

# Modelling non-perturbative corrections in semi-inclusive $B$ decays

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**The work presented in this talk is based on the following papers:**

- U. Aglietti, G. Ferrera and G. Ricciardi, Phys. Rev. D74 (2006), [hep-ph/0507285].*
- U. Aglietti, G. Ferrera and G. Ricciardi, Phys. Rev. D74 (2006), [hep-ph/0509095].*
- U. Aglietti, G. Ferrera and G. Ricciardi, Phys. Rev. D74 (2006), [hep-ph/0509271].*
- U. Aglietti, G. Ferrera and G. Ricciardi, [hep-ph/0608047].*
- U. Aglietti, G. Corcella and G. Ferrera, [hep-ph/0610035].*
- U. Aglietti, L. Di Giustino, G. Ferrera and L. Trentadue, [hep-ph/0612073].*



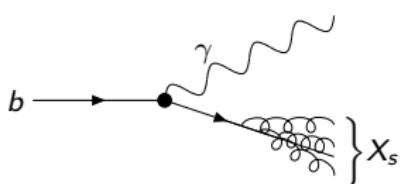
# Outline

- 1 Semi-inclusive  $B$  decays
- 2 Analytic QCD coupling
- 3 Phenomenological Analysis
- 4 Conclusions and Perspectives

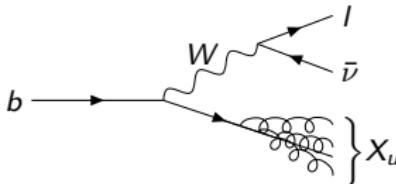


# Semi-inclusive $B$ decays

$$b \rightarrow X_q + \langle \text{non QCD partons} \rangle, \quad q \equiv u, s, d$$



Radiative decay:  
 $b \rightarrow X_s + \gamma$



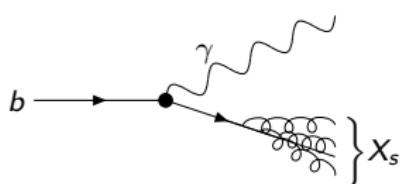
Semileptonic charmless decay:  
 $b \rightarrow X_u + l + \bar{\nu}$

- Energy scales:  $m_b \geq E_X \geq m_X$ ,  $Q = E_X + \sqrt{E_X^2 - m_X^2}$
- Threshold region:  $m_X \ll E_X$ ,  $Q = 2E_X$

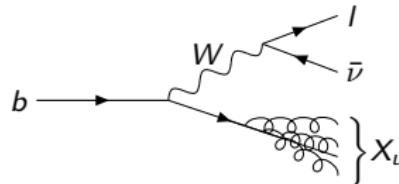


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- The LO probability for a light quark to evolve into a jet with an invariant mass smaller than  $m_X$  is:

$$\begin{aligned} P_<(m_X) &= 1 + \alpha_S \frac{C_F}{\pi} \int_0^1 \frac{d\omega}{\omega} \int_0^1 \frac{d\theta^2}{\theta^2} \left[ \Theta\left(\frac{m_X^2}{Q^2} - \omega\theta^2\right) - 1 \right] + O(\alpha_S^2) \\ &= 1 - \alpha_S \frac{C_F}{2\pi} \ln^2\left(\frac{Q^2}{m_X^2}\right) + O(\alpha_S^2). \end{aligned}$$

- Multiple gluon emission gives rise to the double logarithmic series, i.e. to powers of the term  $\alpha_S \ln^2(Q^2/m_X^2)$ .

$$\sigma(y) = \sum_{n=1}^{\infty} \sum_{k=1}^{2n} c_{n,k} \alpha_S^n(Q) \ln^k y, \quad y = \frac{m_X^2}{Q^2}.$$



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# Radiative and Semileptonic charmless decays: factorization

- Resummation formula for radiative decays [Aglietti ('01)]:

$$\frac{1}{\Gamma_r} \frac{d\Gamma_r}{dt} = C_r[\alpha_S(Q)] \sigma[t; \alpha_S(Q)] + d_r[t; \alpha_S(Q)] ,$$

where  $t \equiv m_X^2/m_b^2$  and  $Q = 2E_X$ .

$$Q = m_b(1 + m_X^2/m_b^2) \simeq m_b, \quad \alpha_S(Q) \simeq \alpha_S(m_b) \simeq 0.22 .$$

- Resummation formula for semileptonic charmless decays [Aglietti ('01)]:

$$\frac{1}{\Gamma_s} \frac{d^3\Gamma_s}{dx du dw} = C_s[x, w; \alpha_S(Q)] \sigma[u; \alpha_S(Q)] + d_s[x, u, w; \alpha_S(Q)]$$

where  $x \equiv \frac{2E_l}{m_b}$ ,  $w \equiv \frac{2E_X}{m_b}$ ,  $u \equiv \frac{1 - \sqrt{1 - 4y}}{1 + \sqrt{1 - 4y}}$ ,  $y \equiv \frac{m_X^2}{4E_X^2}$ .

$$Q = m_b(1 + m_X^2/m_b^2 - q^2/m_b^2); \quad q^2 \text{ is the dilepton invariant mass.}$$

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- In the Mellin space the threshold resummed form factor reads [Sterman ('87), Catani & Trentadue ('89)]:

$$\begin{aligned} \ln \sigma_N &= \int_0^1 dy \frac{(1-y)^{N-1} - 1}{y} \left\{ \int_{Q^2 y^2}^{Q^2 y} \frac{dk_\perp^2}{k_\perp^2} A[\alpha_S(k_\perp^2)] + B[\alpha_S(Q^2 y)] + D[\alpha_S(Q^2 y^2)] \right\} \\ &= \sum_{n=1}^{\infty} \sum_{m=1}^{n+1} G_{nm} \alpha_S^n(Q) \ln^m N, \quad A(\alpha_S) = \sum_{n=1}^{\infty} A_n \alpha_S^n, \quad B(\alpha_S) = \dots, \quad D(\alpha_S) = \dots . \end{aligned}$$

- The new expansion parameter is  $\alpha_S L$  (where  $L \equiv \ln N$ ):
  - LL:  $\alpha_S^n L^{n+1}$ ;
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  - ... and so on.
- When  $y = \frac{m_X^2}{Q^2} \rightarrow 0$  the integration in  $k_\perp^2$  involves  $\alpha_S$  at the Landau pole: it is necessary a prescription, e. g. the Minimal Prescription [Catani, Mangano, Nason & Trentadue ('96)].
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# Analytic QCD coupling

- Standard QCD coupling: physical cut at  $\mu^2 < 0$  and unphysical pole at  $\mu^2 = \Lambda_{QCD}^2$ :

$$\alpha_S^{lo}(\mu^2) = \frac{1}{\beta_0 \ln \frac{\mu^2}{\Lambda_{QCD}^2}}.$$

- Analytic QCD coupling: same discontinuity along the cut but analytic elsewhere in the complex plane [Shirkov & Solovtsov ('97)]:

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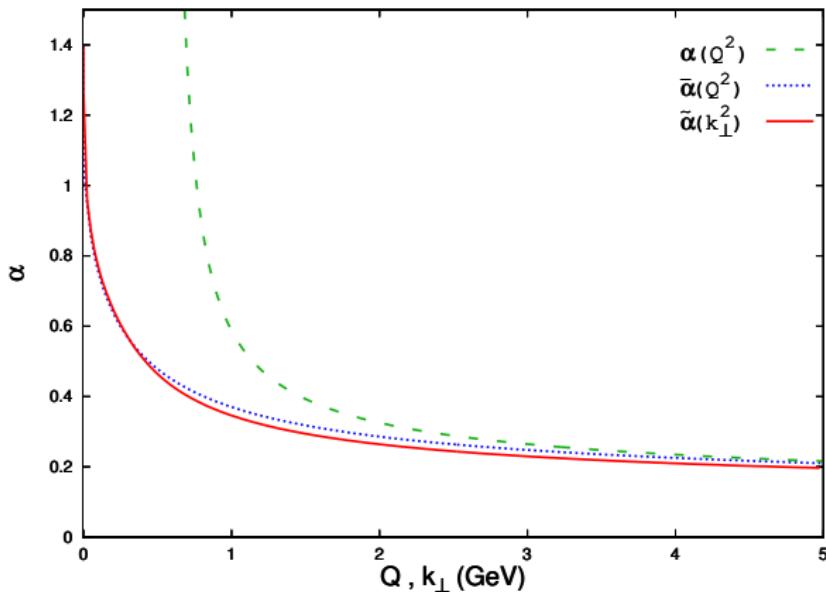


Figure 1: Time-like and space-like analytic couplings compared with the standard one.



# Phenomenological Analysis

$b$ -quark fragmentation:  $e^+e^- \rightarrow Z^0 \rightarrow B + X$ ,  $x_b = \frac{2E_b}{m_Z}$

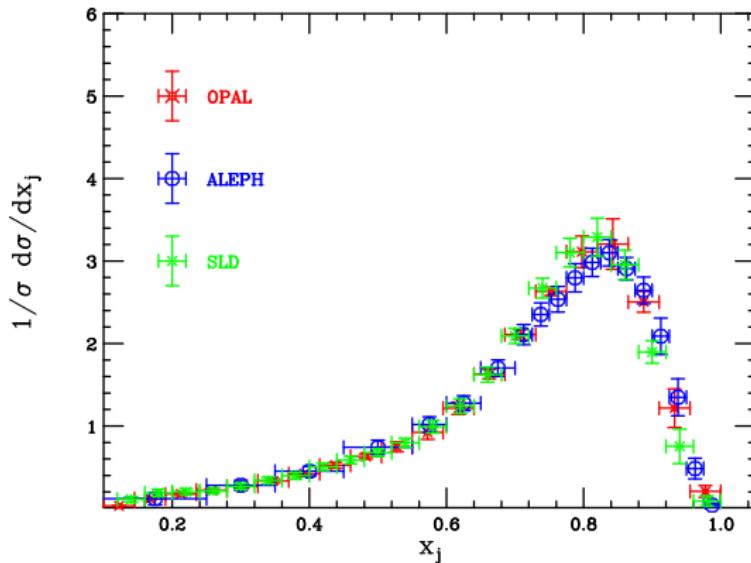


Figure 2:  $B$ -hadron spectrum in  $e^+e^-$  annihilation at  $Z^0$  peak: precision of the model compared with experimental data [Aleph ('01), Delphi ('02), SLD ('00)].



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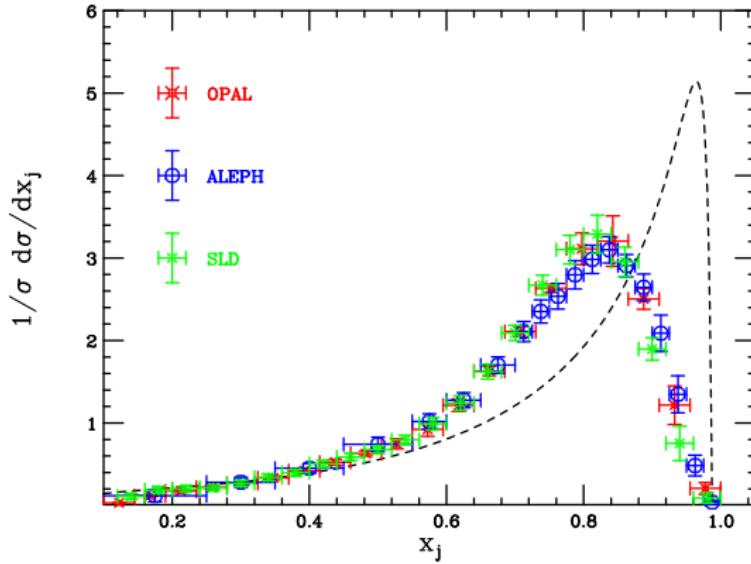


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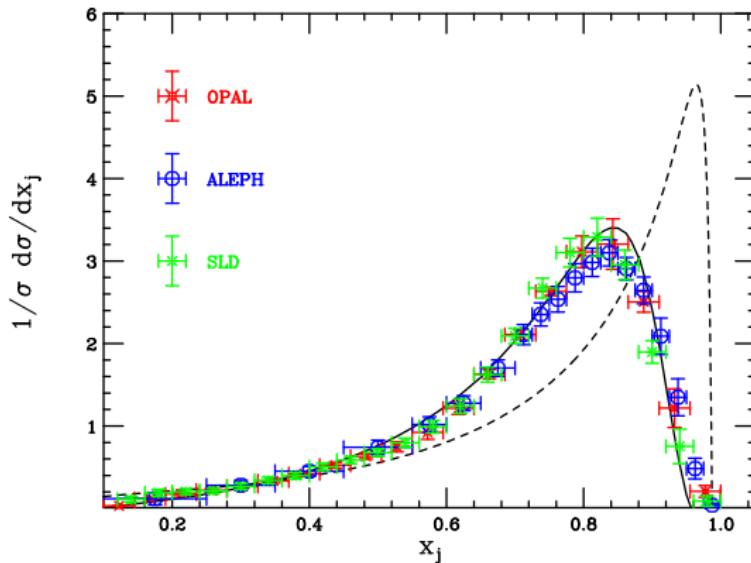


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## Radiative decay: hadron mass distribution

