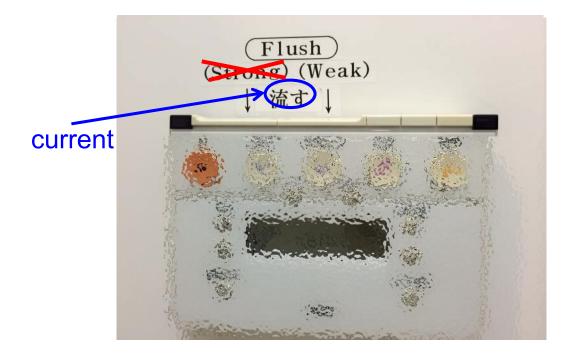
LONG-BASELINE NEUTRINO EXPERIMENTS

ATSUKO K. ICHIKAWA, KYOTO UNIVERSITY, JAPAN

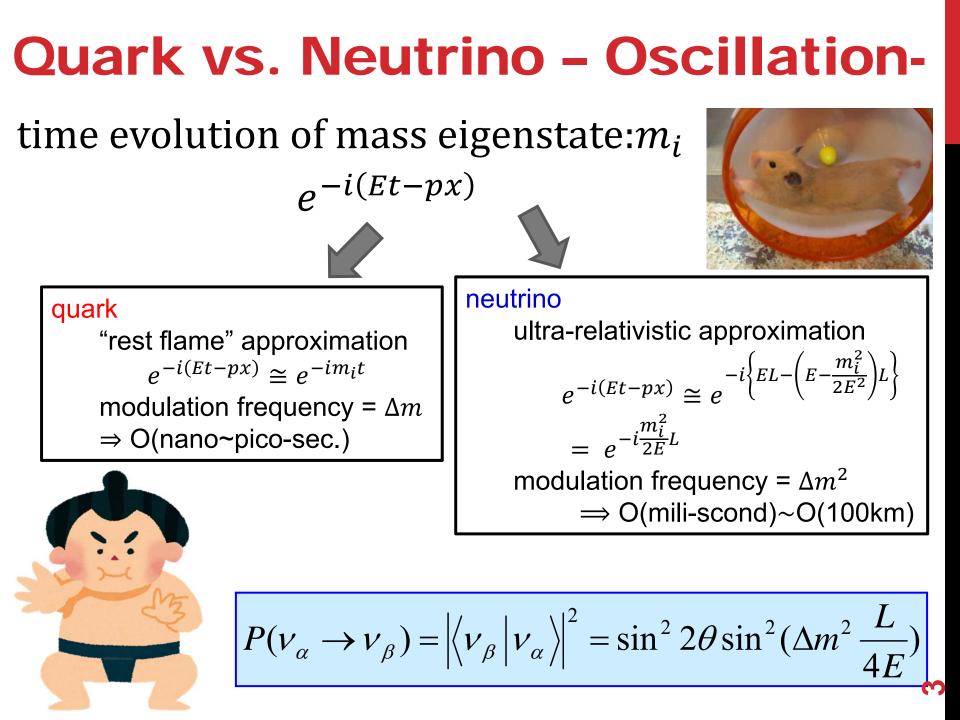






Interaction length of neutrinos from typical sources for $\rho \sim 1$ g/cm³ material

Atmospheric or accelerator neutrinos (~1GeV) : 10⁸km ~ 1 AU
 Solar or reactor neutrinos (~MeV) : 10¹⁴km ~ 100 light-year



Diverse Page Summary
Mixing matrix of leptons

$$U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & +c_{23} & +s_{23} \\ 0 & -s_{23} & +c_{23} \end{pmatrix} \begin{pmatrix} +c_{13} & 0 & +s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & +c_{13} \end{pmatrix} \begin{pmatrix} +c_{12} & +s_{12} & 0 \\ -s_{12} & +c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

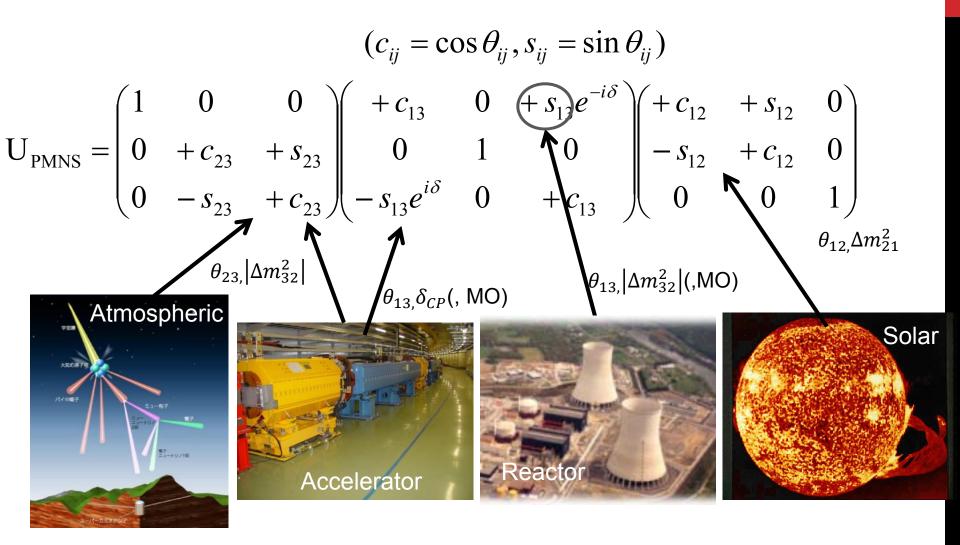
$$\begin{pmatrix} c_{ij} = \cos \theta_{ij}, s_{ij} = \sin \theta_{ij} \end{pmatrix}$$

