Deep Learning In the Wild: Application on ATLAS

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# **Challenges for ML on ATLAS?**

- Successfully physics program at the LHC requires overcoming major hardware, computing, and analysis challenges!
- How will the increasing power of ML play a role?

#### **Analysis Challenges** 120 SM



#### **Trigger Challenges** 10<sup>5</sup> TLAS Simulation HL-LHC Trigger, Vs=14 TeV $HH \rightarrow b\overline{b}b\overline{b}, \langle \mu \rangle = 200$ 10<sup>4</sup> Jet hl<2.5



#### **Tracking Challenges**



#### **Simulation Challenges**



# **The ATLAS Experiment**

 $\sim 10^8$  detector channels

<u>Data:</u> ~300 MB / sec ~3000 TB / year

<u>Weight:</u> 7000 tons <u>Size:</u> 46 m long, 25 m high, 25 m wide









**Generative Model** 

# From Theory to Experiment... and Back

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$$\begin{split} & -\frac{1}{2}(d_{2}^{2}(0,d_{2}^{2}-g_{1}^{-H}\partial_{2}^{2})d_{2}^{2}d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-W_{1}^{2})d_{2}^{2}d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-W_{1}^{2})d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-W_{1}^{2})d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-W_{1}^{2})d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-W_{1}^{2})d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-W_{1}^{2}-d_{2}^{2}-W_{1}^{2}-\frac{1}{2})d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-d_{2}^{2}-W_{1}^{2}-\frac{1}{2})d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-\frac{1}{2})d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-\frac{1}{2})d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-d_{2}^{2}-\frac{1}{2})d_{2}^{2}-\frac{1}{2}(d_{2}^{2}-d_{$$

$$\begin{split} &\gamma^{3}(w_{1}^{3})|_{1}^{4} = \frac{w_{1}}{2\sqrt{2}} \frac{1}{m^{2}} \left[ -\phi^{2}(k^{2}(1-\gamma^{3})e^{\lambda}) + \phi^{-}(e^{\lambda}(1+\gamma^{3})\mu^{\lambda}) \right] - \\ &\frac{2}{3} \frac{1}{m^{2}_{2}} \left[ H(e^{\lambda}e^{\lambda}) + i\phi^{2}(e^{\lambda}\gamma^{2}e^{\lambda}) \right] + \frac{1}{m^{2}_{2}\sqrt{2}} \phi^{-1}(m^{2}_{3}(w^{2}_{2}(L-\gamma^{3})d^{2}_{2}) + \\ &m^{2}_{6}(e^{\lambda})^{2}C_{M}(1+\gamma^{5})d^{2} \right] + \frac{1}{m^{2}_{2}\sqrt{2}} \phi^{-1}(m^{2}_{3}(d^{2})C_{M}(1+\gamma^{5})d^{2}_{2}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{3} \frac{1}{m^{2}} H(d^{2}_{1}d^{2}_{1}) - \frac{2}{3} \frac{1}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{3} \frac{1}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{3} \frac{1}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{3} \frac{1}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{3} \frac{1}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{3} \frac{1}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{3} \frac{1}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{3} \frac{1}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{3} \frac{1}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{3} \frac{1}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{3} \frac{1}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{3} \frac{1}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{3} \frac{1}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) - \frac{2}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) - \frac{2}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{m^{2}_{3}} H(d^{2}_{1}d^{2}_{1}) + \frac{2}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) - \\ &\gamma^{2}(w^{2}_{1}) - \frac{2}{m^{2}_{3}} H(d^{2}_{1}) + \frac{2}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) + \frac{2}{m^{2}_{3}} \phi^{2}(d^{2}_{1}\gamma^{2}_{1}) + \\ &\gamma^{2}(w^{2}_{1}$$