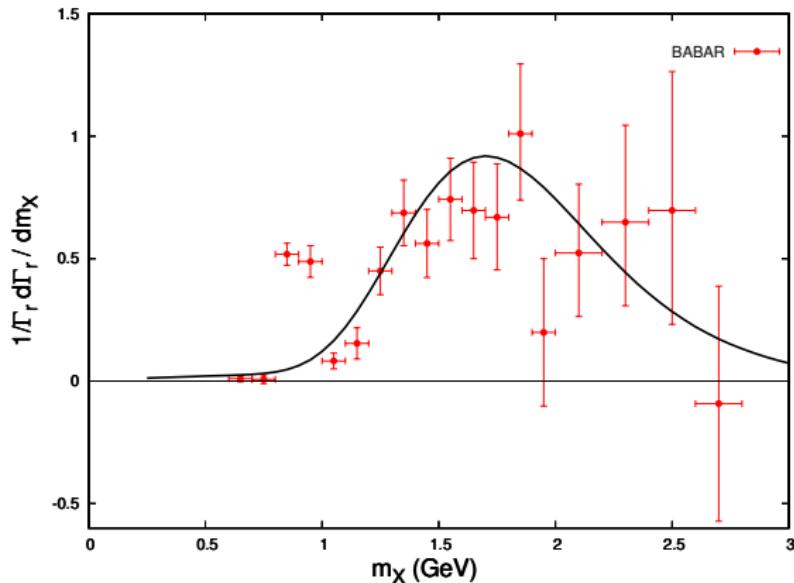


Figure 3: Invariant hadron mass distribution in the radiative decay: prevision of the model compared with the experimental data [BaBar ('05)]. The  $K^*$  peak cannot clearly be accounted for in a perturbative QCD framework.



## Radiative decay: photon energy distribution

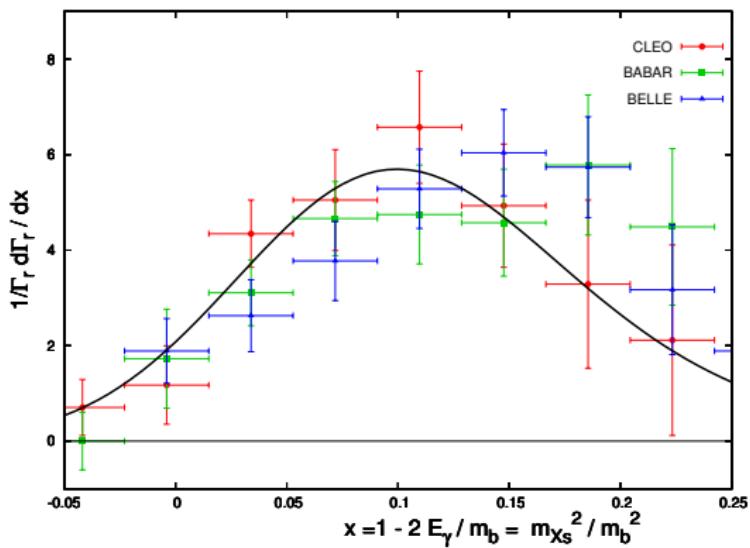


Figure 4: Photon energy spectrum in the radiative decay: prevision of the model compared with data [Cleo ('01), BaBar ('05), Belle ('05)]. To model the Doppler effect related to the motion of the  $B$  mesons , we have convoluted the theoretical curve with a Gaussian distribution with  $\sigma \sim 180$  MeV.

## Semileptonic decay: hadron mass distribution

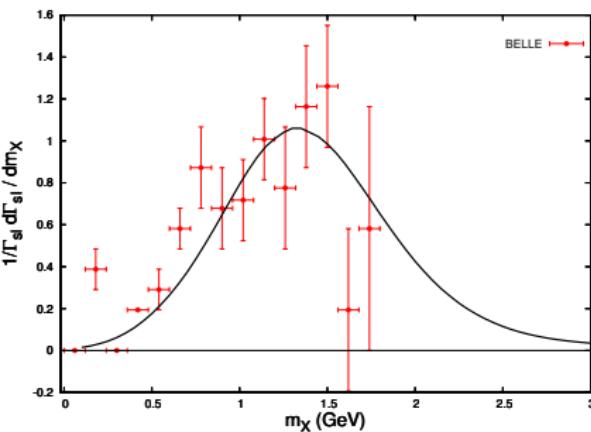
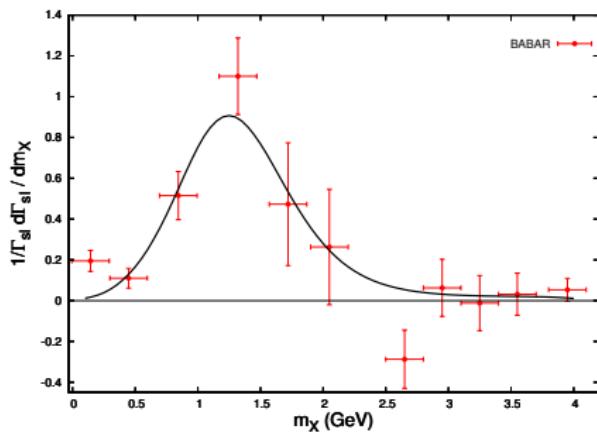


Figure 5: Invariant hadron mass distribution in the semileptonic decay: precision of the model compared with the experimental [Belle ('04), BaBar ('05)]. Note the  $\pi$  and the  $\rho$  peaks at small hadron masses.



## Semileptonic decay: electron energy distribution

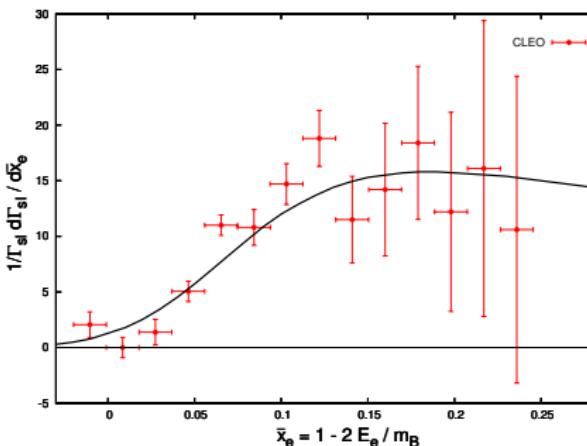


Figure 6: Inclusive charmless electron spectrum in the semileptonic decay: prediction of the model compared with data [Cleo ('01)]. Doppler effect included convoluting with a Gaussian with  $\sigma \sim 100$  MeV.



