^3 , we have⁴ (with (pseudo)scalar, (axial) vector and tensor NP operators, no *CP* average)

$$\frac{\mathsf{d}^{3}\mathsf{\Gamma}}{\mathsf{d}\chi\,\mathsf{d}\theta_{\ell}\,\mathsf{d}\theta_{D^{*}}} = \sum_{a} \overline{W}_{a}\,B_{a}(\chi,\theta_{\ell},\theta_{D^{*}})$$

With 12 different

- Angle dependencies $B_a(\chi, \theta_\ell, \theta_{D^*})$ (known) e.g. $\cos(2\chi) \sin^3(\theta_{D^*}) \sin^3(\theta_\ell)$
- NP dependent coefficients
 W_a (unknown)



Fig. 1: The decay angles χ , θ_{ℓ} , θ_{D^*}

²In relation to the V_a from 1405.3719: $W_1^0 := V_1^0 + V_2^0$, $W_1^T := V_1^T + V_2^T$, $W_2^0 := -2V_2^0$, $W_2^T := -2V_2^T$; for the rest $W_a := V_a$. Furthermore $\overline{W}_a := \int dq^2 W_a(q^2)$. ³Duraisamy, Sharma, Datta ⁴the kinematic variable q^2 will not be considered, i.e. the relevant quantities are integrated over q^2

The Idea (I)

How to extract the \overline{W}_a ? ' Linear algebra: Eq. (1): *N* linear equations with 12 unknowns \implies there are weights $\omega_a(i)$ such that

$$\overline{W}_{a} = \sum_{i} \omega_{a}(i) \prod_{\substack{\text{unknown} \\ \text{NP-dependent} \\ \text{coefficient}}} \Gamma(i)$$

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The Idea (II)

The observables

$$\mathcal{O}_{a}: \underset{CS}{\overset{\Gamma}{\underset{i}\underset{weights}{\text{homod}}}} \mapsto \sum_{i} \omega_{a}(i) \Gamma(i) \text{ such that } \mathcal{O}_{a}(\Gamma_{\text{theo}}) = \overline{W}_{a}$$

Sensitivity

Expected uncertainty $\sqrt{Var(\mathcal{O}_a)}$ as figure of merit \rightarrow the smaller the better (currently simplified setup with only statistical errors, studies of bin migration and flat systematics ongoing)

Degrees of freedom

If more bins than required

- \implies DOFs in weights $\omega_a(i)$
- \Longrightarrow Use to improve sensitivity

Assumptions

- 1 "General bins": Arbitrary subsets $U_i \subseteq [-\pi, \pi] \times [0, \pi] \times [0, \pi]$
- 2a "Product bins": $U_{ijk} = U_i^{\chi} \times U_j^{\theta_\ell} \times U_k^{\theta_{D^*}}$
- 2b "Product weights": Product bins with weights in product form

Minimal number of bins $(l = \tau)$ (under conservative assumptions for the binning)

while number of bins in $\chi \times \theta_{\ell} \times \theta_{D^*}$ required to construct Θ_a									
\mathcal{O}_{a}	\mathcal{O}_1^0 , \mathcal{O}_2^0 , \mathcal{O}_3^0 , \mathcal{O}_1^T , \mathcal{O}_3^T	\mathcal{O}_2^T	$\mathcal{O}_4^{\mathcal{T}}$, $\mathcal{O}_5^{\mathcal{T}}$	\mathcal{O}_1^{0T} , \mathcal{O}_2^{0T} , \mathcal{O}_3^{0T} , \mathcal{O}_4^{0T}					
Minimal # bins	$1\times 3\times 3$	$3\times 3\times 2$	$3\times1\times2$	$2 \times 2 \times 3$					
	$1\times5\times2$		$3\times 3\times 1$	$2\times5\times1$					
	3 imes 3 imes 2		$5\times1\times1$	3 imes 2 imes 2					
				$3\times 3\times 1$					
				5 imes 2 imes 1					

Min. number of bins in $\chi imes heta_\ell imes heta_{D^*}$ required to construct \mathcal{O}_a

All observables can be constructed for $3\times3\times2$ binning.