$$\theta_{12} = 33^{\circ} \pm 1^{\circ}$$

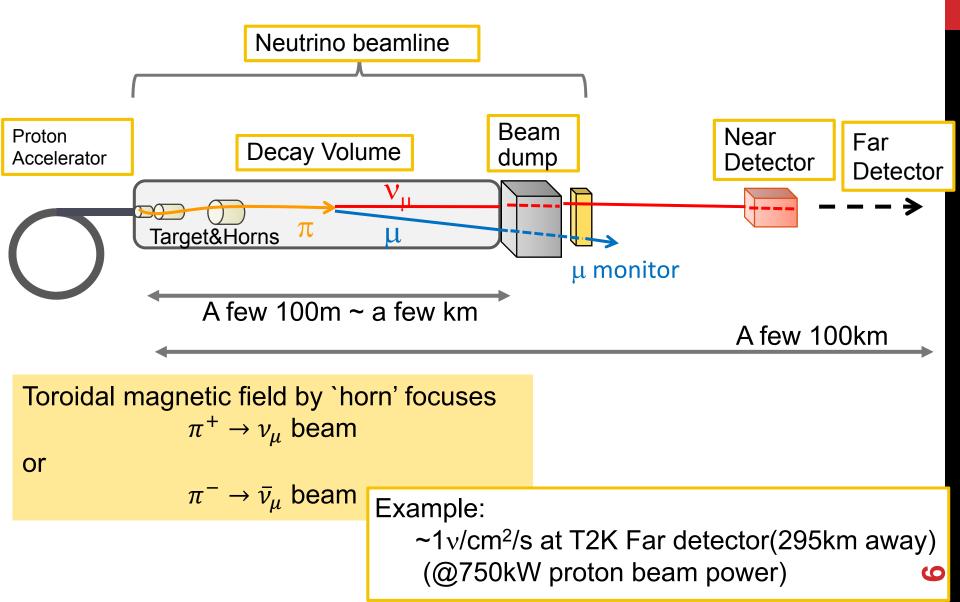
$$\theta_{23} = \text{How close to 45}^{\circ} ?$$

$$\theta_{13} = 8.9^{\circ} \pm 0.4^{\circ}$$
Diverse Mass ordering
normal: $m_1 < m_2 \ll m_3$
inverted: $m_3 \ll m_1 < m_2$
 $\Delta m_{21}^2 = 7.5 \pm 0.2 \times 10^{-5} \text{ eV}^2$
 $|\Delta m_{32}^2| = 2.44 \pm 0.06 \times 10^{-3} \text{ eV}^2$
Diverse Mass ordering
normal ordering

Neutrino oscillations



Components of the Long Baseline Neutrino Experiment



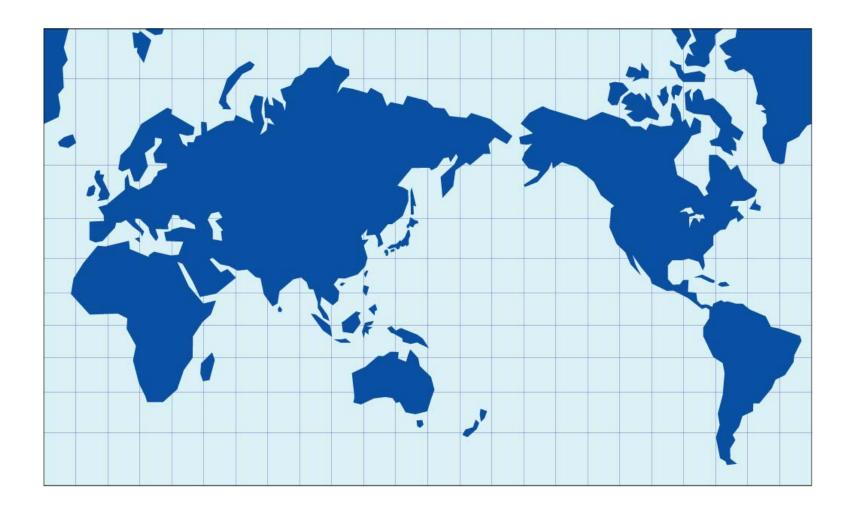
Oscillations at the Long baseline Accelerator experiments

 Δm_{21}^2 terms are small. $\Delta m_{32}^2 (\sim \Delta m_{31}^2)$ terms are dominant \rightarrow Can be simplified.