$$\begin{split} \gamma^{-}(y_0) &= \frac{1}{2}\pi^{-}W(y_0^{-}y_0) - \frac{1}{2}\pi^{-}W(a_0^{-}y_0) + \frac{1}{2}\pi^{-}\varphi(w_0^{-}\gamma)u_0^{-}, \\ &= \frac{1}{2}\pi^{-}_{0}\phi(a_0^{-}\gamma)a_0^{-}X^{+}(\delta_0^{-}-\delta_0^{-}X^{+}X^{+}) + igs_w W_{\mu}^{+}(\partial_{\mu}X^{-}-\partial_{\mu}X^{+}X^{+}) + igs_w W_{\mu}^{+}(\partial_{\mu}X^{-}X^{-}-\partial_{\mu}X^{+}X^{+}) + igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}Y^{-}-\partial_{\mu}X^{-}X^{+}) + igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}Y^{-}-\partial_{\mu}X^{-}X^{+}) + igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}Y^{-}-\partial_{\mu}X^{-}X^{+}) - igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}Y^{-}-\partial_{\mu}X^{-}X^{-}) - igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}Y^{-}-\partial_{\mu}X^{-}X^{-}) - igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}Y^{-}-\partial_{\mu}X^{-}X^{-}) - igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}Y^{-}) - igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}) - igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}) - igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}) - igs_w W_{\mu}^{-}(\partial_{\mu}X^{-}) - igs_w W_{\mu}^{$$

 $\begin{array}{l} \frac{1-2c_w^2}{2c_w}igM[\bar{X}^+X^0\phi^+-\bar{X}^-X^0\phi^-]+\frac{1}{2c_w}igM[\bar{X}^0X^-\phi^+-\bar{X}^0X^+\phi^-]+\\ igMs_w[\bar{X}^0X^-\phi^+-\bar{X}^0X^+\phi^-]+\frac{1}{2}igM[\bar{X}^+X^+\phi^0-\bar{X}^-X^-\phi^0] \end{array}$ 

Parameters  $\theta$ 



O(10) particles

O(100) particles

O(10<sup>8</sup>) detector elements



Slide credit: K. Cranmer

# ML Across the ATLAS Analysis Pipeline

- **Reconstruction** discrimination / regression problems, leading to fastest uptake of new ideas
  - Jets: Tagging, Calibration, Decorrelation
  - Missing Energy Pileup Subtraction
  - Jet Flavour Tagging
  - Tau Particle ID
  - Pixel Clustering for Tracking
- Simulation problems of density estimation and sampling
   Fast Calorimeter Simulation
- Analysis largely well known ML methods for signal vs background discrimination and event reconstruction

   Very impactful, but not going to talk about this much
- ML applications growing more sophisticated
   − Classification → Density estimation → Differentiable Programs

# ML Across the ATLAS Analysis Pipeline

- **Reconstruction** discrimination / regression problems, leading to fastest uptake of new ideas
  - Jets: Tagging Calibration Decorrelation



How can existing ML allows us to approach new challenges

How can we design ML systems to our needs

ML applications growing more sophisticated
 − Classification → Density estimation → Differentiable Programs

# **Looming Gap in Theory vs Practice**

- Large difference between what is done in phenomenology papers and on the experiment (at least what is public)
- Why????
  - Real detector models have MUCH more complex noise than simplified simulation
    - Method performance doesn't necessarily transfer
    - Even the ones that transfer can be hard to tune
  - Calibration even our best simulations for training are not perfect
    - After training the algorithm, we still have to calibrate!
  - Information disconnect
    - Model expertise may be outside experiment
    - Different people build and calibrate algorithm within ATLAS
  - Experiment computational resources may not be well suited to ML
    - Moreover, full data re-processing alone can take months
  - Resistance to change



# Classification and Regression in Reconstruction

# Jets at the LHC

- Jets are formed by clustering energy depositions in calorimeter with the anti- $k_{\rm T}$  algorithm
- Jet identification = Classification: p(parent particle | jet cluster)
- Energy estimation = Inference, regression:  $p(E_{true}^{jet} | jet cluster)$





#### **Canonical Discrimination Problem: Jet Identification**



# **Combining Substructure Variables**



- Wide array of physics insight has gone into developing jet substructure observables
- Direct application of ML for combining power of multiple partially correlated substructure features
- First calibrations look quite reasonable!