Performance Comparison $(I = \tau)$



Finding optimal bin edges (example)

Example: $3\times3\times2$ bins (prod bins, gen. weights) with edge points

$$\{-\pi, \chi_1, \chi_2, \pi\}_{\chi} \times \{\mathbf{0}, \theta_{\ell 1}, \theta_{\ell 2}, \pi\}_{\theta_{\ell}} \times \{\mathbf{0}, \underline{\Delta \theta_{D^*}}, \pi\}_{\theta_{D^*}}, \pi\}_{\theta_{D^*}}$$

- \implies How does the optimal achievable sensitivity depend on $\Delta \theta_{D^*}$?
- \longrightarrow Generate sample of toy binnings (i.e. random values $\chi_1, \chi_2, \theta_{\ell 1}, \theta_{\ell 2}, \Delta_{\theta_{D^*}}$)



What can we conclude about NP contributions? (I)

Suppose we can extract all \overline{W}_a from the angular distribution

Split up \overline{W}_a :

$$\overline{W}_{a} = \sum_{i=1}^{13} \overline{W}_{a}^{(i)} c_{i}(g_{A}, g_{V}, g_{S}, g_{P}, T_{L})$$

$$\underset{\text{coeff} \in \mathbb{R}}{\underset{\text{quadratic in NP coupl.}}{\underset{\text{quadratic in NP coupl.}}}$$

Want to extract c_i given \overline{W}_a :

- System of linear equations
- Sometimes $\overline{W}_a \propto c_i \Longrightarrow$ easy!
- Generally hard ⇒ additional assumptions for NP coefficients

Effective Lagrangian:

$$\begin{aligned} \mathcal{H}_{\text{eff}} &= \frac{4G_F V_{cb}}{\sqrt{2}} \Bigg[(1 + \mathcal{C}_{V_L} [\bar{c}\gamma_\mu P_L b] [\bar{l}\gamma^\mu P_L \nu_l] + \mathcal{C}_{V_R} [\bar{c}\gamma^\mu P_R b] [\bar{l}\gamma_\mu P_L \nu_l] + \\ &+ \mathcal{C}_{S_L} [\bar{c}P_L b] [\bar{l}P_L \nu_l] + \mathcal{C}_{S_R} [\bar{c}P_R b] [\bar{l}P_L \nu_l] + T_L [\bar{c}\sigma^{\mu\nu} P_L b] [\bar{l}\sigma_{\mu\nu} P_L \nu_l] \Bigg], \end{aligned}$$

With $g_{V,A} = C_{V_R} \pm C_{V_L}$, $g_{S,P} = C_{S_R} \pm C_{S_L}$, $T_L = C_T$, $P_{L/R} := (1 \mp \gamma_5)/2$, $\sigma_{\mu\nu} := i(\gamma_{\mu}\gamma_{\nu} - \gamma_{\nu}\gamma_{\mu})/2$.

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With $g_{V,A} = C_{V_R} \pm C_{V_L}$, $g_{S,P} = C_{S_R} \pm C_{S_L}$, $T_L = C_T$, $P_{L/R} := (1 \mp \gamma_5)/2$, $\sigma_{\mu\nu} := i(\gamma_{\mu}\gamma_{\nu} - \gamma_{\nu}\gamma_{\mu})/2$.