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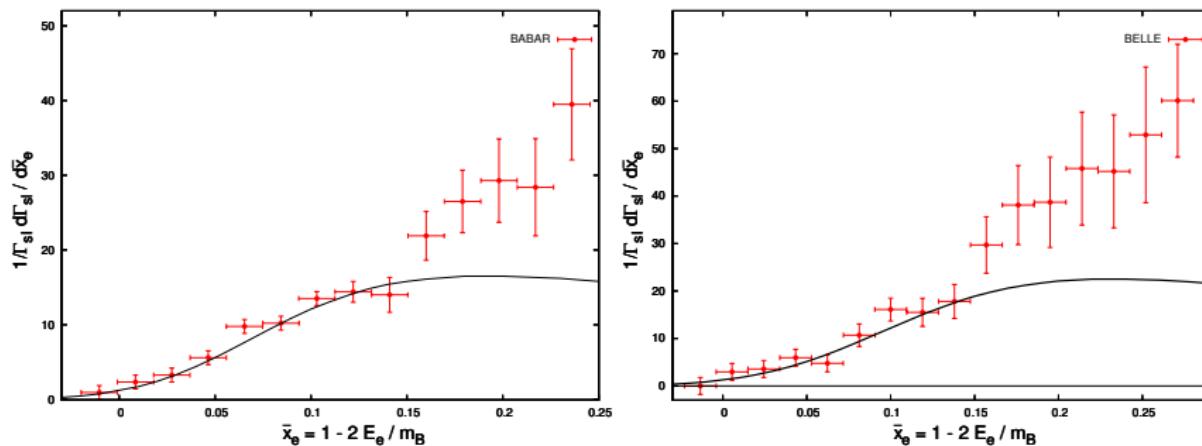


Figure 7: Inclusive charmless electron spectrum in the semileptonic decay: prevision of the model compared with data [Babar ('05) and Belle ('04)].

Doppler effect included convoluting with a Gaussian with  $\sigma \sim 100$  MeV. We do not know whether the discrepancy is related to a deficiency of our model or to an under-estimate charm background.



# Extraction of $\alpha_S(m_Z)$ and $V_{ub}$ from the data

$b \rightarrow s \gamma$

$$\begin{aligned}\alpha_S(m_Z) &= 0.117 \pm 0.004 & (E_\gamma : CLEO, \sigma_\gamma = 150 \text{ MeV}) \\ \alpha_S(m_Z) &= 0.129 \pm 0.005 & (E_\gamma : BABAR, \sigma_\gamma = 200 \text{ MeV}) \\ \alpha_S(m_Z) &= 0.130 \pm 0.005 & (E_\gamma : BELLE, \sigma_\gamma = 200 \text{ MeV})\end{aligned}$$

$b \rightarrow u / \nu$

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# A possible new measure

## Hadronic energy distribution in the semileptonic decay

- The QCD form factor can be experimentally measured from the  $m_X$  or the  $E_\gamma$  distribution of the radiative decay:

$$\frac{1}{\Gamma_r} \frac{d\Gamma_r}{dt} = C_r(\alpha_S) \sigma[t; \alpha_S(m_b)] + d_r(t; \alpha_S),$$

unfortunately the data are not sufficiently accurate to do this.

- The only single differential distribution in the semileptonic decay which permits the direct extraction of the QCD form factor is the hadronic energy distribution for  $w \equiv 2E_X/m_b > 1$ :

$$\frac{1}{\Gamma_s} \frac{d\Gamma_s}{dw} = C_{w_1}(\alpha_S) \left\{ 1 - C_{w_2}(\alpha_S) \Sigma[w-1; \alpha_S(m_b)] + H(w; \alpha_S) \right\} \quad (w > 1)$$

where

$$\Sigma[u; \alpha_S] = \int_0^u du' \sigma(u'; \alpha_S).$$



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$$\frac{1}{\Gamma_r} \frac{d\Gamma_r}{dt} = C_r(\alpha_S) \sigma[t; \alpha_S(m_b)] + d_r(t; \alpha_S),$$

unfortunately the data are not sufficiently accurate to do this.

- The only single differential distribution in the semileptonic decay which permits the direct extraction of the QCD form factor is the hadronic energy distribution for  $w \equiv 2E_X/m_b > 1$  :

$$\frac{1}{\Gamma_s} \frac{d\Gamma_s}{dw} = C_{w_1}(\alpha_S) \left\{ 1 - C_{w_2}(\alpha_S) \Sigma[w-1; \alpha_S(m_b)] + H(w; \alpha_S) \right\} \quad (w > 1)$$

where

$$\Sigma[u; \alpha_S] = \int_0^u du' \sigma(u'; \alpha_S).$$



## Semileptonic decay: hadronic energy distribution

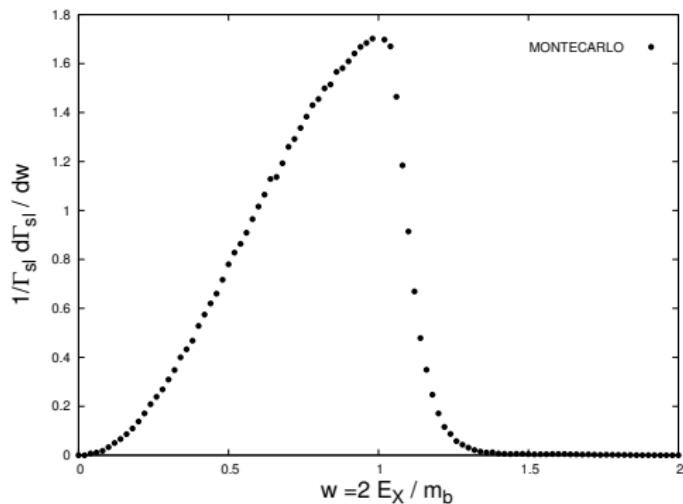


Figure 8: Hadronic energy distribution in the semileptonic decay generated by a montecarlo based on the model.



## Semi-leptonic charmed decay and $V_{cb}$

- To describe the process  $B \rightarrow X_c / \nu$  we need a new formalism to take in account the non-vanishing charm mass  $m_c$ : [Aglietti, Di Giustino, G.F. & Trentadue ('06)]

$$\sigma_N(Q^2, m^2) = \sigma_N(Q^2) \delta_N(Q^2, m^2), \quad r \equiv \frac{m^2}{Q^2} \simeq 0.1$$

$$\ln \delta_N = \int_0^1 dy \frac{(1-y)^{r(N-1)} - 1}{y} \left\{ - \int_{m^2 y^2}^{m^2 y} \frac{dk_\perp^2}{k_\perp^2} A[\alpha_S(k_\perp^2)] - B[\alpha_S(m^2 y)] + D[\alpha_S(m^2 y^2)] \right\}$$

- Using this formula, the full  $O(\alpha_S)$  triple differential distribution [Trott ('04), Aquila, Gambino, Ridolfi & Uraltsev ('05)] and the model previously described we are confident that we can provide quantitative description of the data.