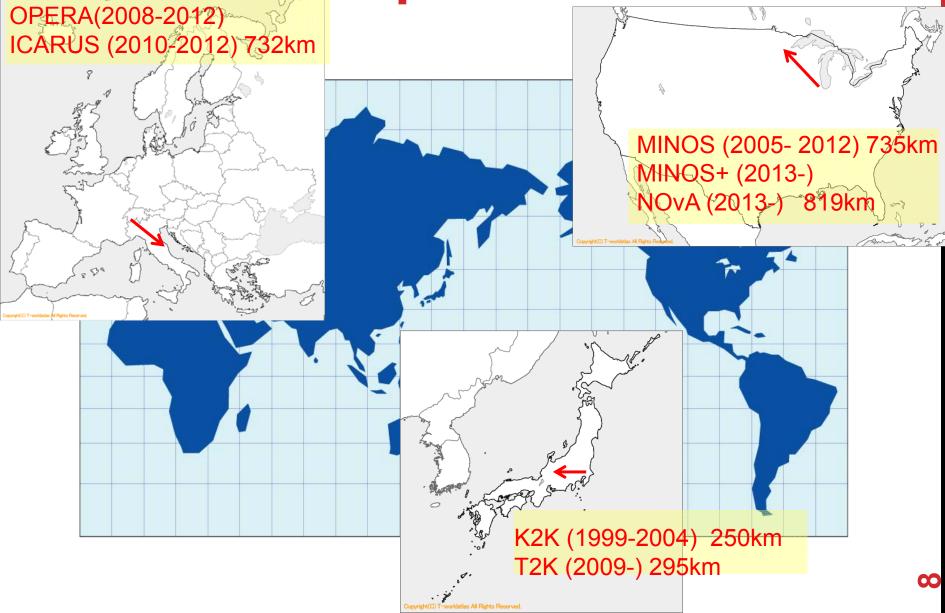
Leading Term

L/E (km/GeV)

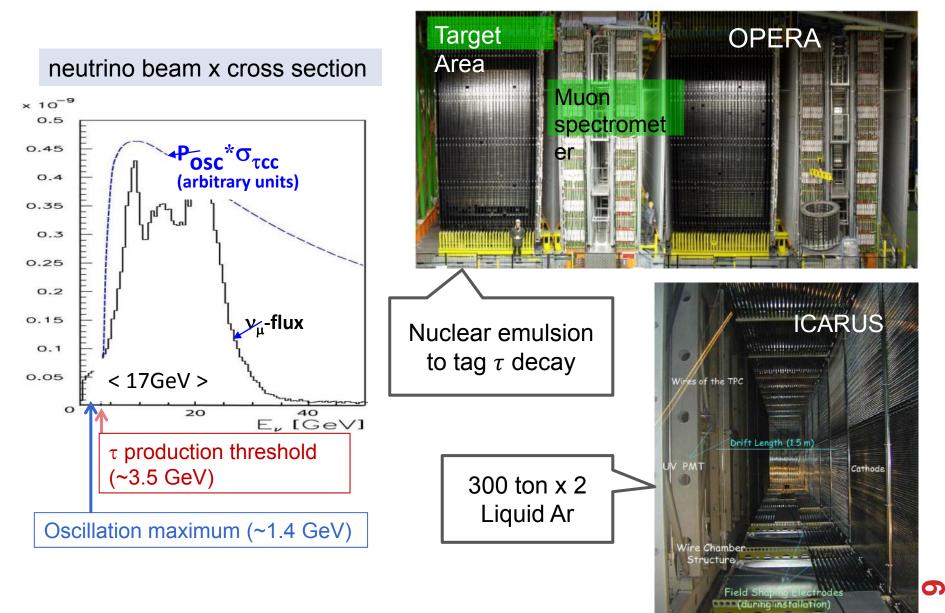
World Long baseline v oscillation experiments



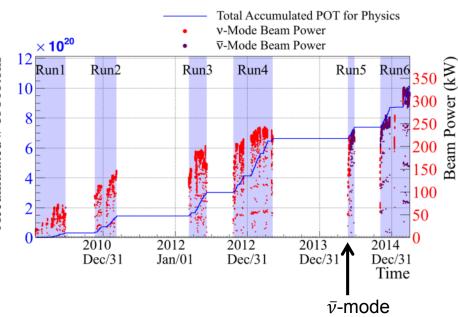
World Long baseline v oscillation experiments

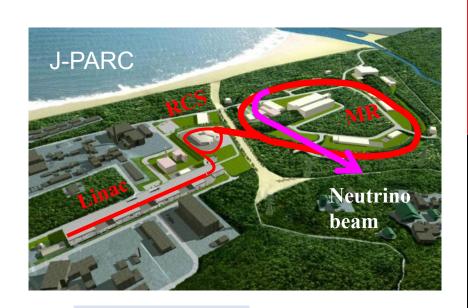


OPERA/ICARUS

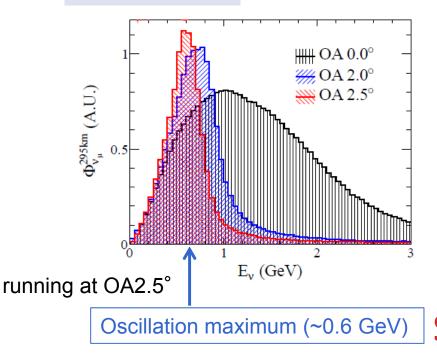


T2K -beam-





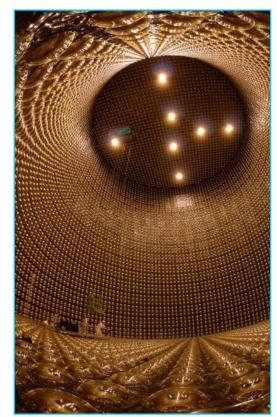
neutrino beam



- Maximum beam power achieved 345kW (goal 750kW)
- ➢ Integrated protons-on-target 7.0x10²⁰(v beam) + 3.1x10²⁰(v beam) (goal 78x10²⁰ in total)