What can we conclude about NP contributions? (II)

Ci	In terms of coupling const	ants
c 1	$ 1-g_A ^2$	
C 2	$ g_V + 1 ^2$	
C 3	$ g_P ^2$	
C 4	$ T_L ^2$	
C 5	$Re((1-{ extsf{g}}_{ extsf{A}})({ extsf{g}}_{ extsf{V}}^*+1))$	
<i>c</i> ₆	$Re((1 - \underline{g}_A)\underline{g}_P^*)$	
C 7	$Re((1-g_A)T_L^*)$	
C 8	$Re((g_V+1)T_L^*)$	
C 9	$\operatorname{Re}(g_P T_L^*)$	
c_{10}	$Im((1-\underline{g_A})(\underline{g_V^*}+1))$)
c ₁₁	$Im((1 - g_A)T_L^*)$	
<i>c</i> ₁₂	$Im((\underline{g}_V+1)\underline{g}_P^*)$	<pre>drop out in CP average</pre>
<i>c</i> ₁₃	$\operatorname{Im}(g_P T_L^*)$	J

Fig. 2: $c_i \longrightarrow$ coupling constants

What can we conclude about NP contributions? (III)

Consider base of operators with ${}^{\rm 5}$

- Renormalizable couplings relevant for $\bar{B} \longrightarrow D^* l^- \bar{\nu}_l$
- Dimension 4 or 6
- $\blacksquare \ \ \text{Non-flavor universal couplings} \Longrightarrow \mathsf{right-handed vector current not considered}$

Denote operators by transformation under $SU(3)_C \times SU(2)_W \times U(1)_Y$:



⁵source: Freytsis, Ligeti, Ruderman 1506.08896

What can we conclude about NP contributions? (IV)

Considering base of operators corresponding to relevant mediators

	SM	All NP	$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	$({f 1},{f 2})_{1/2}$, $({f \bar 3},{f 2})_{5/6}$	(3,1) _{2/3}	(3,2) _{7/6}	$(\bar{3},1)_{1/3}$
W_1^0		c_1, c_3, c_4, c_6, c_7	<i>c</i> 1	<i>c</i> ₃ , <i>c</i> ₆	<i>c</i> ₁ , <i>c</i> ₃ , <i>c</i> ₆	<i>c</i> ₃ , <i>c</i> ₆	<i>c</i> ₁ , <i>c</i> ₃ , <i>c</i> ₆
W_2^0		c_1, c_4	<i>c</i> 1		<i>c</i> 1	c 3	c_1, c_3
W_{3}^{0}		c_1, c_6, c_7, c_9	<i>c</i> 1	<i>с</i> 6	c_1, c_6	<i>c</i> ₆ , <i>c</i> ₃	c_1, c_6, c_3
W_1^T		c_1, c_2, c_4, c_7, c_8	c 1		c1	<i>c</i> ₃ , <i>c</i> ₆	<i>c</i> ₁ , <i>c</i> ₃ , <i>c</i> ₆
W_2^T		c_1, c_2, c_4	<i>c</i> 1		<i>c</i> 1	с 3	c_1, c_3
W_3^T		c_4, c_5, c_7, c_8	<i>c</i> 1		<i>c</i> 1	<i>c</i> ₃ , <i>c</i> ₆	c_3, c_1, c_6
W_4^T		c_1, c_2, c_4	<i>c</i> 1		c1	C 3	c_1, c_3
(W_5^T)		<i>c</i> ₁₀					
W_{1}^{0T}		c_1, c_4	<i>c</i> 1		<i>c</i> 1	c 3	c_1, c_3
W_{2}^{0T}		$c_1, c_4, c_5, c_6, c_7, c_8, c_9$	<i>c</i> 1	с 6	<i>c</i> ₁ , <i>c</i> ₆	<i>c</i> ₃ , <i>c</i> ₆	c_1, c_3, c_6
(W_{3}^{0T})		$c_{10}, c_{11}, c_{12}, c_{13}$		<i>c</i> ₁₂	<i>c</i> ₁₂	<i>c</i> ₁₁	<i>c</i> ₁₁
(W_4^{0T})		<i>c</i> ₁₀					

Colors: $W_a = 0$ $W_a c_i$ -indep. $W_a \propto c_i$ $(W_a - \text{const.}) \propto c_i$ Multiple c_i Bold:Contributes to fully integrated CS;Parenthesized:Drops out in CP average

Fig. 3: All operators with dim. 4 or 6 and renormalizable non-flavor universal couplings

relevant for $\bar{B} \longrightarrow D^* l^- \bar{\nu}_l$ (cf. 1506.08896) applied to observables for $l = \tau$.