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# Conclusions and Perspectives

- Through the analytic QCD coupling and NNLL threshold resummation we have developed a model that includes the large non-perturbative effects in the semi-inclusive  $B$  decays: it describes with good accuracy the measured spectra without introducing any ad hoc non-perturbative component.
- We have a disagreement with the electron spectrum in the semileptonic decay (BaBar and Belle). It could be a deficiency of our model or an underestimate of the charm background.
- It would be interesting if experimentalist made a statistical analysis of the compatibility of our model with the data to extract  $V_{ub}$ .
- We propose to measure the  $E_X$  spectrum of the semileptonic charmless decay: it gives direct information on the QCD form factor.
- Using a new resummation formalism recently developed together with available fixed order calculation we will carry out a phenomenological analysis of semileptonic charmed decays soon.



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# Back-up Slides



## Semileptonic decay: electron energy distribution

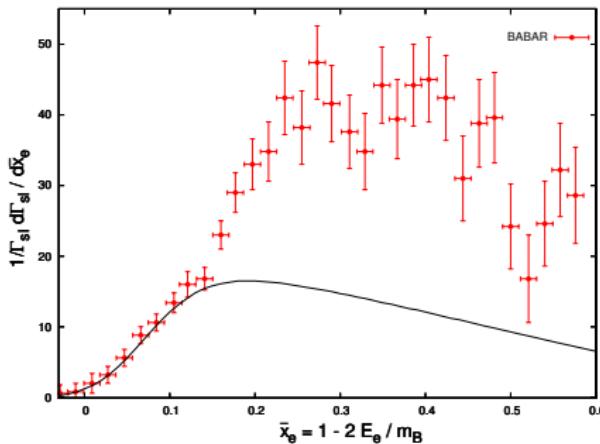
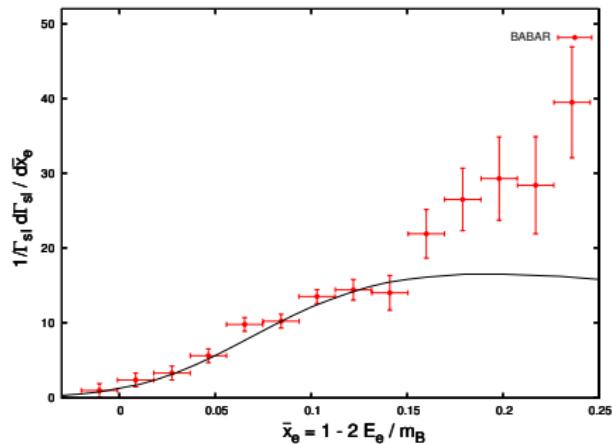


Figure 7: Inclusive charmless electron spectrum in the semileptonic decay: prevision of the model compared with data [Babar ('05)]. Doppler effect included convoluting with a Gaussian with  $\sigma \sim 100$  MeV. We do not know whether the discrepancy is related to a deficiency of our model or to an under-estimate charm background.



# Radiative and Semileptonic charmless decays

- Resummation formula for radiative decays [Aglietti ('01)]:

$$\frac{1}{\Gamma_r} \frac{d\Gamma_r}{dt} = C_r[\alpha_S(Q)] \sigma[t; \alpha_S(Q)] + d_r[t; \alpha_S(Q)],$$

where  $t \equiv m_X^2/m_b^2$  and  $Q = 2E_X$ .

$Q = m_b(1 + m_X^2/m_b^2) \simeq m_b$ ,  $\alpha_S(Q) \simeq \alpha_S(m_b) \simeq 0.22$ .

- Resummation formula for semileptonic charmless decays [Aglietti ('01)]:

$$\frac{1}{\Gamma_s} \frac{d^3\Gamma_s}{dx du dw} = C_s[x, w; \alpha_S(Q)] \sigma[u; \alpha_S(Q)] + d_s[x, u, w; \alpha_S(Q)]$$

where  $x \equiv \frac{2E_l}{m_b}$ ,  $w \equiv \frac{2E_X}{m_b}$ ,  $u \equiv \frac{1-\sqrt{1-4y}}{1+\sqrt{1-4y}}$ ,  $y \equiv \frac{m_X^2}{4E_X^2}$ .

$Q = m_b(1 + m_X^2/m_b^2 - q^2/m_b^2)$ ;  $q^2$  is the dilepton invariant mass.

We can not put  $\alpha_S(Q) \simeq \alpha_S(m_b)$  in the form factor  $\sigma$ :

$$\alpha_S(Q) = \alpha_S(wm_b).$$



## Analytic QCD coupling

- Standard QCD coupling: physical cut at  $\mu^2 < 0$  and unphysical pole at  $\mu^2 = \Lambda_{QCD}^2$ :

$$\alpha_s^{lo}(\mu^2) = \frac{1}{\beta_0 \ln \frac{\mu^2}{\Lambda_{QCD}^2}} .$$

- Analytic QCD coupling: same discontinuity along the cut but analytic elsewhere in the complex plane [Shirkov & Solovtsov ('97)]

$$\bar{\alpha}_S(Q^2) = \frac{1}{2\pi i} \int_0^\infty \frac{ds}{s+Q^2} Disc_s \alpha_S(-s), \quad \text{space-like.}$$

- The infrared pole is subtracted without modify high energy behaviour

$$\bar{\alpha}_S^{lo}(Q^2) = \frac{1}{\beta_0} \left[ \frac{1}{\ln Q^2/\Lambda_{QCD}^2} - \frac{\Lambda_{QCD}^2}{Q^2 - \Lambda_{QCD}^2} \right],$$

$$\lim_{Q^2 \rightarrow 0} \bar{\alpha}_S(Q^2) = \frac{1}{\beta_0} , \quad \lim_{Q^2 \rightarrow \infty} \bar{\alpha}_S(Q^2) = \lim_{Q^2 \rightarrow \infty} \alpha_S(Q^2) .$$



- Semi-inclusive  $B$  decays are time-like processes:

$$\tilde{\alpha}_S(k_\perp^2) = \frac{i}{2\pi} \int_0^{k_\perp^2} ds \text{Disc}_s \frac{\bar{\alpha}_S(-s)}{s}, \quad \text{time-like.}$$

- At leading order we have:

$$\tilde{\alpha}_S^{lo}(k_\perp^2) = \frac{1}{\beta_0} \left( \frac{1}{2} - \frac{1}{\pi} \arctan \frac{\ln \frac{k_\perp^2}{\Lambda_{QCD}^2}}{\pi} \right),$$

$$\lim_{k_\perp^2 \rightarrow 0} \tilde{\alpha}_S(k_\perp^2) = \frac{1}{\beta_0}, \quad \lim_{k_\perp^2 \rightarrow \infty} \tilde{\alpha}_S(k_\perp^2) = \lim_{k_\perp^2 \rightarrow \infty} \alpha_S(k_\perp^2).$$

- The well defined quantity

$$\alpha_0 = \frac{1}{\mu_I} \int_0^{\mu_I} \tilde{\alpha}_S(k_\perp^2) dk_\perp \simeq 0.44,$$

is similar to the fitted value from shape variables data in the DMW model ( $\alpha_0 \simeq 0.45$ ) [Dokshitzer, Marchesini & Webber ('95)].