# Jets as Images



- A jet induces a distribution of energy over  $\eta \phi$ – Essentially how a jet is seen by calorimeters
- Jet-image fixed size 2D representation of the jet as a distribution of energy

# Jet Images

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#### Unrolled slice of detector jet image Boosted W ---> qq 10<sup>2</sup> Pixel $p_{_{T}}$ [GeV] 0.5 jet 10 proton-proton $\phi$ collision into/ out-of page 1 -0.5 jet -0.5 0.5 -1 0 not to scale [Translated] Pseudorapidity (η) jet image

#### Calorimeter towers as pixels Energy depositions as intensity

# Jet Images

#### Average of large number of Jet Images jet image $250 < p_T/GeV < 260 GeV, 65 < mass/GeV < 95$ Pythia 8, W' $\rightarrow$ WZ, $\sqrt{s} = 13 \text{ TeV}$ [Translated] Azimuthal Angle (<a) Pixel p<sub>T</sub> [GeV] 10<sup>2</sup> 10 0.5 10-1 10<sup>-2</sup> jet W-jets 10<sup>-3</sup> 10-4 10<sup>-5</sup> -0.5 10<sup>-6</sup> proton-proton $\phi$ 10-7 collision into/ 10<sup>-8</sup> out-of page 10<sup>-9</sup> -1 -0.5 0 0.5 [Translated] Pseudorapidity (n) jet 250 < p<sub>x</sub>/GeV < 260 GeV, 65 < mass/GeV < 95 Pythia 8, QCD dijets, $\sqrt{s} = 13 \text{ TeV}$ 0<sup>3</sup> Pixel p<sub>r</sub> [GeV] 10<sup>2</sup> 10 not to scale 0.5 10-1 QCD-jets 10<sup>-2</sup> jet image 10<sup>-3</sup> 10-4 10<sup>-5</sup> -0.5 10<sup>-6</sup> 10<sup>-7</sup> 10<sup>-8</sup> 10<sup>-9</sup>

-1

-0.5

0

0.5 [Translated] Pseudorapidity (η)

#### Image Credit: **B. Nachman**



# Jet Images on ATLAS: Quarks vs Gluons

ATL-PHYS-PUB-2017-017







Quark Jet Efficiency

Note: Other experimental results on CMS with ImageTop, a boosted top tagger based on images <u>CMS-PAS-JME-18-002</u>

### **Calo-Images for Missing Energy Pileup Removal**

- *Input*: images of calorimeter clusters and tracks
- *Output*: NN regress to predict hard scatter energy in each calorimeter tower
- Gains in resolution
  - NNN doesn't learn accurate cell by cell predictions
  - Considering new ways to define loss





# **Reconstructing Bottom Quark Jets**



# **Bottom Quark Jet Identification**





- Goal: Discriminate b-jets from non-b-jets
- Track Impact Parameter based taggers: p(jet flavor | tracks in jet)
  - Dimensionality too high for histogram density estimation
  - Often make naïve Bayes assumption that tracks independent!

### Jets as Sequences



- Jets are a grouping of a variable number of particles
- With physically motivated ordering: jet as a sequence

#### **Recurrent Neural Networks**



Image credit: <u>F. Fleuret</u>



Image credit: <u>F. Fleuret</u>



Image credit: F. Fleuret





Image credit: F. Fleuret

# Long Short Term Memory (LSTM)

- Gating:
  - network can grow very deep,
    in time → vanishing gradients.



- *Critical component*: add pass-through (additive paths) so recurrent state does not go repeatedly through squashing non-linearity.
- LSTM:
  - Add internal state separate from output state
  - Add input, output, and forget gating



### Jets and Sequence Processing



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ATL-PHYS-PUB-2017-003

# **RNN b-tagging**

- Order tracks by impact parameter
- RNN can learn inter-track dependencies



# **Combining With Other Algorithms**



1/(false-positive rate) at fixed true-positive rate vs jet  $p_T$ 

ATLAS-PHY-PLOTS-FTAG-2019-005

### Calibration



\*This is actually a slightly older version of the algorithm than previous slide

arXiv:1910.08447

#### **Real World Impact: Dijet Resonances**



### Tau RNN





### Tau RNN





### Density Estimation and Generative Modeling

# Why Generative Modeling

- Discriminative models:  $f(x) \approx \overline{y} = E_{p(y|x)}[y]$
- How do we model uncertainty on predictions, i.e. learn a posterior on likelihood?
- Generative models aim to estimate density p(x) or conditional p(x | y)
  - Explicitly: can compute the value  $p(\cdot)$
  - Implicitly: can draw samples from  $p(\cdot)$
  - More on Generative Models in A. Butter's Talk