Summary

$$\mathcal{O}_{\mathfrak{a}} \colon \underset{\mathsf{CS}}{\overset{\mathsf{\Gamma}}{\underset{i} \text{measured}}} \longmapsto \sum_{i} \underset{\mathsf{weights}}{\overset{\omega_{\mathfrak{a}}(i)}{\underset{j}{\mathsf{r}}} \Gamma(i)} \text{ such that } \mathcal{O}_{\mathfrak{a}}(\mathsf{\Gamma}_{\mathsf{theo}}) = \overline{W}_{\mathfrak{a}} \underset{\underset{\mathsf{CS}}{\overset{\mathsf{unknown}}{\underset{j}{\mathsf{NP-dependent}}}}{\overset{\mathsf{unknown}}{\underset{j}{\mathsf{r}}}}$$

Some key points

- Can construct some of the observables even for very coarse binnings (especially relevant for the experimentally challenging $I = \tau$ case)
- The same strategies have been applied to *l* = *e*, μ (both experimentally and theoretically easier, but less suspicious of NP so far)
- Can also consider the same observables with q² dependency (but will need a new figure of merit to optimize against and discriminating between NP models will be more complex)
- To cancel systematics: Consider ratios of observables or use a normalization mode

Summary

- Some simple expression of weights to extract (combinations of the) W_a (on a "proof of concept" base)
- Studied the NP dependency by considering the q^2 distribution of $W_a(q^2)$, considering only one of g_A , g_V , g_S , g_P , T_L to be non-vanishing

New results

- **Characterized** the construction of the observables *O*_a for several assumptions on binning and weights
- Determined the min. number of bins to construct each \mathcal{O}_a
- Introduced a figure of merit for the obtainable sensitivity
- Used degrees of freedom in the weights to optimize sensitivity
- Studied the influence of bin widths on the sensitivity
- Studied the **NP dependency** of *W*_a based on a base of operators *corresponding to relevant mediators* (better physical motivation)

⁶Becirevic, Fajfer, Nisandzic, Tayduganov 1602.03030 (ref. closest to our work to best knowledge)

Backup Slides

Backup slides overview

- General bins
- The angular functions B_a
- Correlation
- Plots $3 \times 3 \times 2$ edges
- Plots $3 \times 3 \times 3$ edges





Fig. 4: 12 bins, $3 \times 3 \times 3$ cubes (equidistant)



Fig. 4: 12 bins, $5 \times 5 \times 5$ cubes



Fig. 4: 27 bins, $5 \times 5 \times 5$ cubes

The angular functions B_a (I)

$$B_1^T = \sin^3(\theta_{D^*})\sin(\theta_{\ell})$$

$$B_1^0 = \sin(\theta_{D^*})\cos^2(\theta_{D^*})\sin(\theta_{\ell})$$

$$B_2^0 = \sin(\theta_{D^*})\cos^2(\theta_{D^*})\sin^3(\theta_{\ell})$$

$$B_3^0 = \sin(\theta_{D^*})\cos^2(\theta_{D^*})\sin(\theta_{\ell})\cos(\theta_{\ell})$$

$$B_4^T = \cos(2\chi)\sin^3(\theta_{D^*})\sin^3(\theta_{\ell})$$

$$B_5^T = \sin(2\chi)\sin^3(\theta_{D^*})\sin^3(\theta_{\ell})$$

$$B_1^{0T} = \cos(\chi) \sin(\theta_{D^*}) \sin(2\theta_{D^*}) \sin(\theta_{\ell}) \sin(2\theta_{\ell})$$

$$B_2^{0T} = \cos(\chi) \sin(\theta_{D^*}) \sin(2\theta_{D^*}) \sin^2(\theta_{\ell})$$

$$B_3^{0T} = \sin(\chi) \sin(\theta_{D^*}) \sin(2\theta_{D^*}) \sin^2(\theta_{\ell})$$

$$B_4^{0T} = \sin(\chi) \sin(\theta_{D^*}) \sin(2\theta_{D^*}) \sin(\theta_{\ell}) \sin(2\theta_{\ell})$$