Accumulated # of Protons

T2K - Far detector -

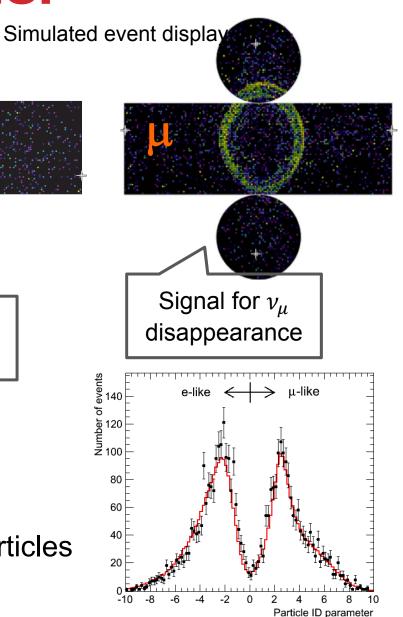


Water Cherenkov Detector

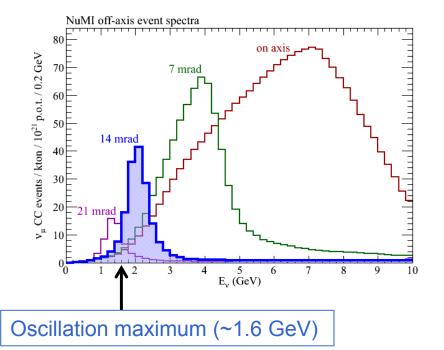
Very good PID for sub-GeV particles mis-identification ~1%

Signal for v_e

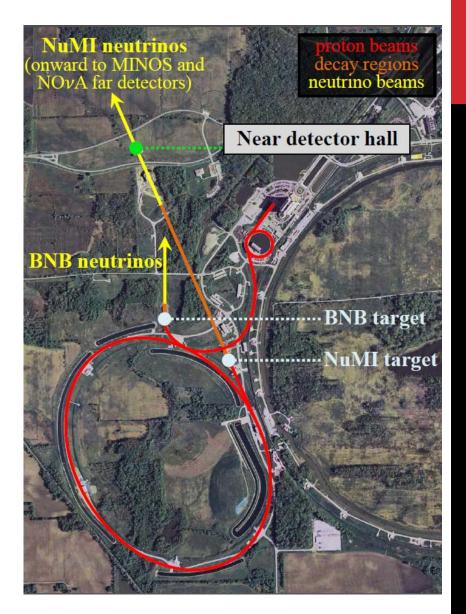
appearance

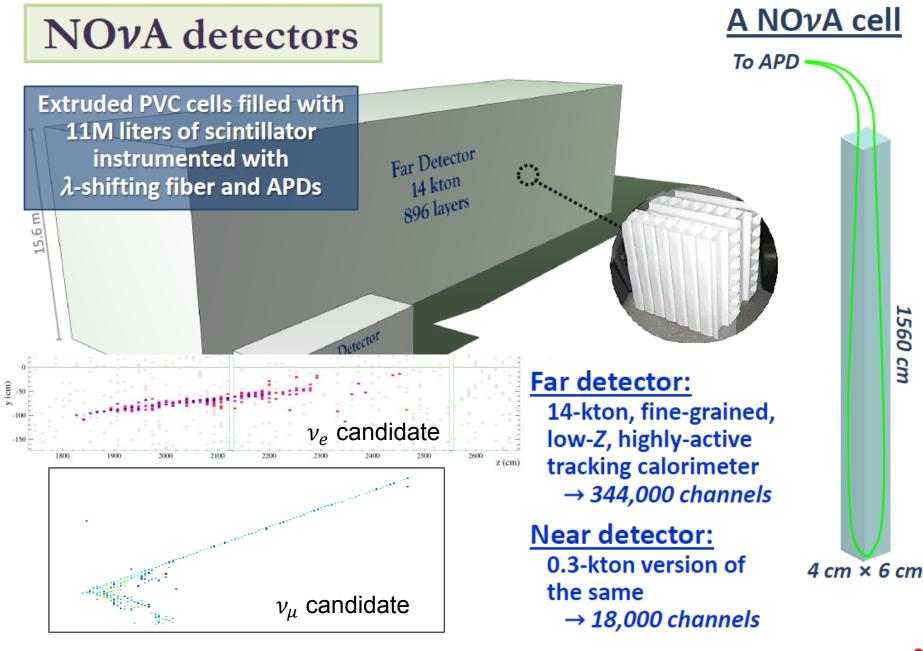


NOvA -beam-



start in 2013. Now running at 420kW (goal:700kW)