# References

- [1] S. Catani and L. Trentadue, Nucl. Phys. B **327** (1989) 323.
- [2] G. Sterman, Nucl. Phys. B **281** (1987) 310.
- [3] U. Aglietti, R. Sghedoni and L. Trentadue, [arXiv:hep-ph/0105322].
- [4] U. Aglietti, [arXiv:hep-ph/0104020, hep-ph/0601242].
- [5] U. Aglietti, G. Ricciardi and G. Ferrera, [arXiv:hep-ph/0507285, hep-ph/0509095, hep-ph/0509271].
- [6] S. Catani, M. L. Mangano, P. Nason and L. Trentadue, [arXiv:hep-ph/9602208].
- [7] U. Aglietti, G. Corcella and G. Ferrera, [arXiv:hep-ph/0610035].
- [8] D. V. Shirkov and I. L. Solovtsov, [arXiv:hep-ph/9704333].
- [9] U. Aglietti and G. Ricciardi, [arXiv:hep-ph/0407225].
- [10] Y. L. Dokshitzer, G. Marchesini and B. R. Webber, [arXiv:hep-ph/9512336].
- [11] BaBar Collaboration, B. Aubert *et al.*, [arXiv:hep-ex/0408068, hep-ex/0507001, hep-ex/0508004, hep-ex/0509040 ].
- [12] Belle Collaboration, P. Koppenburg *et al.*, [arXiv:hep-ex/0403004, hep-ex/0504046, hep-ex/0505088].
- [13] CLEO Collaboration, S. Chen *et al.*, [arXiv:hep-ex/0108032].
- [14] B. Mele and P. Nason, Nucl. Phys. B **361** (1991) 626.
- [15] L.N. Lipatov, Sov. J. Nucl. Phys. **20** (1975) 95; V.N. Gribov and L.N. Lipatov, Sov. J. Nucl. Phys. **15** (1972) 438; Yu.L. Dokshitzer, Sov. Phys. **46** (1977) 641; G. Altarelli and G. Parisi, Nucl. Phys. B **126** (1977) 298.
- [16] SLD Collaboration, K. Abe *et al.*, Phys. Rev. Lett. **84** (2000) 4300.
- [17] ALEPH Collaboration, A. Heister *et al.*, Phys. Lett. B **512** (2001) 30.
- [18] OPAL Collaboration, G. Abbiendi *et al.*, Eur. Phys. J. C **29** (2003) 463.



# Threshold Resummation

In the Mellin space the threshold resummed form factor reads  
 [Sterman ('87), Catani & Trentadue ('89)]:

$$\ln \sigma_N = \int_0^1 dy \frac{(1-y)^{N-1} - 1}{y} \left\{ \int_{Q^2 y^2}^{Q^2 y} \frac{dk_\perp^2}{k_\perp^2} A[\alpha_S(k_\perp^2)] + B[\alpha_S(Q^2 y)] + D[\alpha_S(Q^2 y^2)] \right\}$$

$$= \sum_{n=1}^{\infty} \sum_{m=1}^{n+1} G_{nm} \alpha_S^n(Q) L^m = L g_1(\lambda) + g_2(\lambda) + \alpha_S g_3(\lambda) + \dots$$

where  $A(\alpha_S) = \sum_{n=1}^{\infty} A_n \alpha_S^n$ ,  $B(\alpha_S) = \sum_{n=1}^{\infty} B_n \alpha_S^n$ ,  $D(\alpha_S) = \sum_{n=1}^{\infty} D_n \alpha_S^n$ .

and  $g_i(\lambda) = \sum_{n=0}^{\infty} g_{i,n} \lambda^n$ ,  $\lambda \equiv \beta_0 \alpha_S L$ ,  $L \equiv \ln N$



# Radiative $B$ decay: $B \rightarrow X_s \gamma$

- The resummation formula for the invariant mass distribution reads [Aglietti ('01)]:

$$\frac{1}{\Gamma_r} \frac{d\Gamma_r}{dt} = C_r[\alpha_S(Q)] \sigma[t; \alpha_S(Q)] + d_r[t; \alpha_S(Q)] ,$$

where  $t \equiv m_X^2/m_b^2$  and  $Q = 2E_X$ .

$$Q = m_b(1 + m_X^2/m_b^2) \simeq m_b, \quad \alpha_S(Q) \simeq \alpha_S(m_b) \simeq 0.22 .$$

- $C_r(\alpha_S) = C_r^{(0)} + \alpha_S C_r^{(1)} + \dots$   
short-distance (process dependent) hard factor.
- $\Sigma(t; \alpha_S) = \int_0^t \sigma(t'; \alpha_S) dt' = \sum_{n=0}^{\infty} \sum_{k=0}^{2n} \Sigma_{n,k} \alpha_S^n \ln^k(1/t)$   
long-distance dominated (universal) QCD form factor.
- $d_r(t; \alpha_S) = d_r^{(0)}(t) + \alpha_S d_r^{(1)}(t) + \dots$   
short-distance (process dependent) remainder function, to have good approximation also in the region  $m_X \leq E_X$ :  $\lim_{t \rightarrow 0} \int_0^t d_r(t'; \alpha_S) dt' = 0$ .



# Semileptonic $B$ decay: $B \rightarrow X_u \mid \nu$

- The resummation formula for the triple differential distribution reads [Aglietti ('01)]:

$$\frac{1}{\Gamma_s} \frac{d^3 \Gamma_s}{dx du dw} = C_s[x, w; \alpha_S(Q)] \sigma[u; \alpha_S(Q)] + d_s[x, u, w; \alpha_S(Q)]$$

where  $x \equiv \frac{2E_L}{m_b}$ ,  $w \equiv \frac{2E_X}{m_b}$ ,  $u \equiv \frac{1-\sqrt{1-4y}}{1+\sqrt{1-4y}}$ ,  $y \equiv \frac{m_X^2}{4E_X^2}$  and,

$Q = m_b(1 + m_X^2/m_b^2 - q^2/m_b^2)$ ;  $q^2$  is the dilepton invariant mass. In this case we can **not** put  $\alpha_S(Q) \simeq \alpha_S(m_b)$  in the resummed form factor where  $\alpha_S(Q) = \alpha_S(w m_b)$ .