# **Fast Simulation**

- Increased pileup at HL-LHC will push boundaries of our computational capabilities for simulation
- Full Simulation
  - Accurate but costly to sample
- Fast Simulation
  - Sample from parametric average shower model
  - Doesn't account for correlations in shower shape fluctuations







ATLAS Simulation Preliminary

# **Fast Simulation**

- Increased pileup at HL-LHC will push boundaries of our computational capabilities for simulation
- Full Simulation
  - Accurate but costly to sample
- Fast Simulation
  - Sample from parametric average shower model



- Correct average: Model correlated fluctuations on top of average
- *Full ML approach*: Learn generative mode of distribution of showers, p(x), and produce samples



• Method: Fit parametric approximation to correlated noise distribution

• Sklar's Theorem: given a random vector  $(X_1, ..., X_n)$ , the joint cumulative distribution function  $H(x_1, ..., x_n) = P(X_1 \le x_1, ..., X_n \le x_n)$ 

$$\Pi(x_1, ..., x_n) = \Gamma(x_1 \ge x_1, ..., x_1 \ge x_1)$$

can be expressed using marginals  $F_i(x_i) = P(X_1 \le x_1)$  as  $H(x_1, ..., x_n) = C(F_1(x_1), ..., F_n(x_n))$ 

where  $C(\cdot)$  is the copula

# **Correcting the Average**

• Method: Fit parametric approximation to correlated noise distribution with a *Gaussian Copula* 

Gaussian Copula

$$C_R^{ ext{Gauss}}(u) = \Phi_R\left(\Phi^{-1}(u_1),\ldots,\Phi^{-1}(u_d)
ight)$$

Gaussian Copula density

$$c_R^{ ext{Gauss}}(u) = rac{1}{\sqrt{\det R}} \exp\left(-rac{1}{2} egin{pmatrix} \Phi^{-1}(u_1)\dots\ \Phi^{-1}(u_d) \end{pmatrix}^T \cdotig(R^{-1}-Iig) \cdotegin{pmatrix} \Phi^{-1}(u_1)\dots\ \Phi^{-1}(u_d) \end{pmatrix}
ight)$$

- 1. CDF transform inputs  $x_i$  to uniform  $u_i$
- 2. Fit copula to sample of correlated uniform variables
- 3. Sample Copula to get  $u_i$  and invert CDF to get  $x_i$

### **Correcting the Average with Copula**





# **Deep Generative Models**





#### Generative Adversarial Networks (GAN) [arXiv:1406.2661]



- Generator produces images from random noise and tries to trick discriminator into thinking they are real
- Classifier tries to tell the difference between real and fake images

 $\min_{G} \max_{D} V(D,G) = \mathbb{E}_{x \sim p_{data}(x)} \left[ \log D(x) \right] + \mathbb{E}_{z \sim p_z(z)} \left[ 1 - \log D(G(z)) \right]$ 

- Two-player minimax game between Generator (G) and Discriminator (D) networks
- Training involves carful and often unstable iteration between updating G parameters ( $\theta$ ) and D parameters ( $\psi$ )
- If perfectly trained, generator converges to implicit model of data density:  $G(z) = x \sim p_{data}(x)$





#### **GANs for Calorimeter Energy Depositions**





### Making use of Generative Modeling Tools

# **Adversarial Learning for Enforcing Invariance**

- With flexibility come complexity:
  - Hard to control how models learn / utilize information
  - Potentially unwanted sensitivity to poorly modeled aspects of simulation
  - Potentially unwanted sculpting of key physics distributions like mass
- *Idea*: Augment training of classifier to enforce invariance to changes in a variable Z (nuisance parameter for systematic uncertainty, kinematic variables, etc.)
  - Several ways to do this, see D. Shih's Talk