IPA, May 6, 201

NOvA detectors



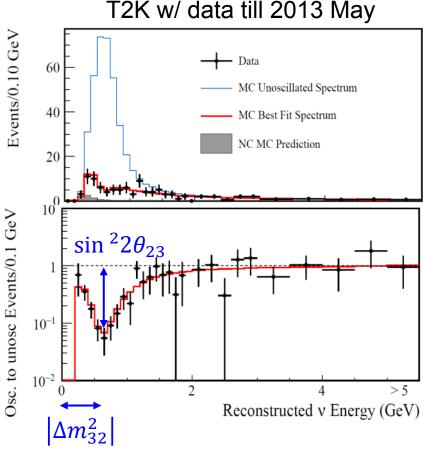




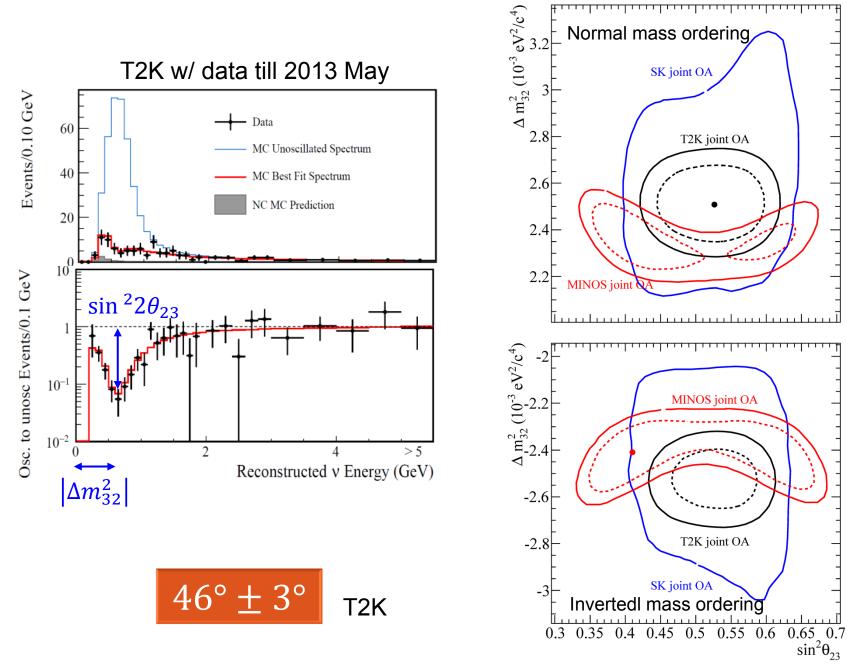
World Long baseline v oscillation experiments

	Accelerator		Detector
	Energy	Power(Planned)	Delector
K2K	12 GeV	7kW	50kt
MINOS	120 GeV	400 kW	5.4kt
OPERA/ICARUS	400 GeV/c	500kW	1.2kt/0.6kt
T2K	30 GeV	350kW (750kW)	50kt
NOvA	120 GeV	420kW (700kW)	14kt

v_{μ} disappearance - Is θ_{23} 45°?-



T2K w/ data till 2013 May



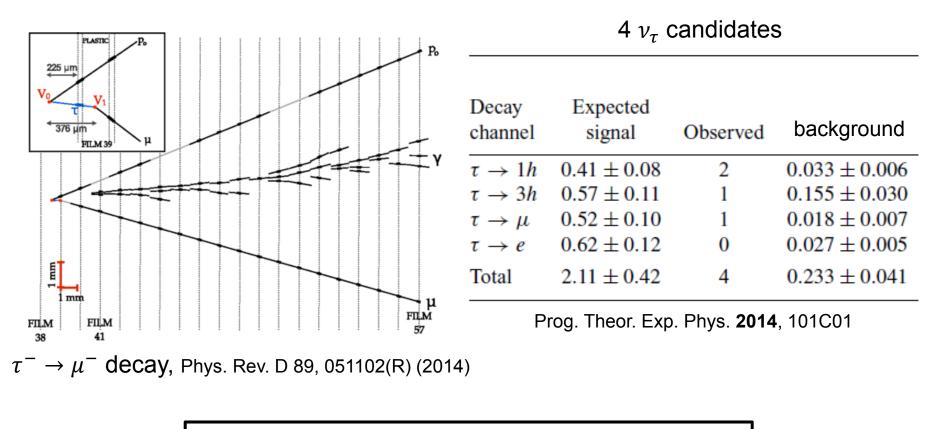
Phys. Rev. D 91, 072010 (2015)

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v_{τ} appearance - Did disappeared v_{μ} really changes to v_{τ} ? -

OPERA

3.8 σ v_{τ} appearance by Super-K atmospheric data *(Abe et al., PRL 110, 181802 (2013))* from a sample of enhanced τ -like events. OPERA identifies τ production in event-by-event basis.