- $C_s(x, w; \alpha_S) = C_s^{(0)}(x, w) + \alpha_S C_s^{(1)}(x, w) + \dots$   
short-distance (process dependent) hard factor.
- $d_s(x, u, w; \alpha_S) = d_s^{(0)}(x, u, w) + \alpha_S d_s^{(1)}(x, u, w) + \dots$   
short-distance (process dependent) remainder function.



# Non universality effects

- Universality of long-distance effects studied by comparing the logarithmic structure of different spectra.
- Spectra not involving integration over hadron energy: same infrared structure of the hadron invariant mass distribution of the radiative decay i.e **pure short-distance relation** [Aglietti, Ricciardi & G.F. ('05)].

$$\Sigma(u; \alpha_S) = \int_0^u \sigma(u'; \alpha_S) du' = \sum_{n=0}^{\infty} \sum_{k=0}^{2n} \Sigma_{n,k} \alpha_S^n \ln^k \frac{1}{u}$$

- Spectra involving integration over hadron energy: different infrared structure from each other and from the hadron invariant mass distribution of the radiative decay i.e **not pure short-distance relation** [Aglietti, Ricciardi & G.F. ('05)].

$$\Sigma_U(u; \alpha_S) = \frac{\int_0^1 \int_0^w C(x, w; \alpha_S) dx \Sigma(u; \alpha_S(wm_b)) dw}{\int_0^1 \int_0^w C(x, w; \alpha_S) dx dw} = \sum_{n=0}^{\infty} \sum_{k=0}^{2n} \Sigma_{U,n,k} \alpha_S^n \ln^k \frac{1}{u}$$



## Model for soft gluon effects

- Non perturbative effects represent substantial contributions to the spectra that cannot be neglected both in semi-inclusive  $B$  decay and in  $b$ -quark fragmentation.
- The usual approach is to factorize non perturbative contributions in a universal distribution, depending on free parameters fitted to experimental data, and convolute it with perturbative spectra.
- By the use of a regular low energy QCD coupling our model includes non perturbative effects. Without introducing an additional ad hoc non perturbative component we are able to compare our predictions directly with data [Aglietti, G.F. & Ricciardi ('06)].



## Comparison with DMW model

- Since the time-like coupling is regular for any value of  $k_\perp$ , we can compute the average of the coupling

$$\alpha_0 = \frac{1}{\mu_I} \int_0^{\mu_I} \tilde{\alpha}_S(k_\perp^2) dk_\perp,$$

which is a free parameter to be determined with a fit to experimental data [Dokshitzer, Marchesini & Webber ('95)].

- Assuming  $\alpha_S(m_b) = 0.22$ ,  $n_f = 3$  and  $\mu_I = 2$  GeV, we obtain at leading order with the time-like coupling:

$$\alpha_0 \approx 0.44.$$

- The fitted value from shape variable in  $e^+e^-$  data is around 0.45, not distant from our estimate.



## Improved threshold resummation

- The improved threshold resummation formula therefore reads

$$\ln \sigma_N = \int_0^1 dy \frac{(1-y)^{N-1} - 1}{y} \left\{ \int_{Q^2 y^2}^{Q^2 y} \frac{dk_\perp^2}{k_\perp^2} \tilde{A}[\tilde{\alpha}_S(k_\perp^2)] + \tilde{B}[\tilde{\alpha}_S(Q^2 y)] + \tilde{D}[\tilde{\alpha}_S(Q^2 y^2)] \right\}$$

- The coefficients for the time-like coupling are obtained by imposing the equality:

$$A(\alpha_S) = \tilde{A}(\tilde{\alpha}_S),$$

where

$$\tilde{A}(\tilde{\alpha}_S) = \sum_{n=1}^{\infty} \tilde{A}_n \tilde{\alpha}_S^n = \tilde{A}_1 \tilde{\alpha}_S + \tilde{A}_2 \tilde{\alpha}_S^2 + \tilde{A}_3 \tilde{\alpha}_S^3 + \dots$$

- Expressing the time-like coupling in terms of the standard one, we obtain:

$$\tilde{A}_1 = A_1; \quad \tilde{A}_2 = A_2; \quad \tilde{A}_3 = A_3 + \frac{(\pi \beta_0)^2}{3} A_1 \simeq 0.31 + 0.72 \simeq 1;$$

analogous relations hold for  $\tilde{B}_i$  and  $\tilde{D}_i$ .



# Phenomenological Analysis

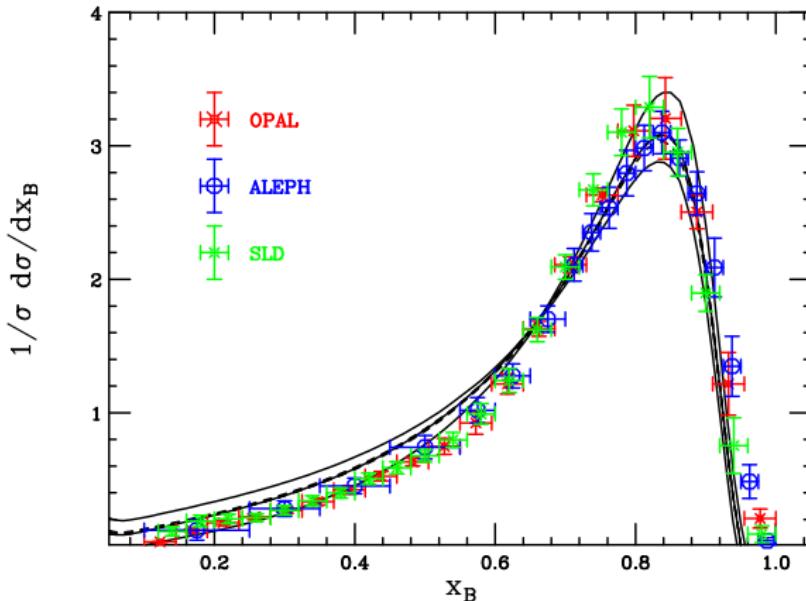


Figure 2: Model dependence on the factorizations scales of  $B$ -hadron spectrum in  $e^+e^-$  annihilation at  $Z^0$  peak. Solid lines:  $\mu_{0F} = m_b/2, m_b$  and  $2m_b$ ; dashed lines:  $\mu_F = m_Z/2, m_Z$  and  $2m_Z$ .