Table 1: The functions $B_a(\chi, \theta_\ell, \theta_{D^*})$

The angular functions B_a (II)



Fig. 5: Combinations of the angle functions.

$$\begin{split} f_1(\chi) &:= 1, \\ f_2(\chi) &:= \sin(\chi), \\ f_3(\chi) &:= \cos(\chi), \\ f_4(\chi) &:= \sin(2\chi), \\ f_5(\chi) &:= \cos(2\chi), \end{split}$$

$$g_1(\theta_\ell) := \sin(\theta_l),$$

$$g_2(\theta_\ell) := \sin^2(\theta_l),$$

$$g_3(\theta_\ell) := \sin^3(\theta_l),$$

$$g_4(\theta_\ell) := \sin(\theta_l)\cos(\theta_l),$$

$$g_5(\theta_\ell) := \sin(\theta_l)\sin(2\theta_l),$$

$$h_1(\theta_{D^*}) := \sin^3(\theta_{D^*}),$$

$$h_2(\theta_{D^*}) := \sin(\theta_{D^*})\sin(2\theta_{D^*}),$$

$$h_3(\theta_{D^*}) := \sin(\theta_{D^*})\cos^2(\theta_{D^*}).$$

Correlation



Fig. 6: Correlation of angle functions resp. observables

$$\{-\pi, \chi_1, \chi_2, \pi\}_{\chi} imes \{\mathbf{0}, \theta_{\ell 1}, \theta_{\ell 2}, \pi\}_{ heta_{\ell}} imes \{\mathbf{0}, \Delta \theta_{\mathsf{D}^*}, \pi\}_{ heta_{D^*}}$$



$$\{-\pi, \chi_1, \chi_2, \pi\}_{\chi} imes \{\mathbf{0}, heta_{\ell 1}, heta_{\ell 2}, \pi\}_{ heta_{\ell}} imes \{\mathbf{0}, \mathbf{\Delta} heta_{\mathsf{D}^*}, \pi\}_{ heta_{D^*}}$$



 $\{-\pi, \chi_1, \chi_2, \pi\}_{\chi} \times \{\mathbf{0}, \theta_{\ell 1}, \theta_{\ell 2}, \pi\}_{\theta_{\ell}} \times \{\mathbf{0}, \Delta \theta_{D^*}, \pi\}_{\theta_{D^*}}$



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 $\{-\pi,\chi_1,\chi_2,\pi\}_{\chi}\times\{\mathbf{0},\theta_{\ell 1},\theta_{\ell 2},\pi\}_{\theta_{\ell}}\times\{\mathbf{0},\theta_{D^*1},\theta_{D^*2},\pi\}_{\theta_{D^*}}$



 $\{-\pi, \chi_1, \chi_2, \pi\}_{\chi} \times \{\mathbf{0}, \theta_{\ell 1}, \theta_{\ell 2}, \pi\}_{\theta_{\ell}} \times \{\mathbf{0}, \theta_{D^{*}1}, \theta_{D^{*}2}, \pi\}_{\theta_{D^{*}}}$



 $\{-\pi, \chi_1, \chi_2, \pi\}_{\chi} \times \{\mathbf{0}, \theta_{\ell 1}, \theta_{\ell 2}, \pi\}_{\theta_{\ell}} \times \{\mathbf{0}, \theta_{D^*1}, \theta_{D^*2}, \pi\}_{\theta_{D^*}}$



 $\{-\pi, \chi_1, \chi_2, \pi\}_{\chi} \times \{\mathbf{0}, \theta_{\ell 1}, \theta_{\ell 2}, \pi\}_{\theta_{\ell}} \times \{\mathbf{0}, \theta_{D^*1}, \theta_{D^*2}, \pi\}_{\theta_{D^*}}$



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