 $\nu_{\mu} \rightarrow \nu_{\tau}$ oscillation is confirmed by 4.2 σ significance

V_e appearance - Golden mode for CP phase and mass ordering-

mixing matrix (PMNS)

$$\nu_{\alpha} \rightarrow \nu_{\beta} \text{ oscillation}$$
$$|\nu_{\alpha}(L)\rangle = U_{\alpha i} e^{-i\frac{m_{i}^{2}}{2E}L}$$
$$\therefore \langle \nu_{\beta} | \nu_{\alpha}(L) \rangle = U_{\beta i}^{*} U_{\alpha i} e^{-i\frac{m_{i}^{2}}{2E}L}$$
if $\beta = \alpha$
$$\langle \nu_{\alpha} | \nu_{\alpha}(L) \rangle = |U_{\alpha i}|^{2} e^{-i\frac{m_{i}^{2}}{2E}L}$$

Imaginary part vanishes! Appearance is necessary to see CP violation

v_e appearance (complete version in vacuum)

Leading term

$$P(\nu_{\mu} \rightarrow \nu_{\ell}) = 4C_{13}^{2}S_{13}S_{23}^{2}\sin^{2}\Phi_{31}$$

$$+8C_{13}^{2}S_{12}S_{13}S_{23}(C_{12}C_{23}\cos\delta - S_{12}S_{13}S_{23})\cos\Phi_{32}\sin\Phi_{31}\sin\Phi_{21}$$

$$-8C_{13}^{2}C_{12}C_{23}S_{12}S_{13}S_{23}\sin\delta\sin\Phi_{32}\sin\Phi_{31}\sin\Phi_{21}$$

$$CPC$$

$$-8C_{13}^{2}C_{12}C_{23}S_{12}S_{13}S_{23}\sin\delta\sin\Phi_{32}\sin\Phi_{31}\sin\Phi_{21}$$

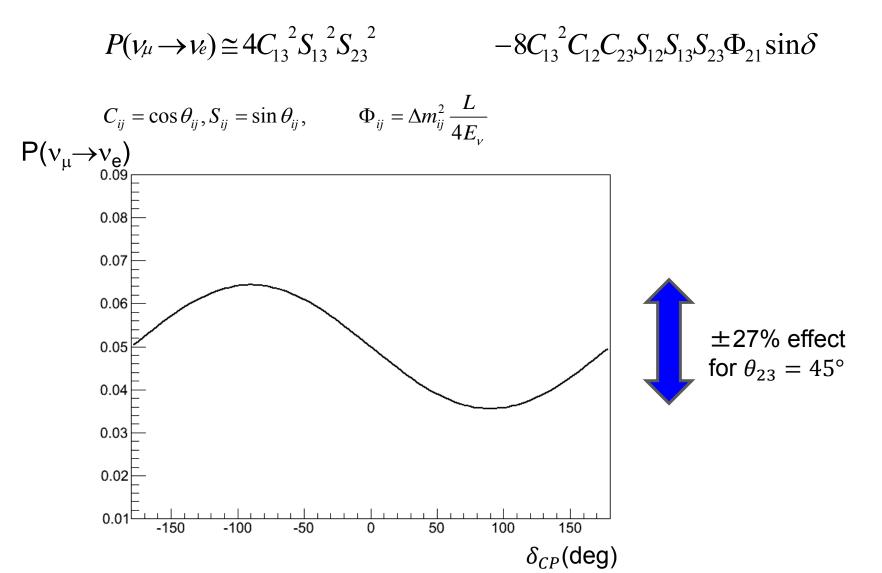
$$CPV$$

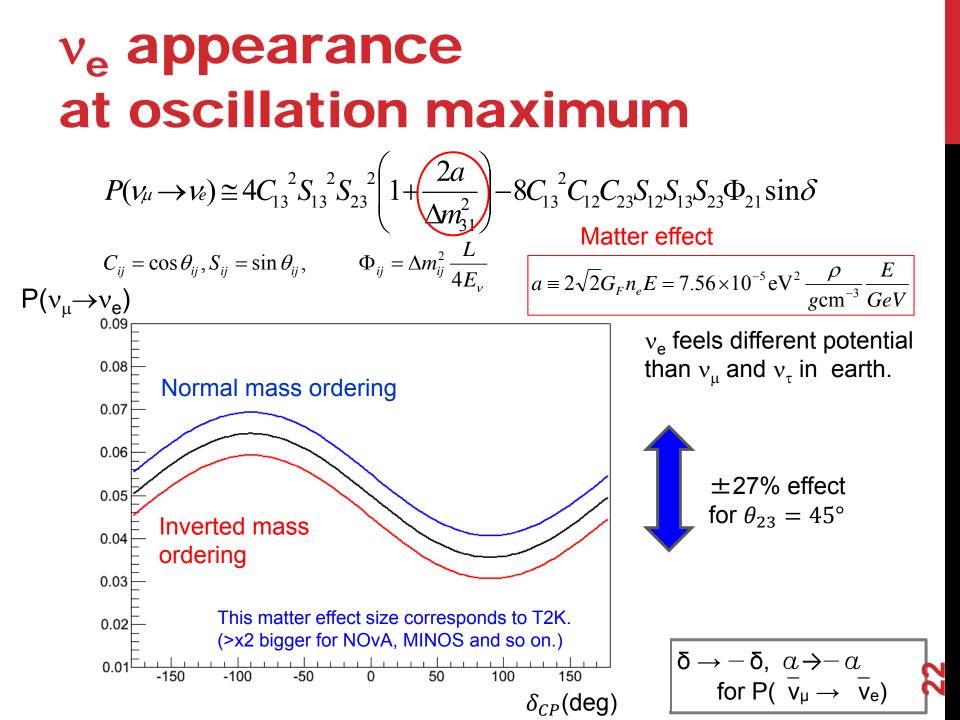
$$+4S_{12}^{2}C_{13}^{2}(C_{12}^{2}C_{23}^{2} + S_{12}^{2}S_{23}^{2})_{13}^{2} - 2C_{12}C_{23}S_{12}S_{23}S_{13}\cos\delta)\sin^{2}\Phi_{21}$$
Solar
$$C_{ij} = \cos\theta_{ij}, S_{ij} = \sin\theta_{ij}$$