## Extraction of $\alpha_S(m_Z)$ and $V_{ub}$ from the data

- Since our model not contain free parameter to be fitted to the data we have been able to extract, for the first time, the value of  $\alpha_S(m_Z)$  from the experimental data of Babar, Belle, and Cleo Collaborations.
- For each spectrum (excluding the electron energy spectrum of BaBar and Belle) the extracted values of  $\alpha_S(m_Z)$  are in agreement with the current world average within at most two standard deviation [PDG ('06)].
- Our statistical analysis is preliminary, a complete analysis requires the exact knowledge of the experimental resolution functions.
- With such detailed statistical analysis it is also possible to extract the value  $V_{ub}$  of the CKM matrix.



$$\begin{aligned}
 g_1(\lambda) &= -\frac{A_1}{2\beta_0\lambda} [(1-2\lambda)\ln(1-2\lambda) - 2(1-\lambda)\ln(1-\lambda)]; \\
 g_2(\lambda) &= \frac{D_1}{2\beta_0} \ln(1-2\lambda) + \frac{B_1}{\beta_0} \ln(1-\lambda) + \frac{A_2}{2\beta_0^2} [\ln(1-2\lambda) - 2\ln(1-\lambda)] \\
 &\quad - \frac{A_1\beta_1}{4\beta_0^3} [2\ln(1-2\lambda) + \ln^2(1-2\lambda) - 4\ln(1-\lambda) - 2\ln^2(1-\lambda)] \\
 &\quad + \frac{A_1\gamma_E}{\beta_0} [\ln(1-2\lambda) - \ln(1-\lambda)]; \\
 g_3(\lambda) &= -\frac{D_2\lambda}{\beta_0(1-2\lambda)} - \frac{2D_1\gamma_E\lambda}{1-2\lambda} + \frac{D_1\beta_1}{2\beta_0^2} \left( \frac{2\lambda}{1-2\lambda} + \frac{\ln(1-2\lambda)}{1-2\lambda} \right) - \frac{B_2\lambda}{\beta_0(1-\lambda)} - \frac{B_1\gamma_E\lambda}{1-\lambda} \\
 &\quad + \frac{B_1}{\beta_0^2}\beta_1 \left( \frac{\lambda}{1-\lambda} + \frac{\ln(1-\lambda)}{1-\lambda} \right) - \frac{A_3}{2\beta_0^2} \left( \frac{\lambda}{1-2\lambda} - \frac{\lambda}{1-\lambda} \right) - \frac{A_2\gamma_E}{\beta_0} \left( \frac{1}{1-2\lambda} - \frac{1}{1-\lambda} \right) \\
 &\quad + \frac{A_2\beta_1}{2\beta_0^3} \left( \frac{3\lambda}{1-2\lambda} - \frac{3\lambda}{1-\lambda} + \frac{\ln(1-2\lambda)}{1-2\lambda} - \frac{2\ln(1-\lambda)}{1-\lambda} \right) - \frac{A_1\gamma_E^2}{2} \left( \frac{4\lambda}{1-2\lambda} - \frac{\lambda}{1-\lambda} \right) \\
 &\quad - \frac{A_1\pi^2}{12} \left( \frac{4\lambda}{1-2\lambda} - \frac{\lambda}{1-\lambda} \right) - \frac{A_1\beta_2}{4\beta_0^3} \left( \frac{2\lambda}{1-2\lambda} - \frac{2\lambda}{1-\lambda} + 2\ln(1-2\lambda) - 4\ln(1-\lambda) \right) \\
 &\quad + \frac{A_1\beta_1\gamma_E}{\beta_0^2} \left( \frac{1}{1-2\lambda} - \frac{1}{1-\lambda} + \frac{\ln(1-2\lambda)}{1-2\lambda} - \frac{\ln(1-\lambda)}{1-\lambda} \right) - \frac{A_1\beta_1^2}{2\beta_0^4} \left( \frac{\lambda}{1-2\lambda} - \frac{\lambda}{1-\lambda} \right) \\
 &\quad - \ln(1-2\lambda) + \frac{\ln(1-2\lambda)}{1-2\lambda} + \frac{\ln(1-2\lambda)^2}{2(1-2\lambda)} + 2\ln(1-\lambda) - \frac{2\ln(1-\lambda)}{1-\lambda} - \frac{\ln(1-\lambda)^2}{1-\lambda}.
 \end{aligned}$$



## Infrared Logarithm Resummation

- The resummation of large logarithms contributions is possible for those observable which exponentiate i.e. the matrix element and the phase space can be factorized by expressing the emission of  $n$  infrared (soft or collinear) gluons as the product of  $n$  single gluon emission.
- With general approximation can be demonstrated that the matrix elements of the gluon emissions exponentiate up a given logarithmic accuracy.
- In general the factorization of the phase space is not true, it depends by the particular process.
- In the case of the threshold resummation the exponentiation of the phase space can be demonstrated in the Mellin space:

$$f_N \equiv \int_0^1 f(z) z^{N-1} dz$$



## Inverse Mellin transform

- Integration over  $y$  in  $G_N$  is performed exactly in numerical way; this is possible because the time-like coupling does not have the Landau pole and is regular for any value of  $N$ .
- The inverse transform from  $N$ -space to  $x$ -space is also made exactly in numerical way by the formula

$$\sigma_N(\tilde{\alpha}_S) \equiv \int_0^1 (1-y)^{N-1} \sigma(y, \tilde{\alpha}_S) dy,$$

$$\sigma(y; \tilde{\alpha}_S) = \int_{C-i\infty}^{C+i\infty} \frac{dN}{2\pi i} (1-y)^{-N} \sigma_N(\tilde{\alpha}_S),$$

where the constant  $C$  is chosen so that the integration contour in the  $N$ -plane lies to the right of all the singularities of  $\sigma_N(\alpha_S)$ .



- Most previous analyses predicted  $B$ -hadron production convoluting the parton-level spectrum with a non-perturbative fragmentation function which contains few parameters which are to be fitted to experimental data.
- Instead of fitting the parameters of a hadronization model we model non-perturbative effects by the use of an analytic effective coupling constant which does not contain any free parameter, extending the analysis carried out in the framework of heavy-flavour decays.
- We shall then be able to compare our predictions directly with data, without using any extra hadronization model.



# Perspective

The work presented in this talk permit a wide range of improvements and develops. Some of theme are in progress as well:

- Analyze the case of  $B \rightarrow X_c l \nu$  decays similarly to the  $B \rightarrow X_u l \nu$  decays. The theoretical complication in this case is given by the not negligible mass of the  $c$ -quark. However the experimental data for such processes are much more accurate.
- There are some other spectra in  $B \rightarrow X_u l \nu$  decays which have phenomenological interest which can be computed with our formalism.
- In the  $b$ -quark fragmentation case it is possible to perform a complete (NNLO+NNLL) resummation. After this we are confident that the reduced theoretical uncertainties will permit us to extract  $\alpha_s(m_Z)$  as we have done with the decay spectra.
- Other possible extensions is to apply our formalism to other process and observables as the  $B$  production in top and Higgs decays or charm production in  $e^+e^-$  annihilations at the  $Z^0$  peak ( $m_Z \sim 90$  GeV) or even much below at  $\Upsilon(4s)$  peak ( $m_\Upsilon \sim 10$  GeV).