### **Adversarial Networks**



• Classifier built to solve problem at hand

### **Adversarial Networks**



- Loss that encodes performance of a classifier and adversary
- Classifier penalized when adversary does well at predicting Z

arXiv:1611.01046

$$\hat{\theta}_f, \hat{\theta}_r = \arg\min_{\theta_f} \max_{\theta_r} E(\theta_f, \theta_r).$$

$$E_{\lambda}(\theta_f, \theta_r) = \mathcal{L}_f(\theta_f) - \lambda \mathcal{L}_r(\theta_f, \theta_r),$$



- Hyper-parameter  $\lambda$  controls trade-off
  - Large  $\lambda$  enforces f(...) to be pivotal, e.g. robust to nuisance
  - Small  $\lambda$  allows f(...) to be more optimal without Z variation

# Learning to Pivot: Toy Example

• 2D example

$$x \sim \mathcal{N}\left((0,0), \begin{bmatrix} 1 & -0.5\\ -0.5 & 1 \end{bmatrix}\right) \quad \text{when } Y = 0,$$
  
 $x \sim \mathcal{N}\left((1,1+Z), \begin{bmatrix} 1 & 0\\ 0 & 1 \end{bmatrix}\right) \quad \text{when } Y = 1.$ 





With Adversary 4.0  $p(f(X)|Z = -\sigma)$ 3.5 p(f(X)|Z=0)3.0  $p(f(X)|Z = +\sigma)$ 2.5 ((X)f)d1.5 1.0 0.5 0.0 0.0 0.2 0.4 0.6 0.8 1.0 f(X)



# **Learning to Pivot: Physics Example**



- Standard training with no systematics during training, evaluate systematics after training
- λ=0
  - Training samples include events with systematic variations, but no adversary used

• λ=10

 Trading accuracy for robustness results in net gain in terms of statistical significance

[AMS = Estimate of statistical significance including systematic uncertainty]

arXiv:1611.01046

# **Decorrelating Variables**

- Same adversarial setup can decorrelate a classifier from a chosen kinematic variable [<u>arXiv:1703.03507</u>]
- Example: decorrelate classifier from jet mass, so as not to sculpt jet mass distribution with classifier cut

#### W-jets vs. QCD Jets



ATL-PHYS-PUB-2018-014

# **Decorrelating Variables**

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- Example: decorrelate classifier from jet mass, so as not to sculpt jet mass distribution with classifier cut



# Non-ATLAS, but related, use of Generative Modeling Tools

# **Black Box Optimization**

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• Goal: Optimize simulator parameters to minimize objective

# **Black Box Optimization**

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• Goal: Optimize simulator parameters to minimize objective

# **Black Box Optimization**



- Goal: Optimize simulator parameters to minimize objective
- Can we approximate the simulator directly?

# **Differentiable Surrogates**



- Train parameterized generative surrogate model S, i.e. GAN or flow, to approximate  $F(x; \psi)$ 
  - Noise input to surrogate can account for stochastic nature of F

# **Differentiable Surrogates**

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# **Differentiable Surrogates**

-2

 $\psi_1$ 

-0.60



• Optimize objective with gradient descent using trained surrogate to produce differentiable samples

# **Optimization on Toy Examples**

Under Review S. Shirbikov, V. Belavin, M. K., A. Baydin, A. Ustyuzhanin



#### **Rosenbrock function**



100 Dim parameter space projected on 10 Dim submanifold

#### **Optimization In Physics Example**

Under Review S. Shirbikov, V. Belavin, M. K., A. Baydin, A. Ustyuzhanin





# Conclusion

- Analysis pipeline is grounded in our detail physics domain knowledge
- Maintain our physics knowledge embedded in this pipeline while using ML to help solve some of the intractable challenges we face on ATLAS
- ML methods have shown strong performance improvements in reconstruction and analysis methods
- Techniques to deal with key challenges such as simulation computational cost and systematic uncertainty mitigation are under study