$$\Phi_{ij} = \Delta m_{ij}^{2}\frac{L}{4E_{\nu}}$$

$$CP \text{ violating term introduced by interference btw. } \theta_{13} \text{ and } \theta_{12}$$

v_e appearance at oscillation maximum

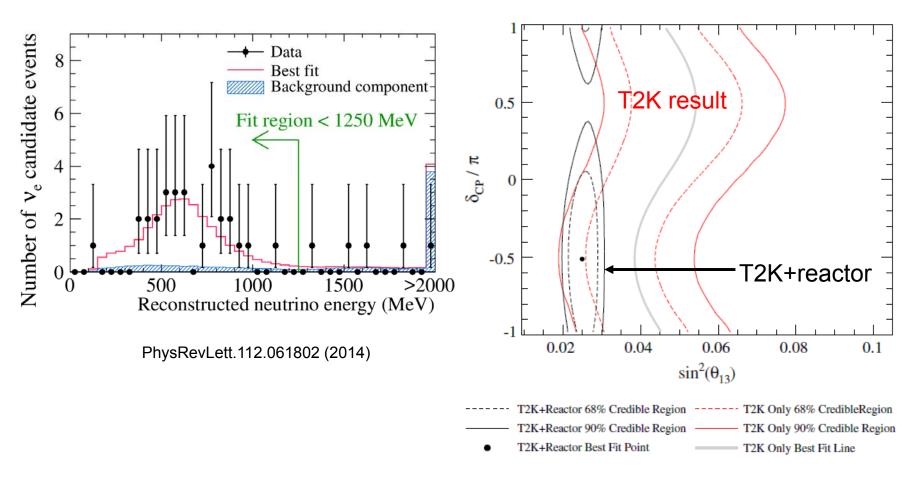




v_e appearance by T2K

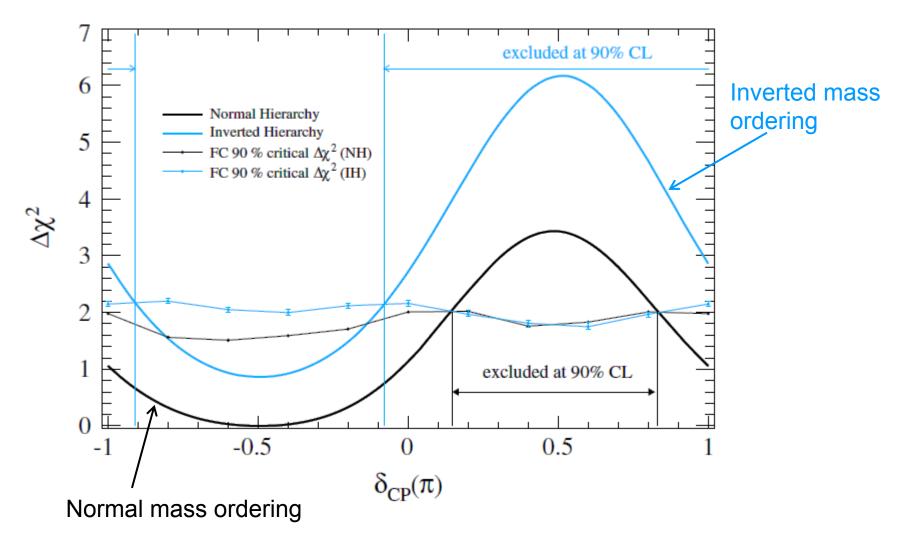
released in August 2013 w/ data till May 2013

Phys. Rev D.91, 072010(2015)



28 events observed over 4.92 ± 0.55 bkgs $\rightarrow 7.3\sigma$ excess First Confirmation of 'Appearance phenomenon' w/ > 5σ significance.

First constraint on δ_{CP} **by T2K**

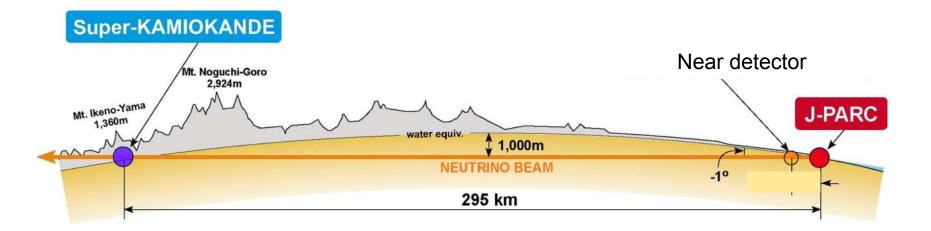


Measurements with antineutrino

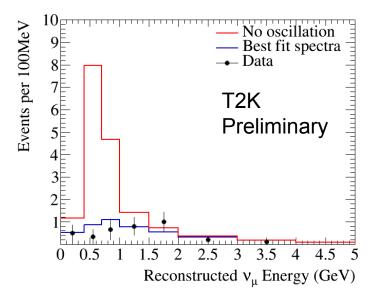
An important step towards CPV measurement Unique test of new physics

$\bar{\nu}_{\mu}$ disappearance

- > CPT theorem $\Rightarrow P(\nu_{\mu} \rightarrow \nu_{\mu}) = P(\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{\mu})$
- > $v_{\mu}/\overline{v}_{\mu}$ disappearance is insensitive to matter effect.
- > If we observe $P(\nu_{\mu} \rightarrow \nu_{\mu}) \neq P(\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{\mu})$, it may be due to CPT violation or non-standard interaction with matter.

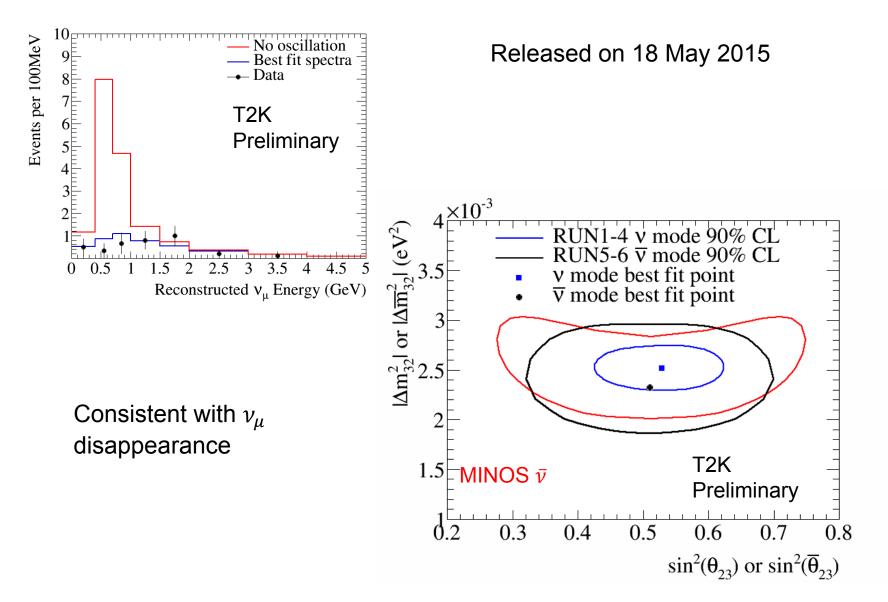


T2K w/ $\bar{\nu}$ beam data till Mar 12 2015 (2.3x10²⁰POT)



Released on 18 May 2015

T2K w/ $\bar{\nu}$ beam data till Mar 12 2015 (2.3x10²⁰POT)



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(Near future) Prospects

T2K and NOvA is accumulating data. T2K : 13% so far NOvA : 5%?

Do they have a sensitivity to CPV and mass ordering with full statistics?

Expected # of events w/ typical assumptions (*)

T2K : 106 v_e , 24 \bar{v}_e

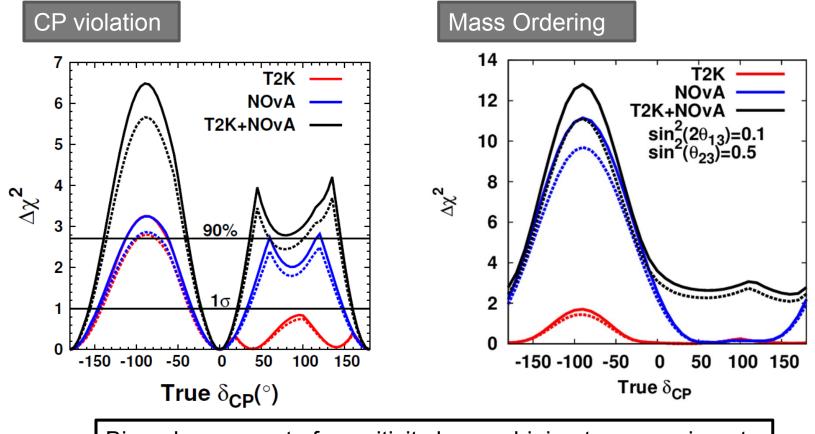
NOvA : 68 v_e , 32 \bar{v}_e

, while CPV effect is 27% at maximum and NOvA matter effect is 30%

(*) Not the latest number

T2K and NOvA combined sensitivity at full statistics

Shown is Normal mass ordering case.



Big enhancement of sensitivity by combining two experiments having different baseline length

solid: stat. only

dashed: assuming 5% normalization uncertainty on signal and 10% normalization uncertainty on background Prog. Theor. Exp. Phys. (2015) 043C01 by T2K collaboration

Summary

- In these 20 years, the skeleton of lepton flavor-mass mixing become significantly clear!
- > Missings are δ_{CP} , maximal or non-maximal θ_{23} and mas ordering
 - > In a few years, we may see answers with T2K and NOvA
 - Future projects (Hyper-Kamiokande and DUNE) will thoroughly explore these targets.
- Neutrino has always produced 'surprising'. New physics may appear
 - oscillation not explained by PMNS framework?
 - Sterile neutrino?
 - Non-standard interaction with matter
 - Lorentz violation at the Plank scale?
 - Precise measurements by both disappearance and appearance and using both CC and NC interactions are desired.
- > $\overline{\nu}_{\mu}$ disappearance by T2K just released.
- > In this summer, new results from T2K and NOvA are expected!